

Jet Propulsion Laboratory
California Institute of Technology

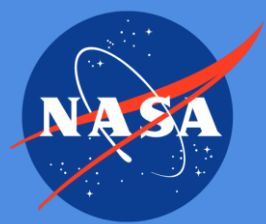
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Antenna Systems for Space Applications at NASA Jet Propulsion Laboratory

Dr. Paolo Focardi
Group Lead for Technology Development



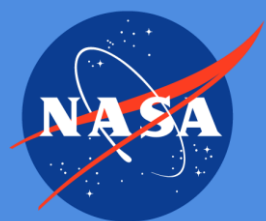


NASA Centers

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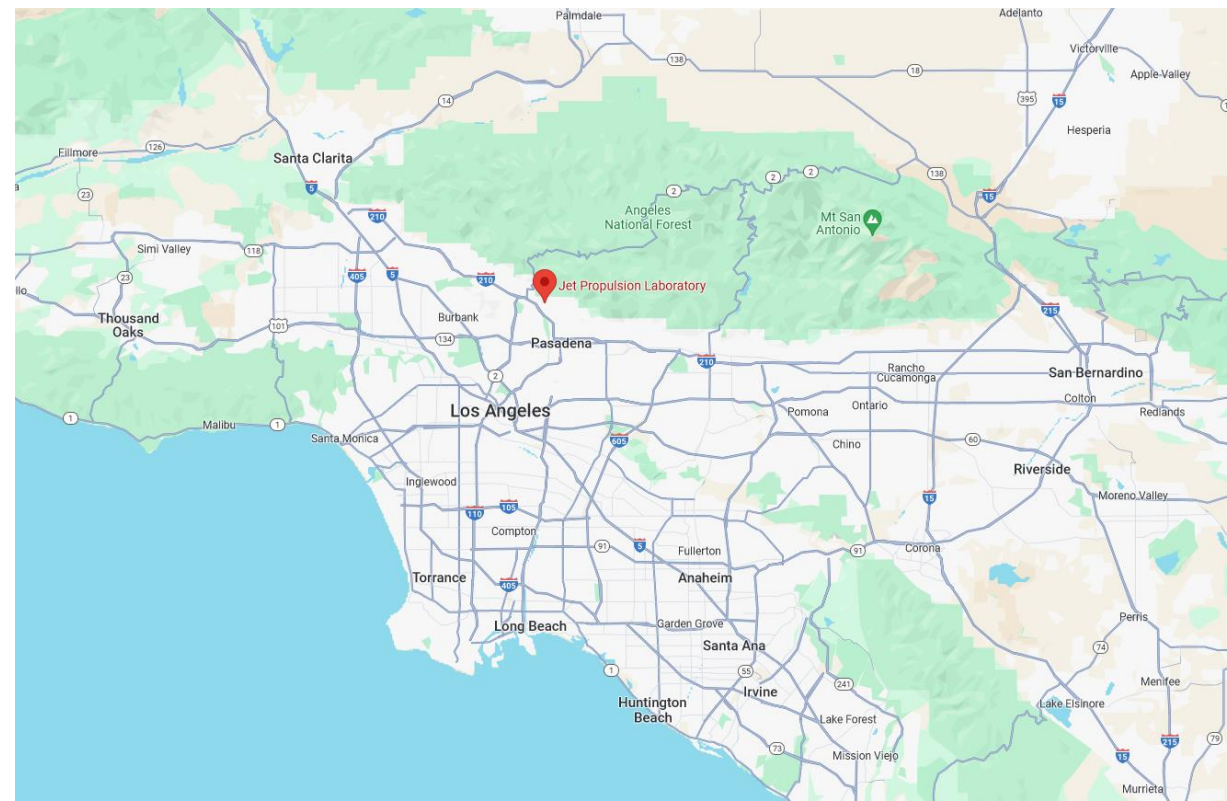
Jet Propulsion Laboratory (JPL)

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- JPL is about 10 miles north of downtown Los Angeles.
- Part of Southern California's historic aerospace industry.



- **JPL is a federally funded research and development center** funded by NASA and managed by Caltech.
- **> 5,500 employees**
- **> \$2B/year budget**, at the mercy of Congressional funding decisions



OUR MISSIONS

VOYAGER 1&2

EUCLID

ARIEL
(CASE)

JWST
(MIRI)

ROMAN
CORONAGRAPH

NuSTAR

SPHEREX

EUROPA
CLIPPER

PSYCHE

JUNO

NEO SURVEYOR

NEOWISE

VERITAS

EnVISION

MAIA

JASON-3

SBG

GRACE-FO

SUNRISE

CADRE

FSS
LUNAR
TRAILBLAZER

ASTER

AVIRIS, PRISM
UAVSAR, HyTES,
PALS, AirMSPI,
DopplerScat,
ScOpSAR

ASTHROS

NISAR

EZIE

MLS

MARS ODYSSEY

MRO

MISR

CLOUDSAT

PREFIRE

AIRS

SMAP

COSMIC

INCUS

OCO-2

MARS SAMPLE
RETURN

INGENUITY

SENTINEL-6

SWOT

CARBON PLUME
MAPPER

ECOSTRESS
EMIT
OCO-3
TEMPEST-H8
COWVR
COLD ATOM LABORATORY

MARS EXPRESS

PERSEVERANCE

MASS CHANGE

CURIOSITY

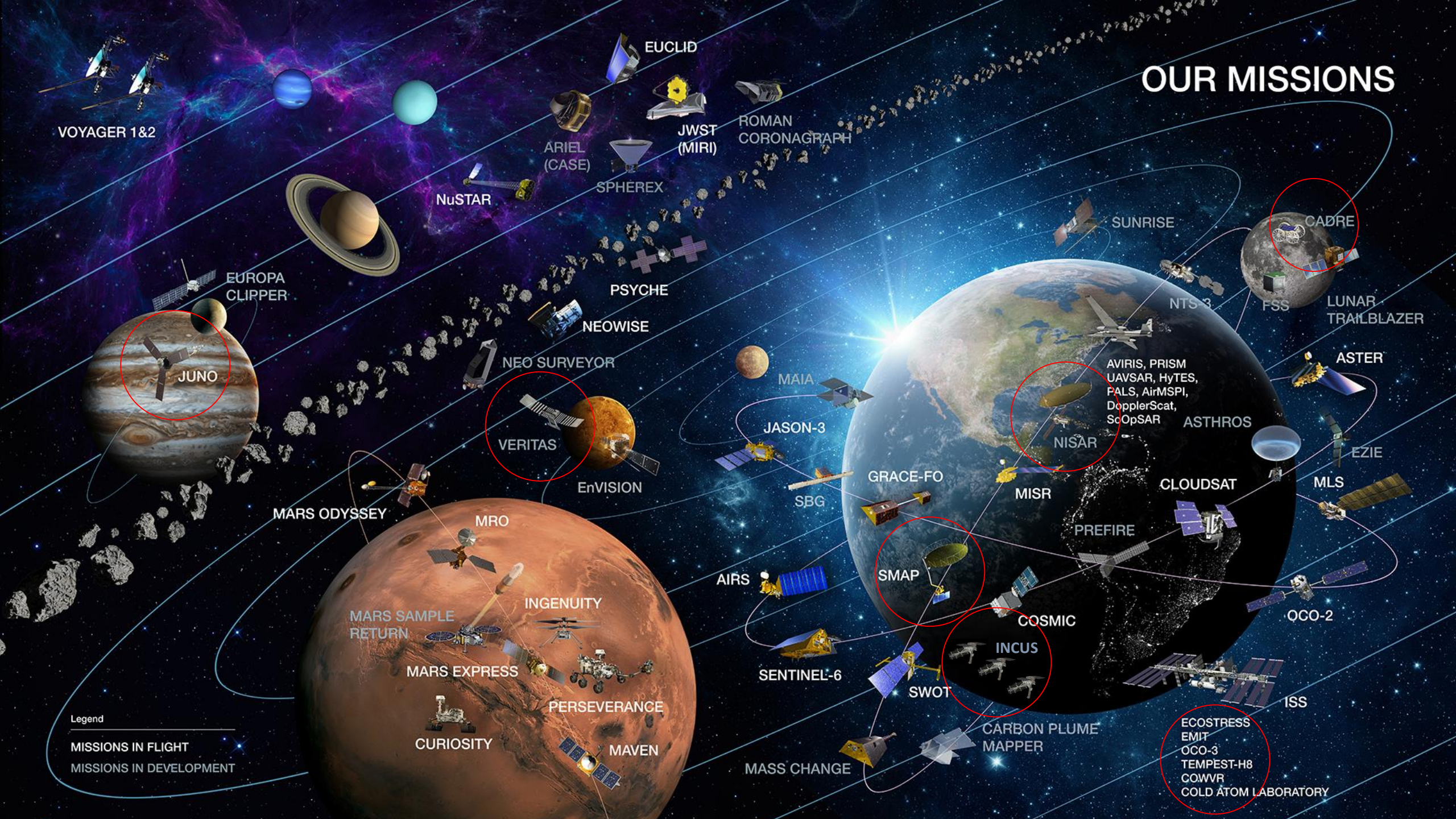
MAVEN

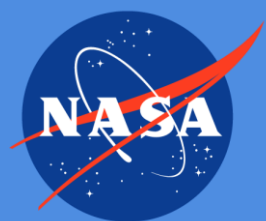
ISS

Legend

MISSIONS IN FLIGHT

MISSIONS IN DEVELOPMENT

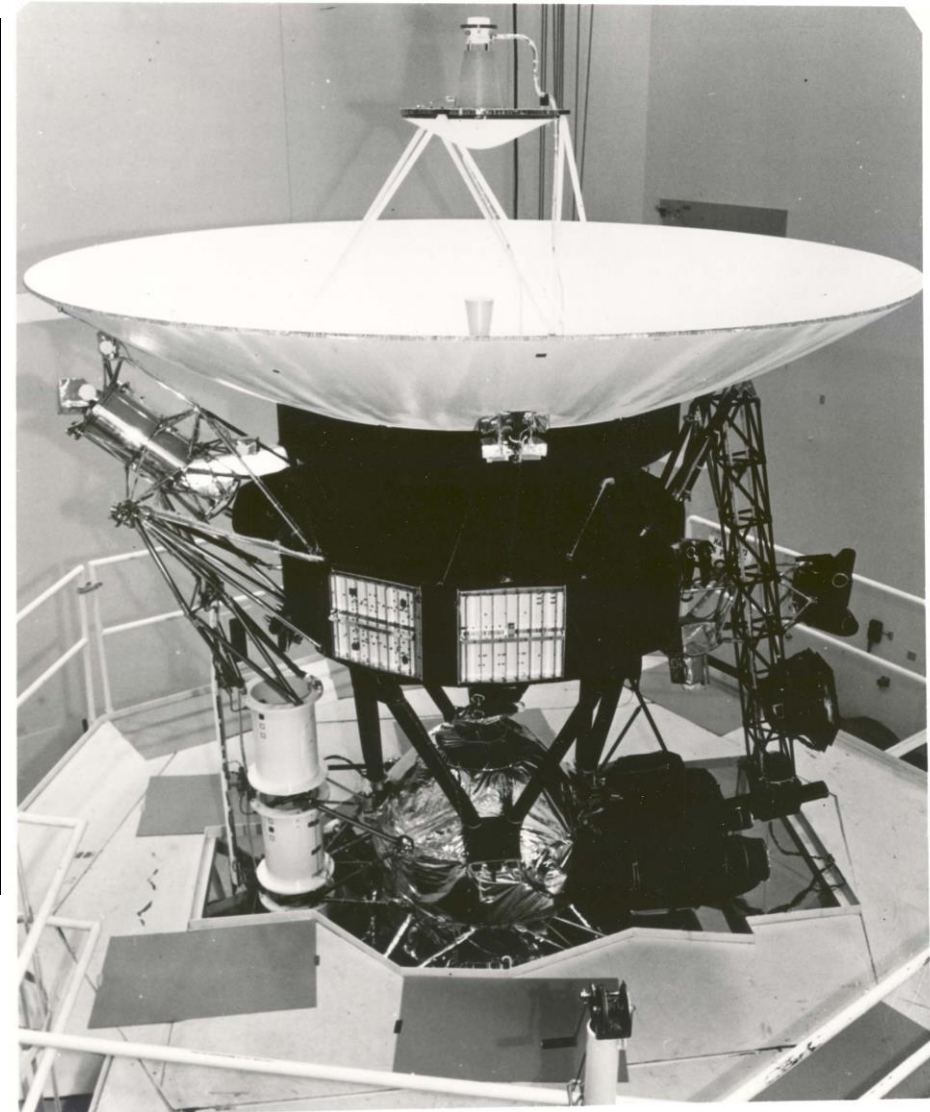




Voyager, Interstellar Space, 1977

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Voyager High-Gain Antenna

- 3.66-m diameter parabolic reflector
- F/D = 0.338
- S-band, TX & RX
- X-band, TX only

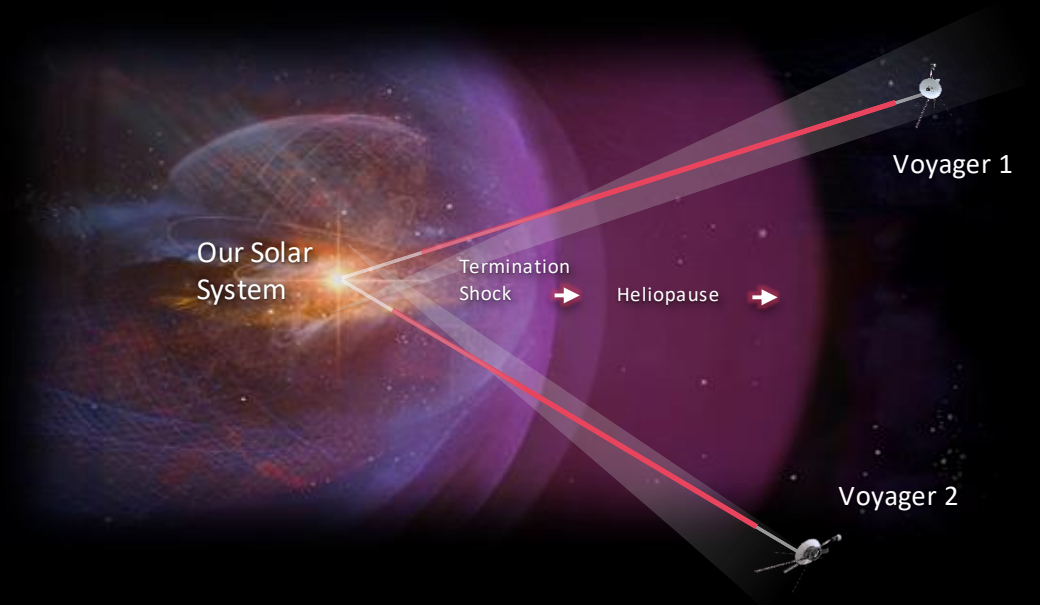
Voyager is powered by 3 radioisotope thermoelectric generators and uses **13-W radio frequency (RF) transmitters.**

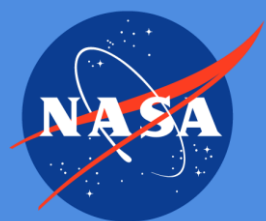
Very Far & Very Fast

Distance	Miles Away	AU
Voyager 1	15.2 billion	163
Voyager 2	12.7 billion	137

Light Times	One Way (HH:MM)	Two Way (HH:MM)
Voyager 1	22:30	45:00
Voyager 2	19:00	38:00

Speed	Miles/Hour	KM/Second	AU/Year
Voyager 1	38,025	17.0	3.6
Voyager 2	34,390	15.4	3.2
Solar System	492,126	220	46.3





Magellan, Venus, 1989

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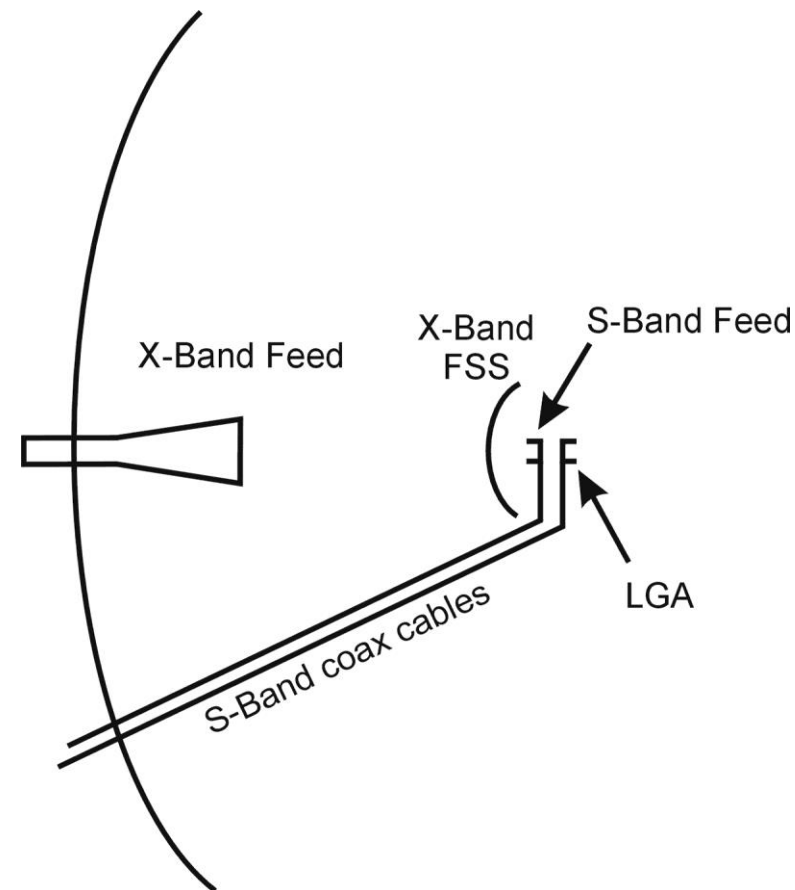
The Magellan spacecraft being deployed by the Space Shuttle Atlantis during STS-30, along with its inertial upper stage. To the left of the reflector antenna, the altimeter antenna instrument is also visible.

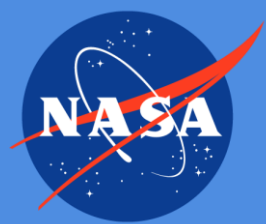
The Voyager high-gain antenna was integrated with an S-band instrument that operated in three modes:

1. synthetic aperture radar (SAR)
2. altimeter
3. radiometer

For the first time, the side-looking SAR provided a remarkably detailed map of Venus's surface, which is not visible due to the planet's dense atmosphere. The technology proven by Magellan was later used on the Shuttle Radar Topography Mission.

Antenna Schematic Diagram

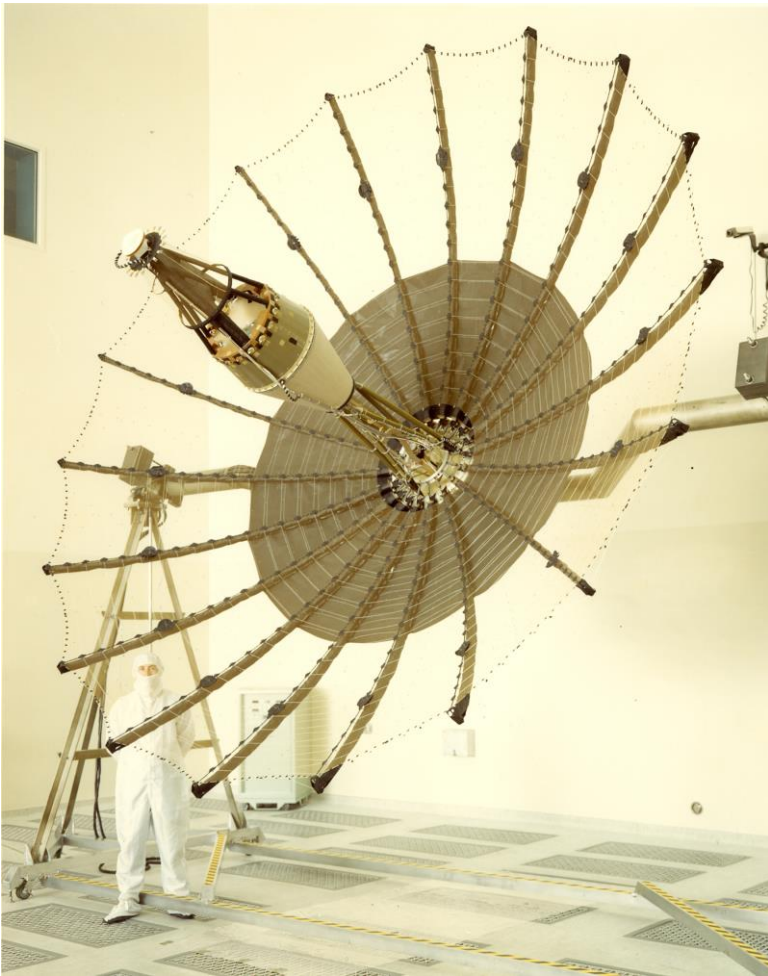




Galileo, Jupiter, 1989

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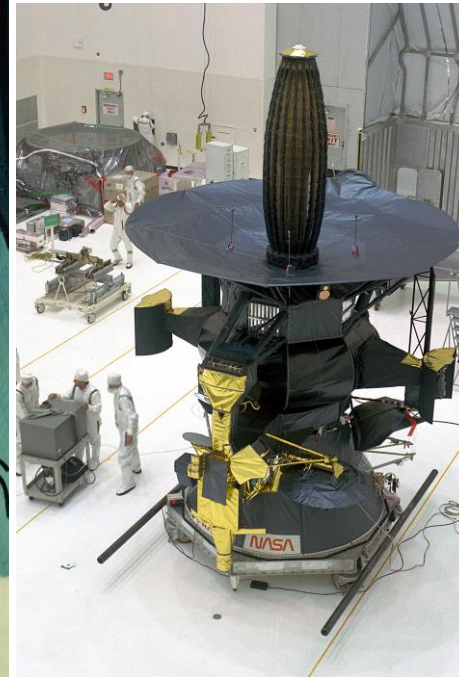


Galileo High-Gain Antenna

- 4.755-m diameter parabolic reflector
- Radial rib mesh deployable reflector, 10 opi mesh
- S-band & X-band, TX & RX

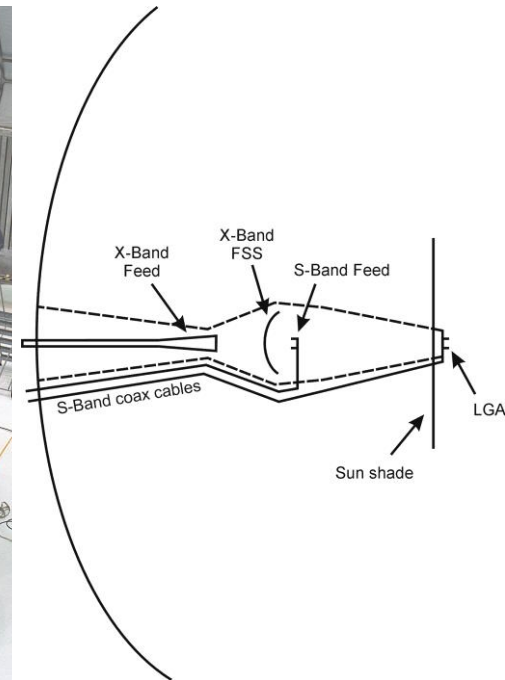


The Galileo antenna fully deployed (left) and partially deployed (above) while being tested on the plane polar nearfield antenna range at JPL's MESA Antenna Test Facility.



Galileo S/C with stowed antenna.

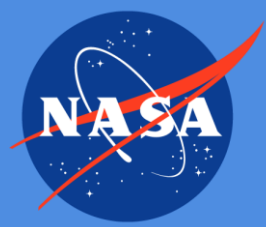
Antenna Schematic Diagram





Space Shuttle Atlantis launching Galileo into orbit on October 18, 1989.





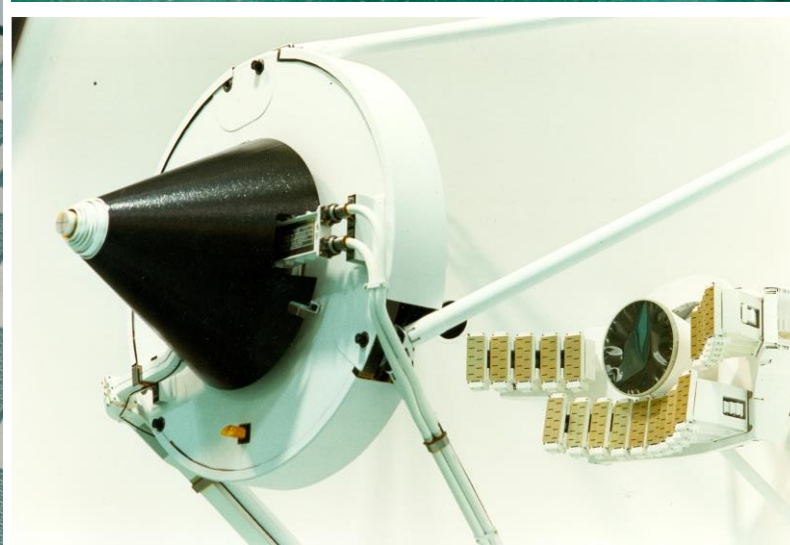
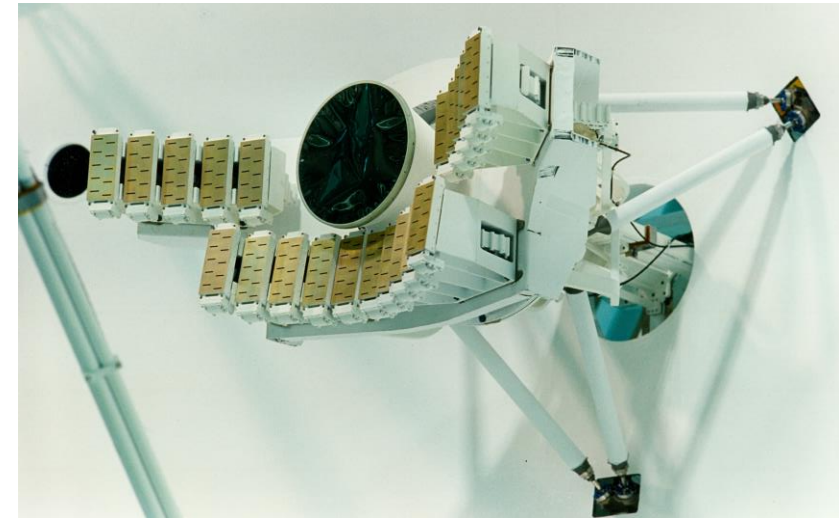
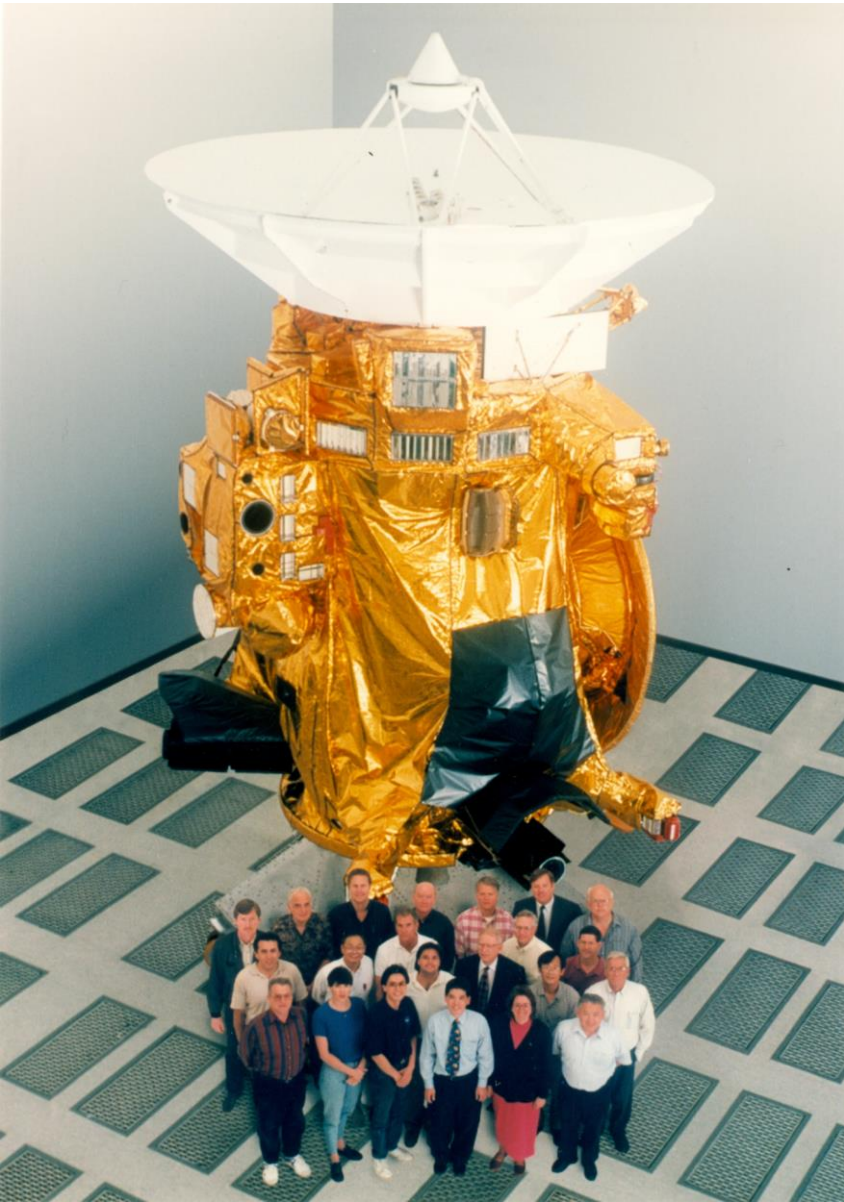
Cassini, Saturn, 1997

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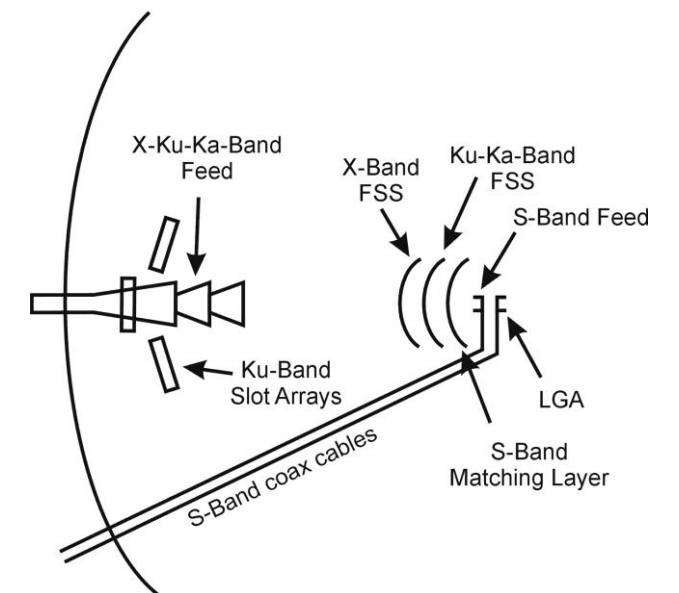
Cassini High-Gain Antenna

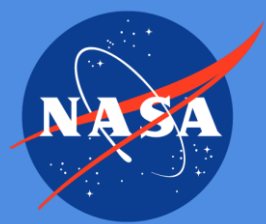
- 4-m diameter parabolic reflector
- $F/D = 0.33$
- S-X-Ku-Ka-band operations

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Antenna Schematic Diagram

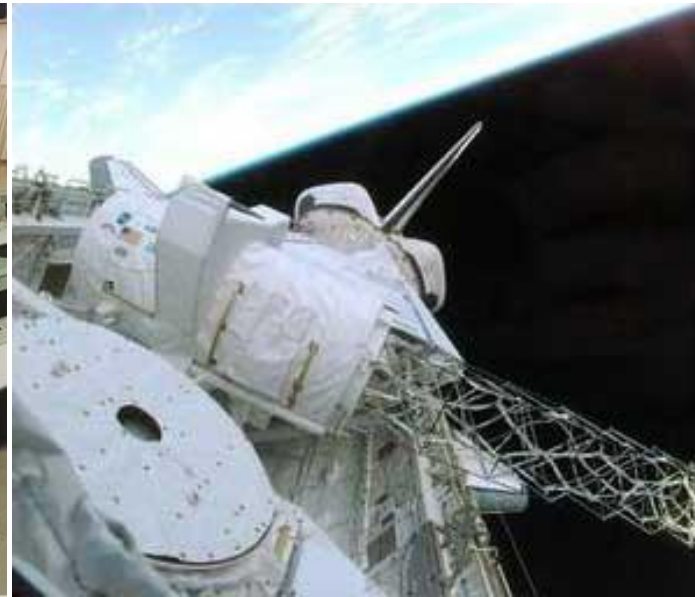
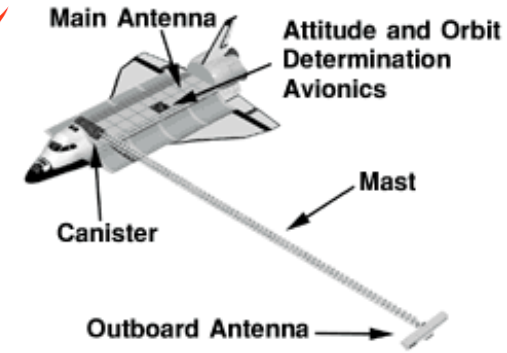
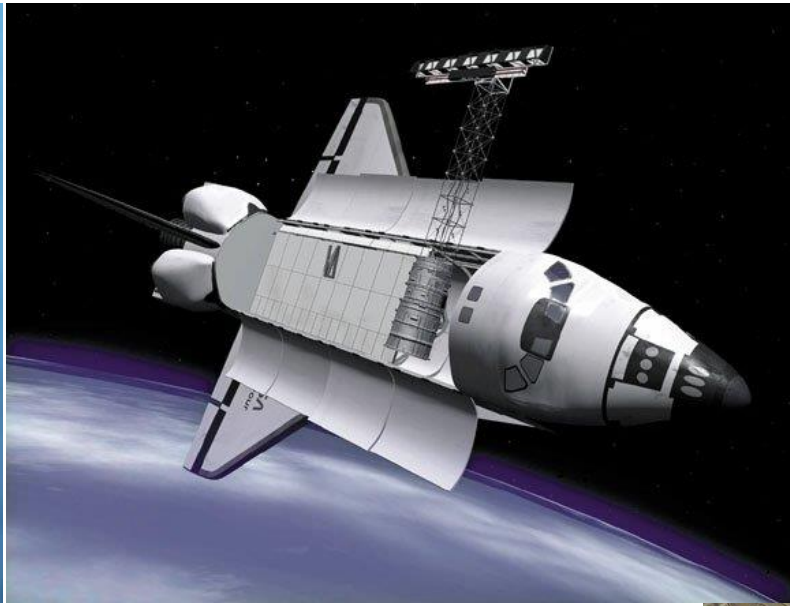


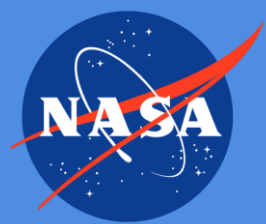


Shuttle Radar Topography Mission, 2000

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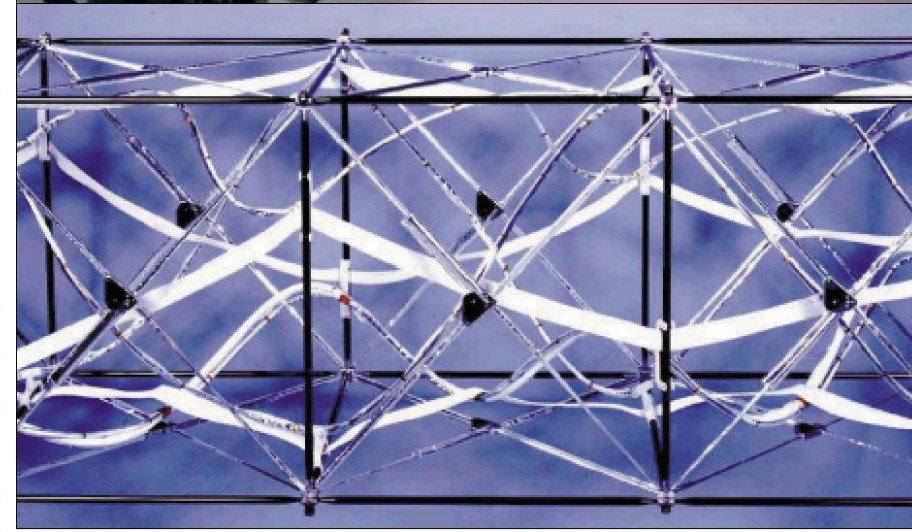
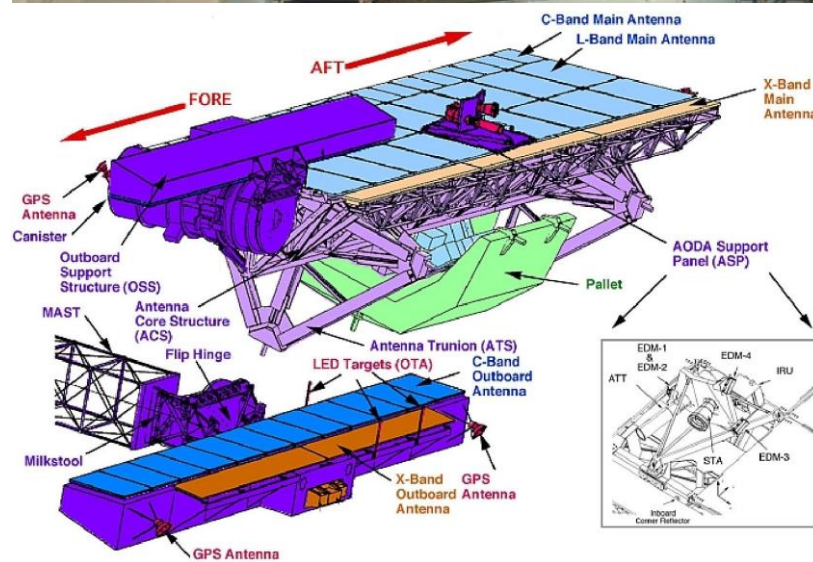




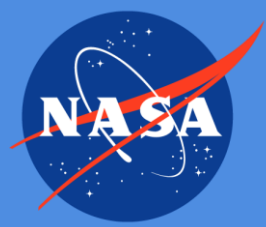
RX-Only Antennas and 60-m Boom

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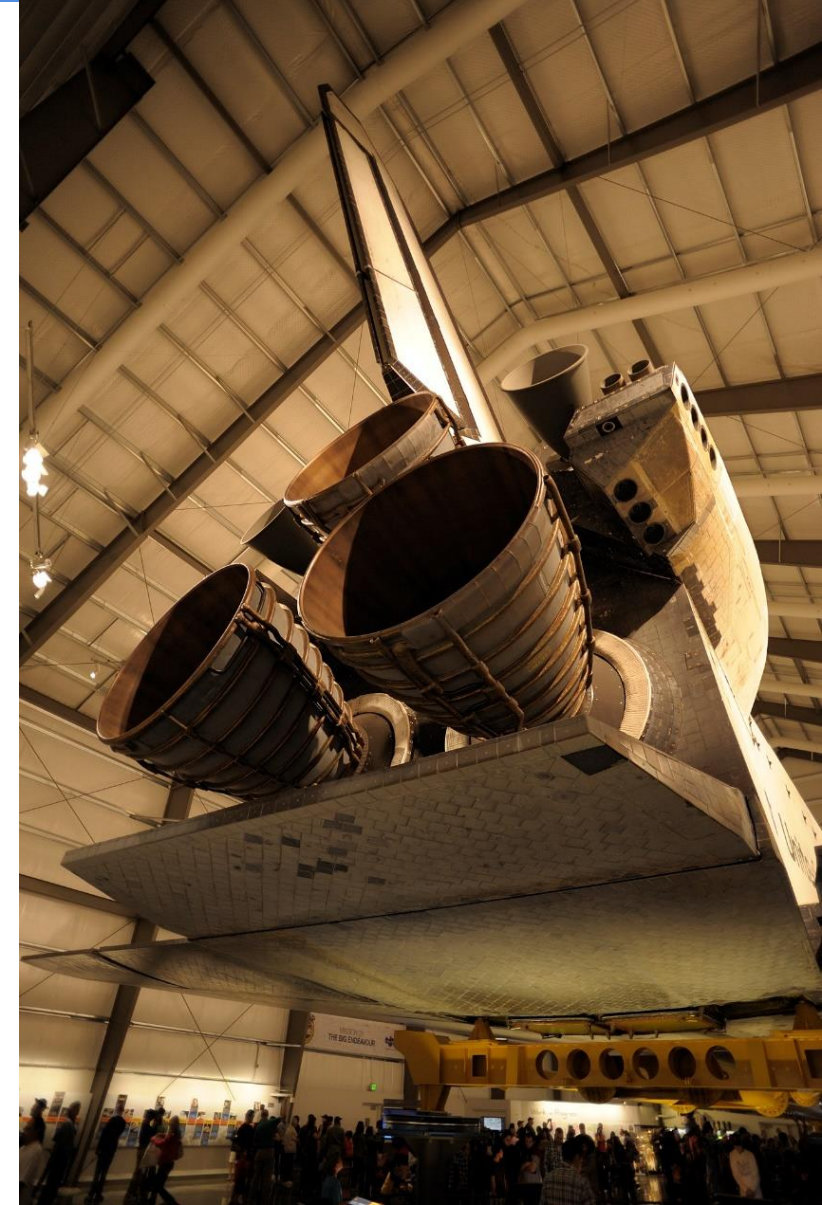
- TX/RX Antennas
 - C-band, 12 m x 0.75 m patch array
 - X-band, 12 m x 0.40 m slotted waveguide
- RX-Only Antennas
 - C-band, 8 m x 0.75 m patch array
 - X-band, 6 m x 0.40 m slotted waveguide



Space Shuttle Endeavour at the California Science Center

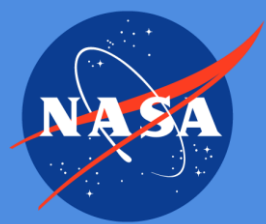
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Space Shuttle Endeavour successfully completed its last mission to space in 2011 and then arrived in LA in 2012 with a flyby over JPL.





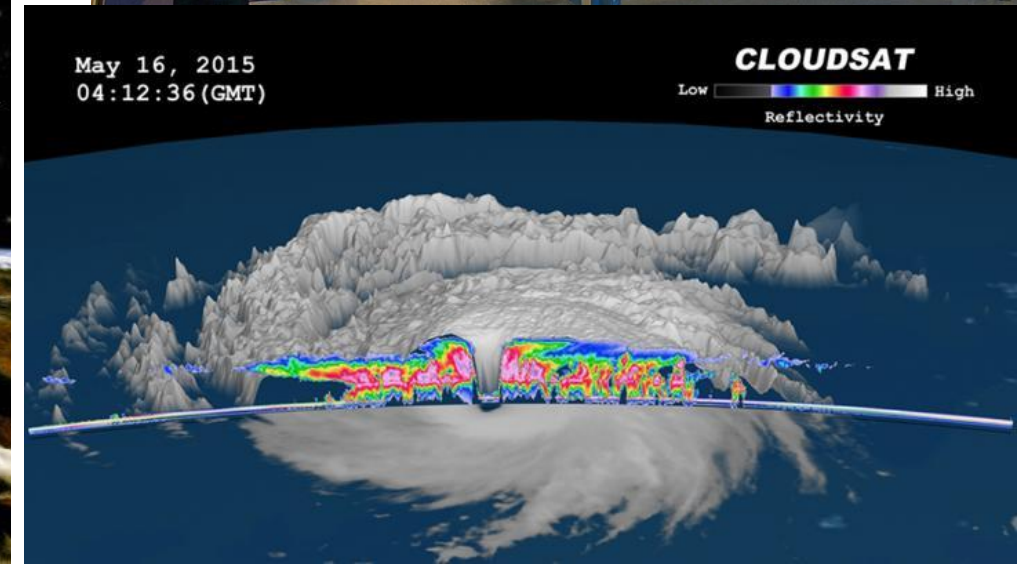
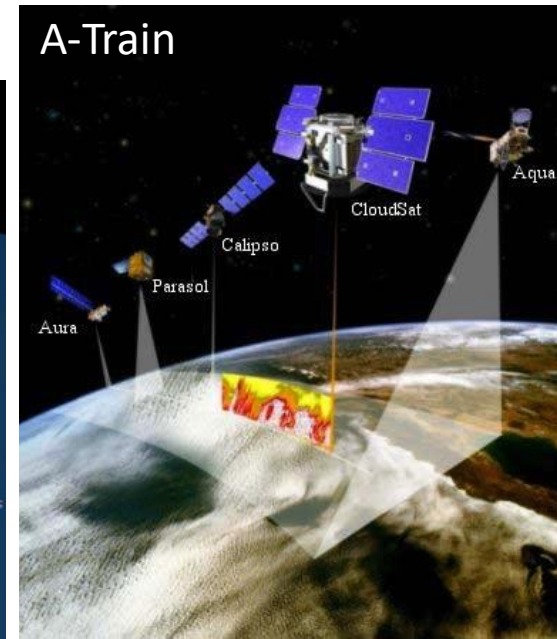
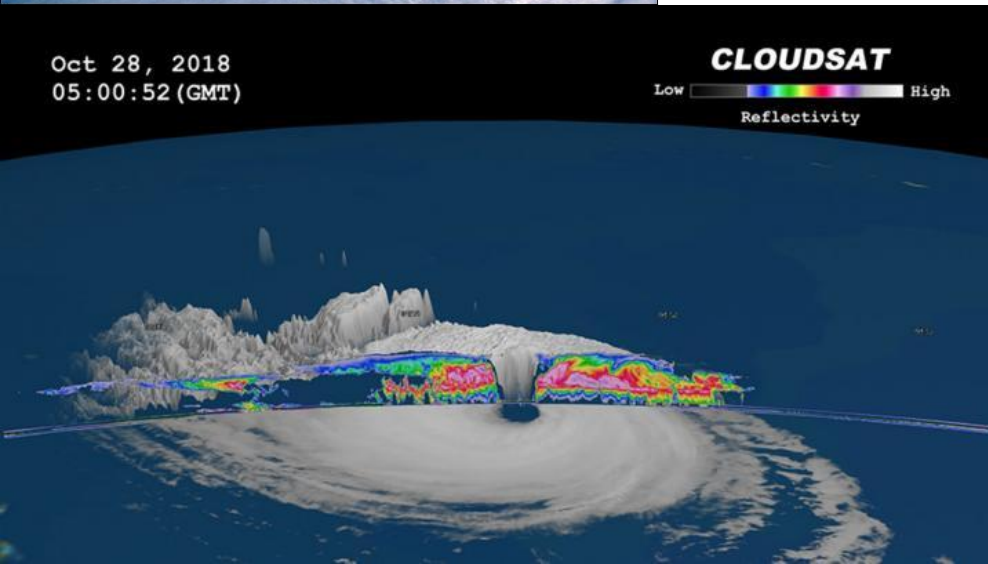
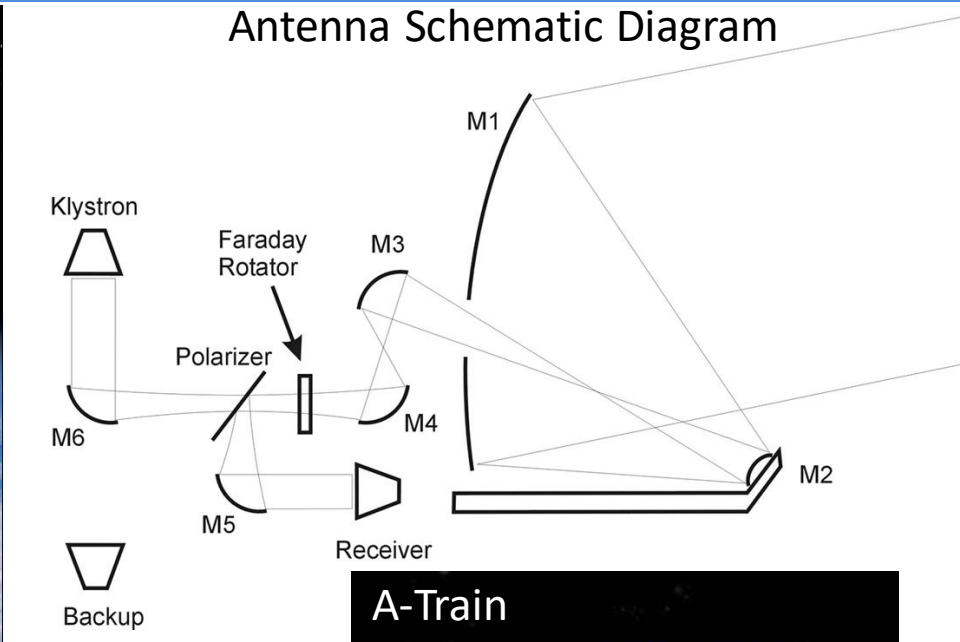
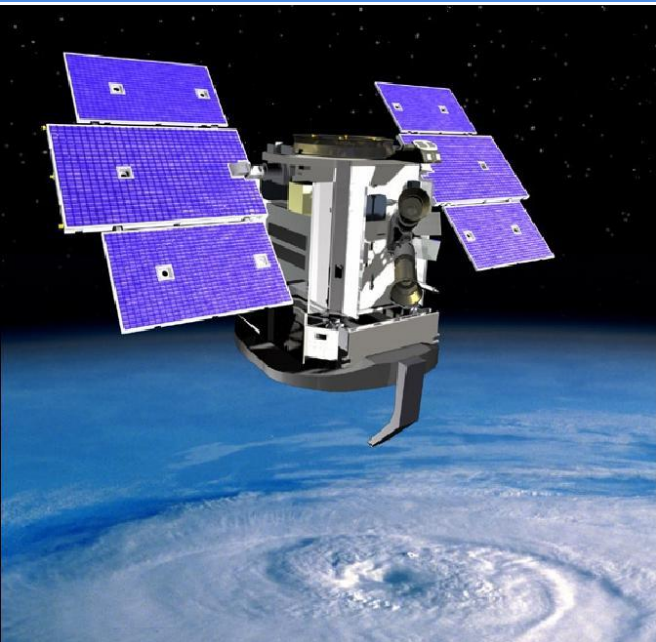
CloudSat, Cloud Profiling Radar, 2006

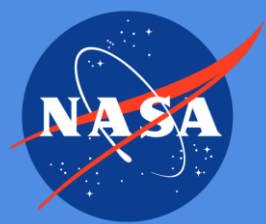
CloudSat Antenna

- 1.85-m offset Cassegrain reflector with QOTL
- W-band operations

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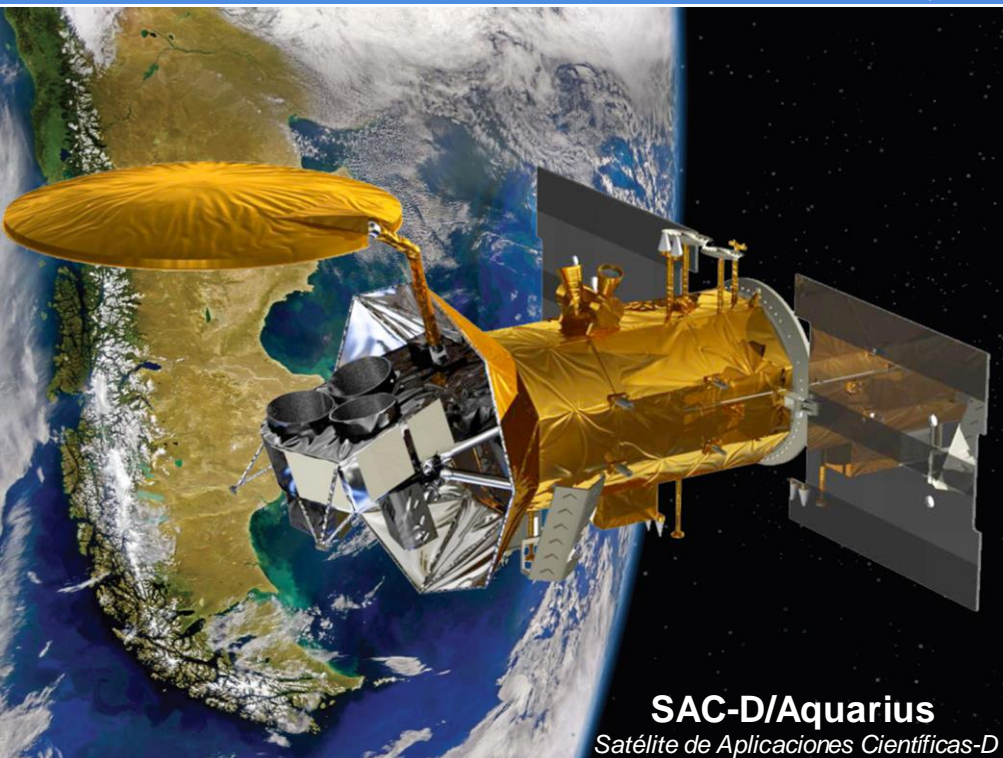




Aquarius, Ocean Water Salinity, 2011

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SAC-D/Aquarius
Satélite de Aplicaciones Científicas-D

- SAC-D/Aquarius was a collaboration between:



Goddard Space Flight Center
Jet Propulsion Laboratory
California Institute of Technology

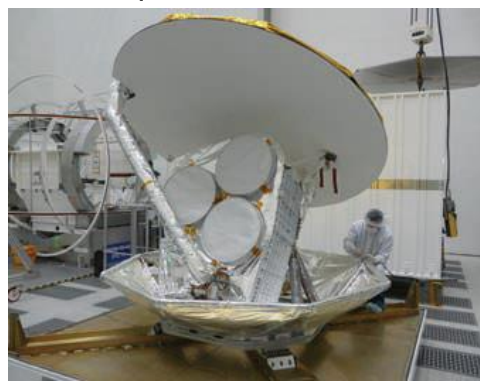
CONAE
Comisión Nacional de
Actividades Espaciales

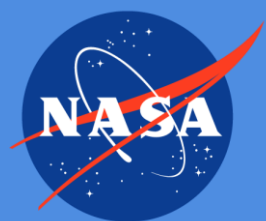
- Its mission was to measure ocean water salinity from space.
- The Aquarius instrument was delivered by Goddard and JPL to CONAE and was integrated with the SAC-D spacecraft.
- It was launched on June 10, 2011, from Vandenberg Air Force Base (AFB), CA.
- It consisted of:
 - 2.5 m solid (graphite composite) deployable offset reflector
 - Three dual-pol, corrugated L-band feed horns
 - Three L-band radiometers (one for each feed)
 - An L-band radar for sea-surface roughness error correction of radiometric measurements
- Mission was terminated on June 8, 2015, because of the failure of a component of onboard power regulation and attitude control.

Aquarius Horn



Aquarius Instrument



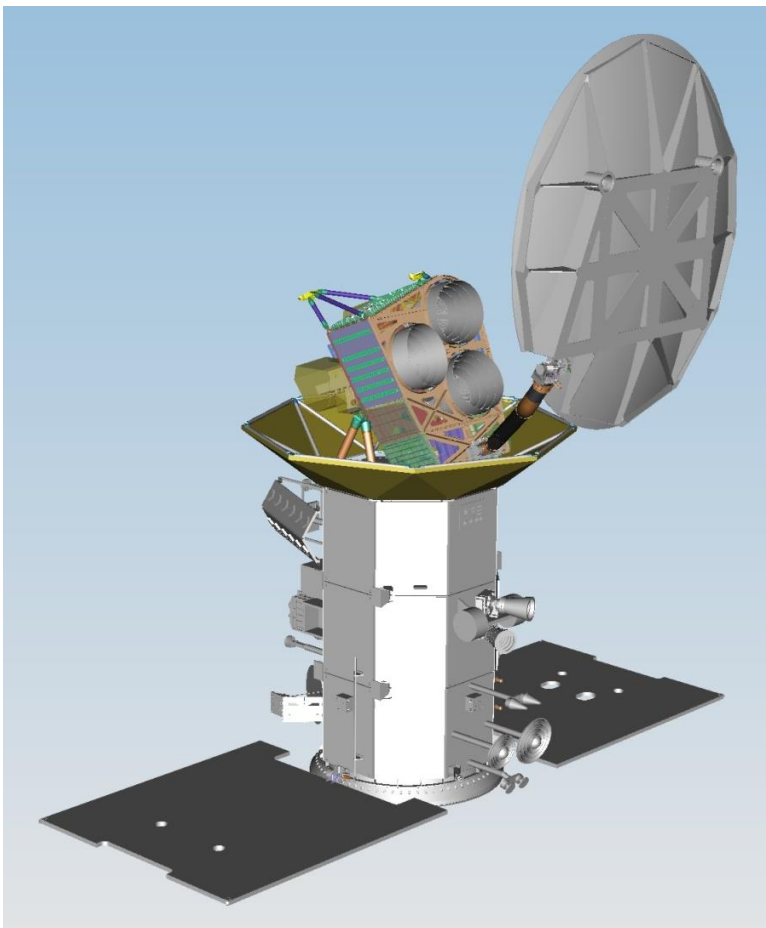


Aquarius RF Modeling

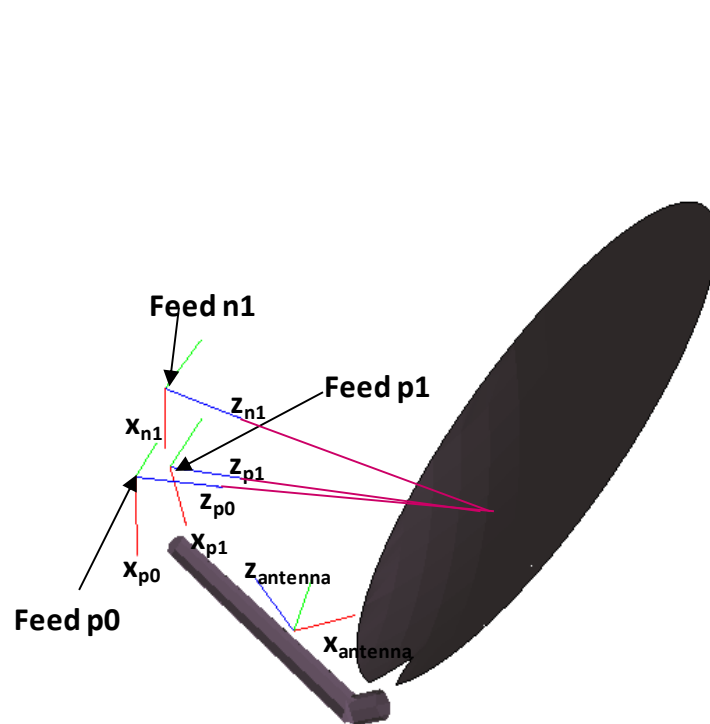
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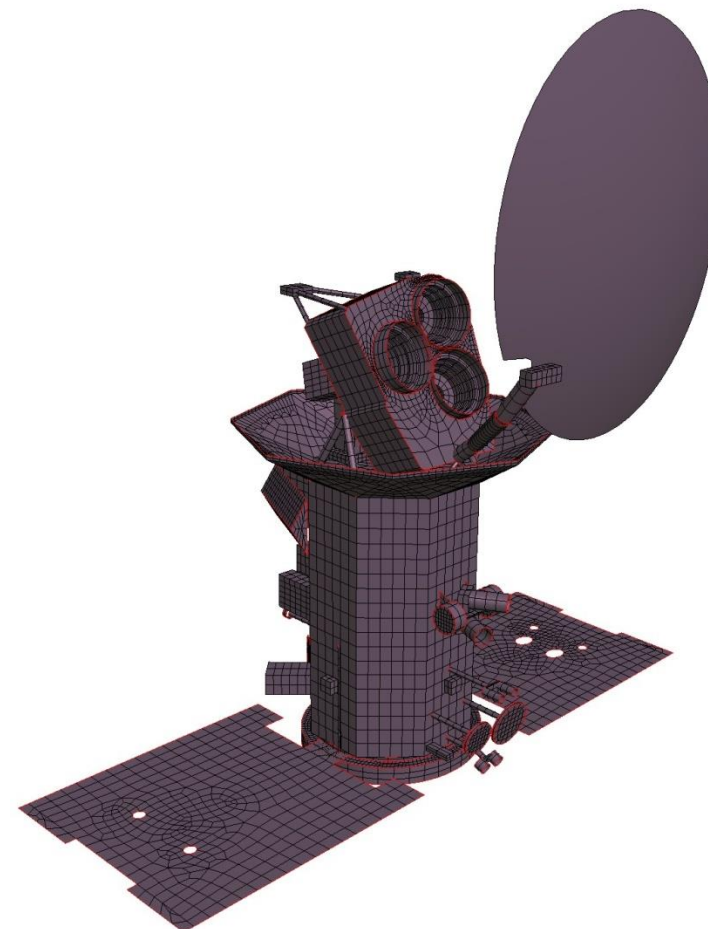
Mechanical Model

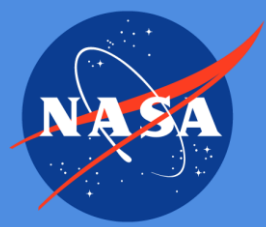


2002 RF Model



2011 RF Model

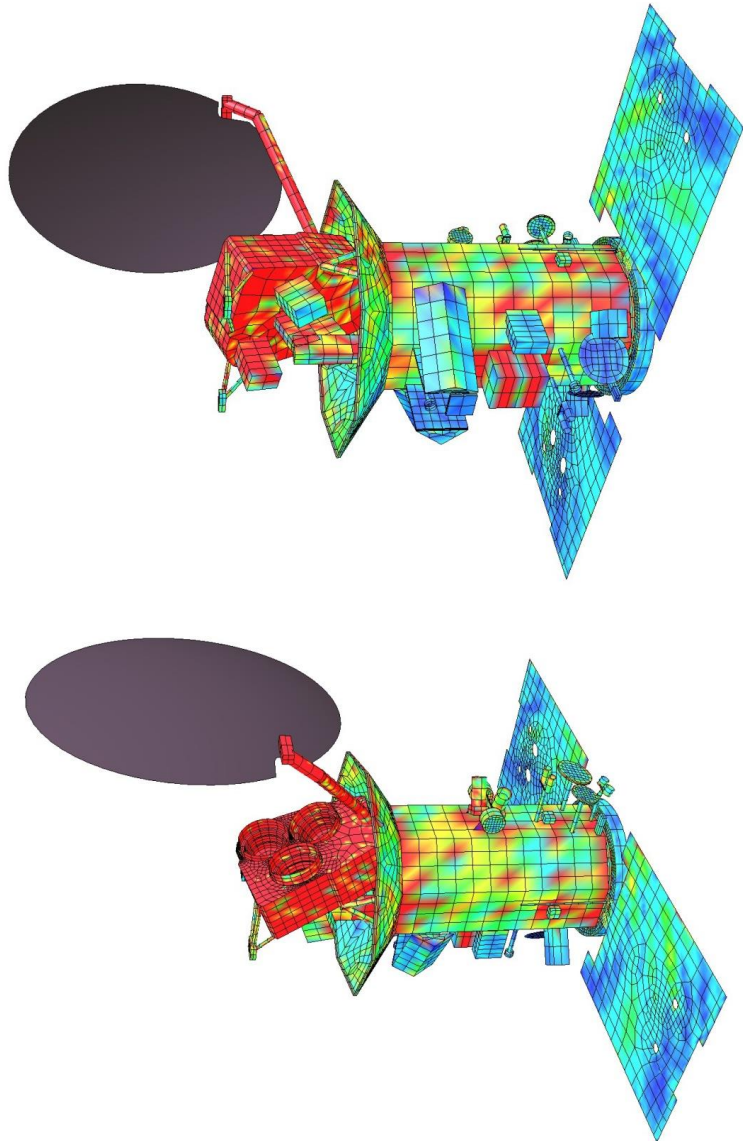


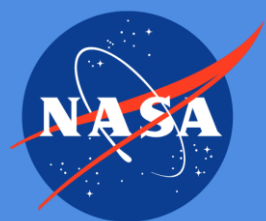


Aquarius Model RF Currents

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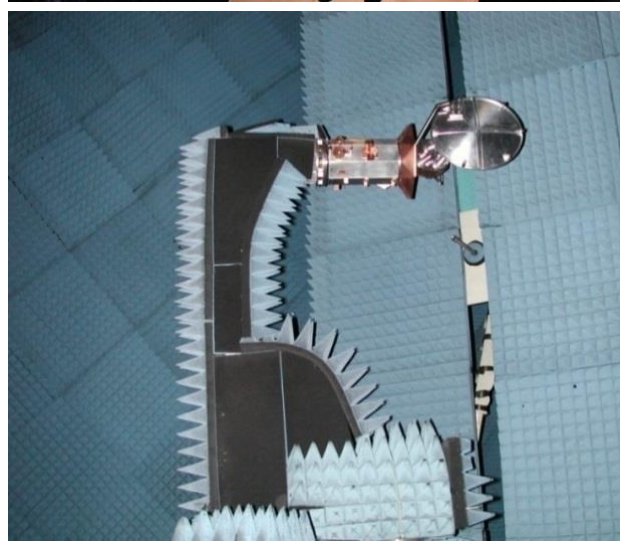
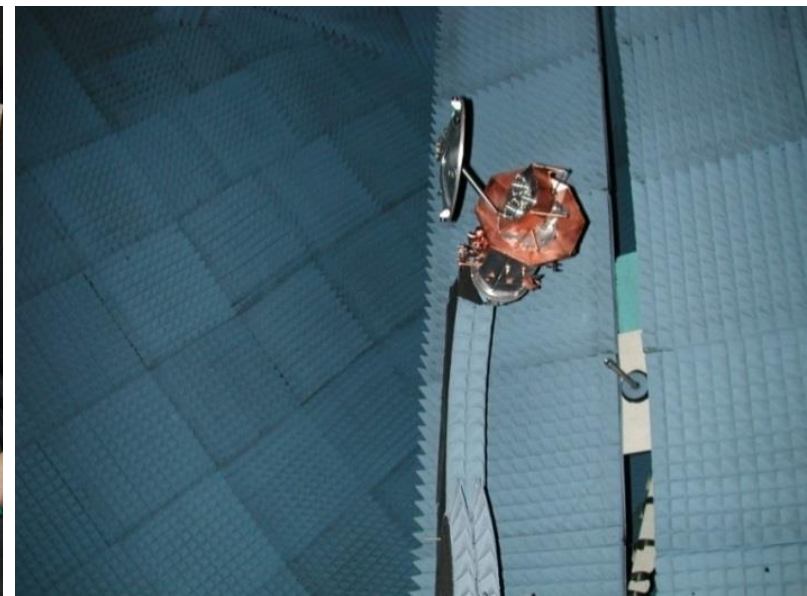
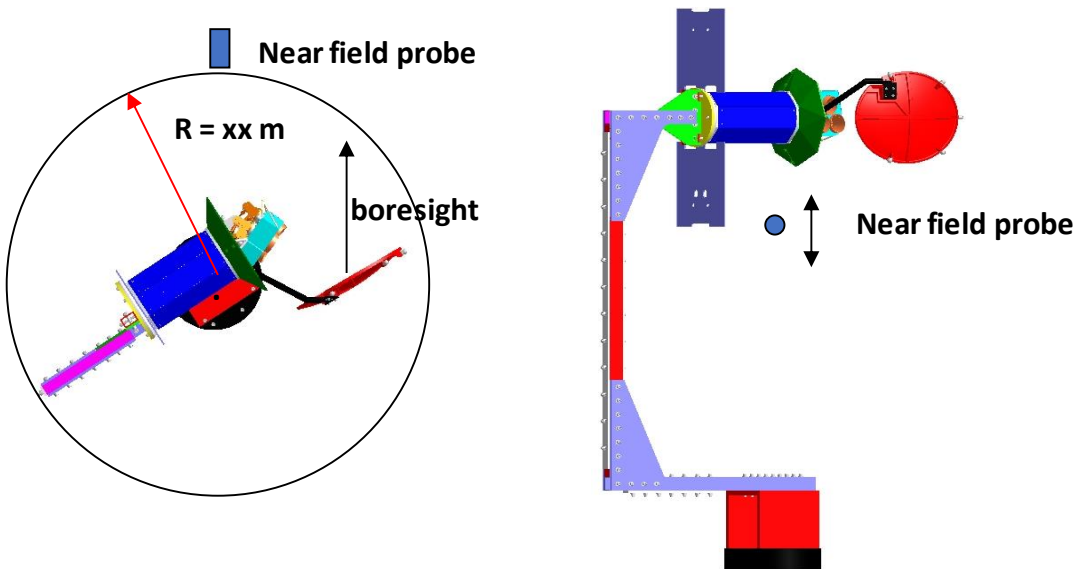




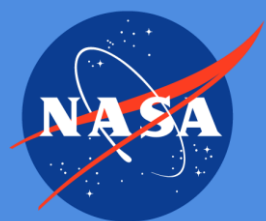
Aquarius Scale Model

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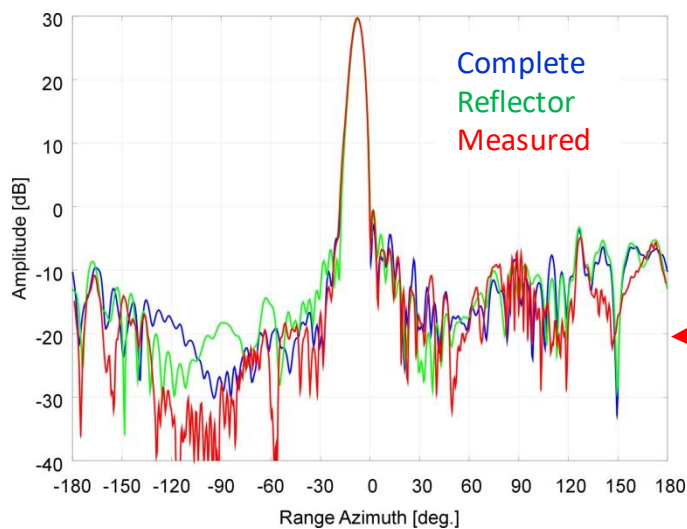
- 1/10th scale model (operating frequency scaled to Ku-band)
- Measured in cylindrical near-field antenna range at JPL



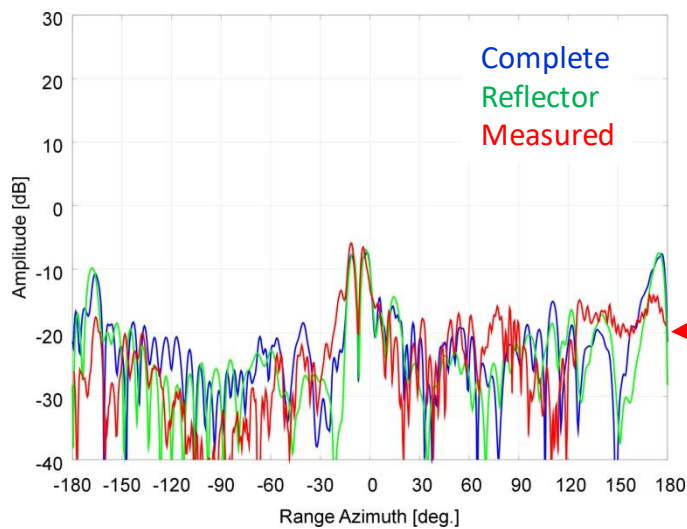
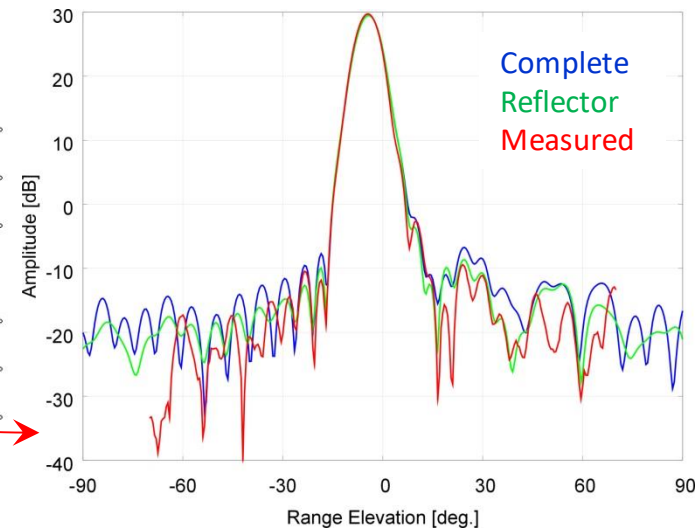
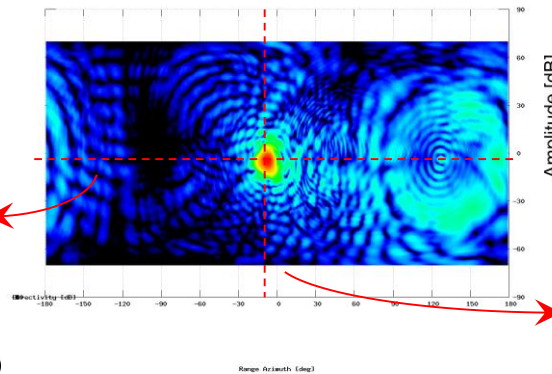
Calculated & Measured Radiation Patterns

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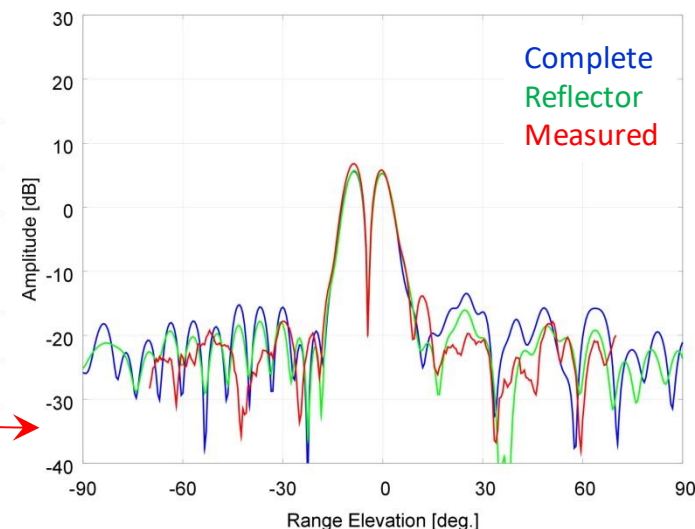
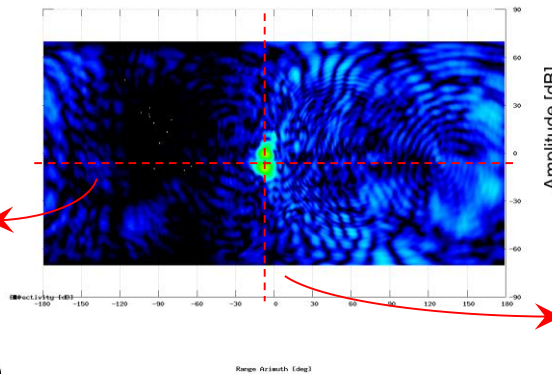
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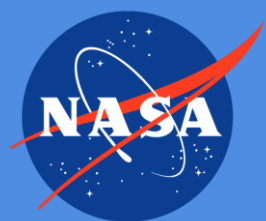


Measured Co-pol patterns



Measured Cx-pol patterns

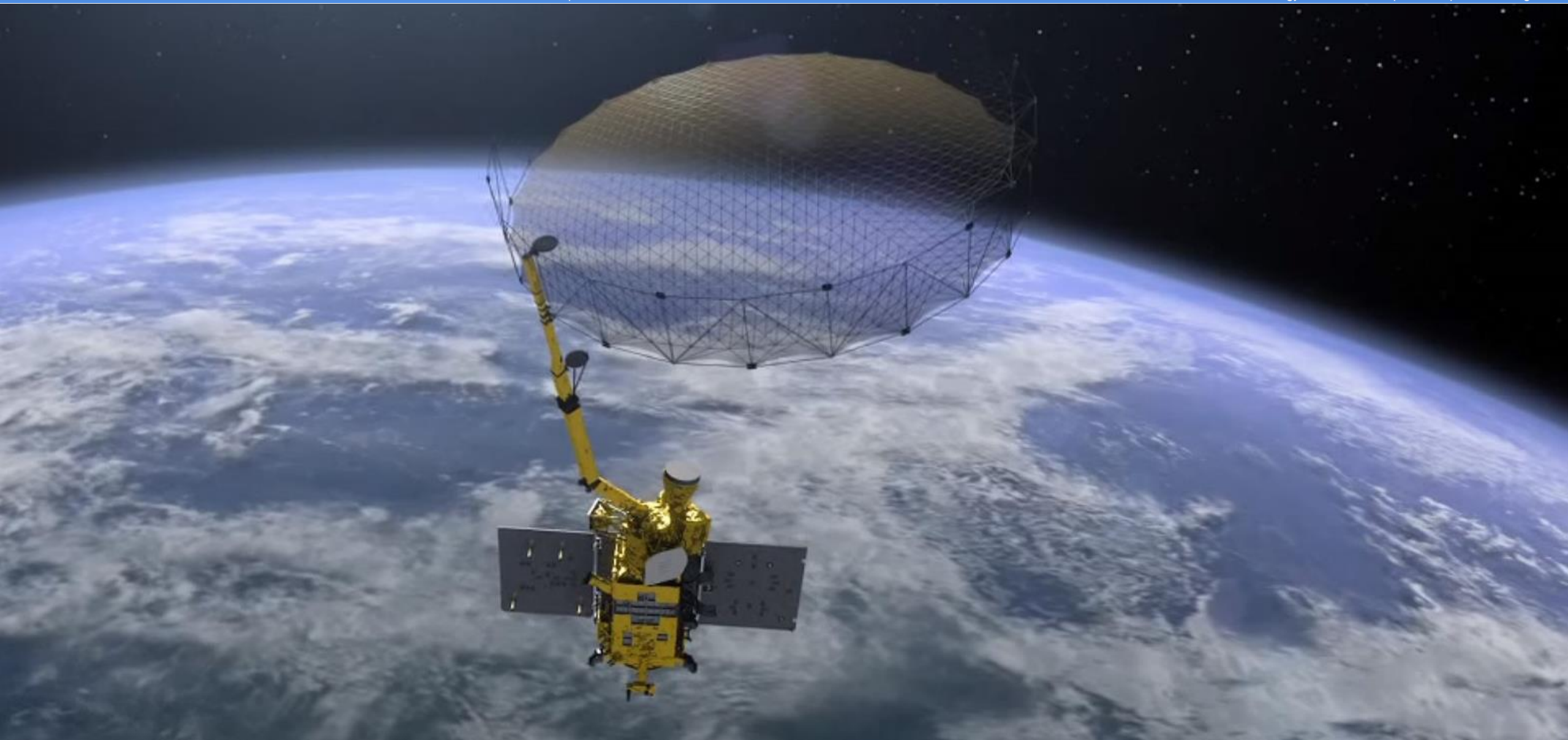


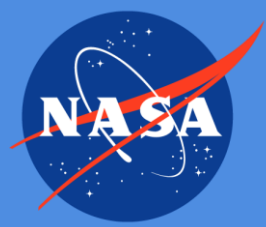


SMAP, Soil Moisture Active Passive, 2015

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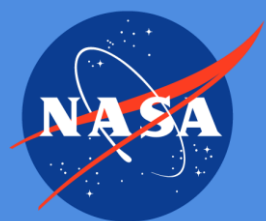
SMAP, Soil Moisture

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- SMAP's mission is to measure soil moisture and its freeze/thaw state.
- It was launched on January 31, 2015, from Vandenberg AFB, CA.
- The instrument consists of:
 - 6 m deployable mesh offset reflector
 - A single dual-band, dual-pol, corrugated feed horn
 - L-band SAR
 - L-band radiometer
- The antenna boresight beam is pointed 35.5° off of Nadir.
- The antenna instrument spins at ~ 14.6 RPM around Nadir.
- The radiometer data is more accurate than the SAR data but has a spatial resolution of about 40 km.
- The SAR spatial resolution is 1-3 km.
- In July 2015, the radar ceased its operations due to a sudden failure of the low-voltage power supply.
- The radiometer is still operational and has provided a large amount of calibrated soil moisture data on a global scale.

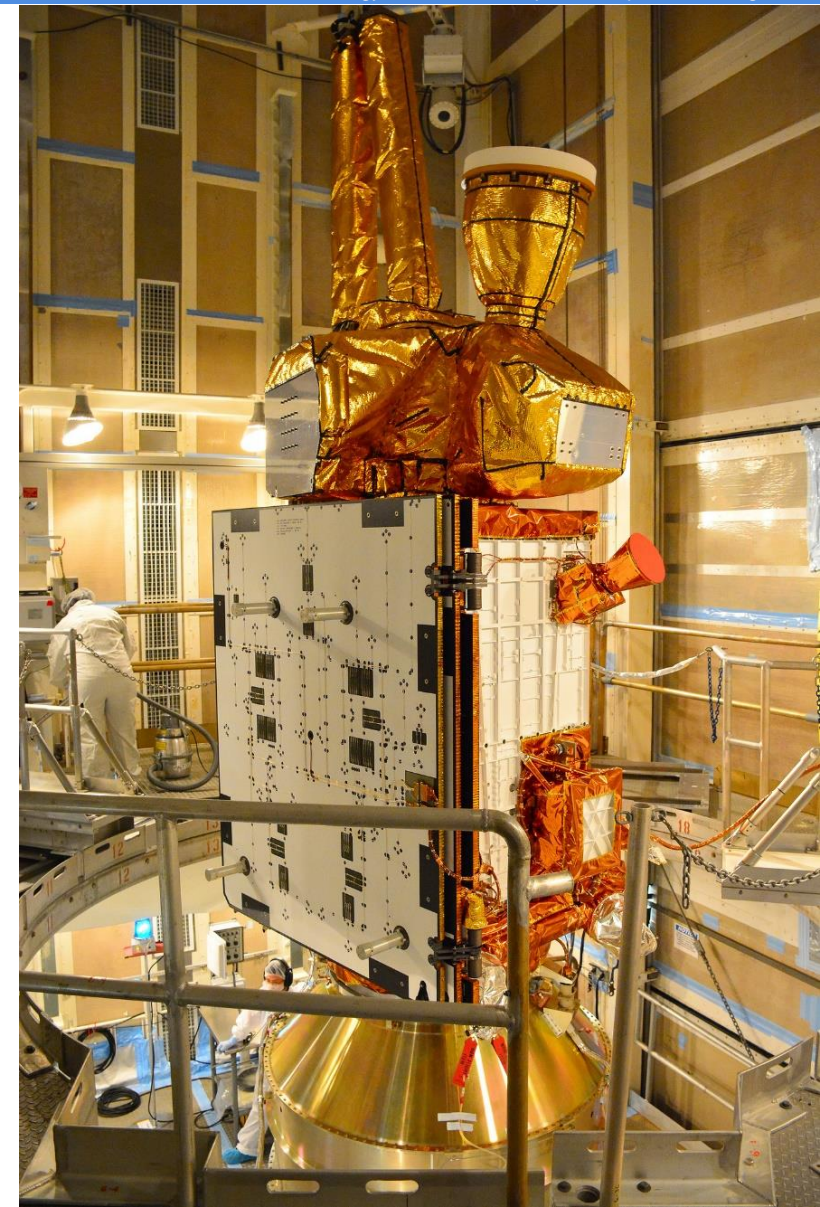


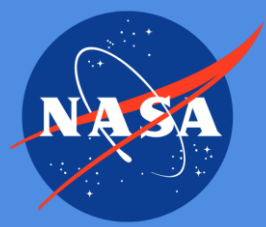


Delivery to Vandenberg AFB

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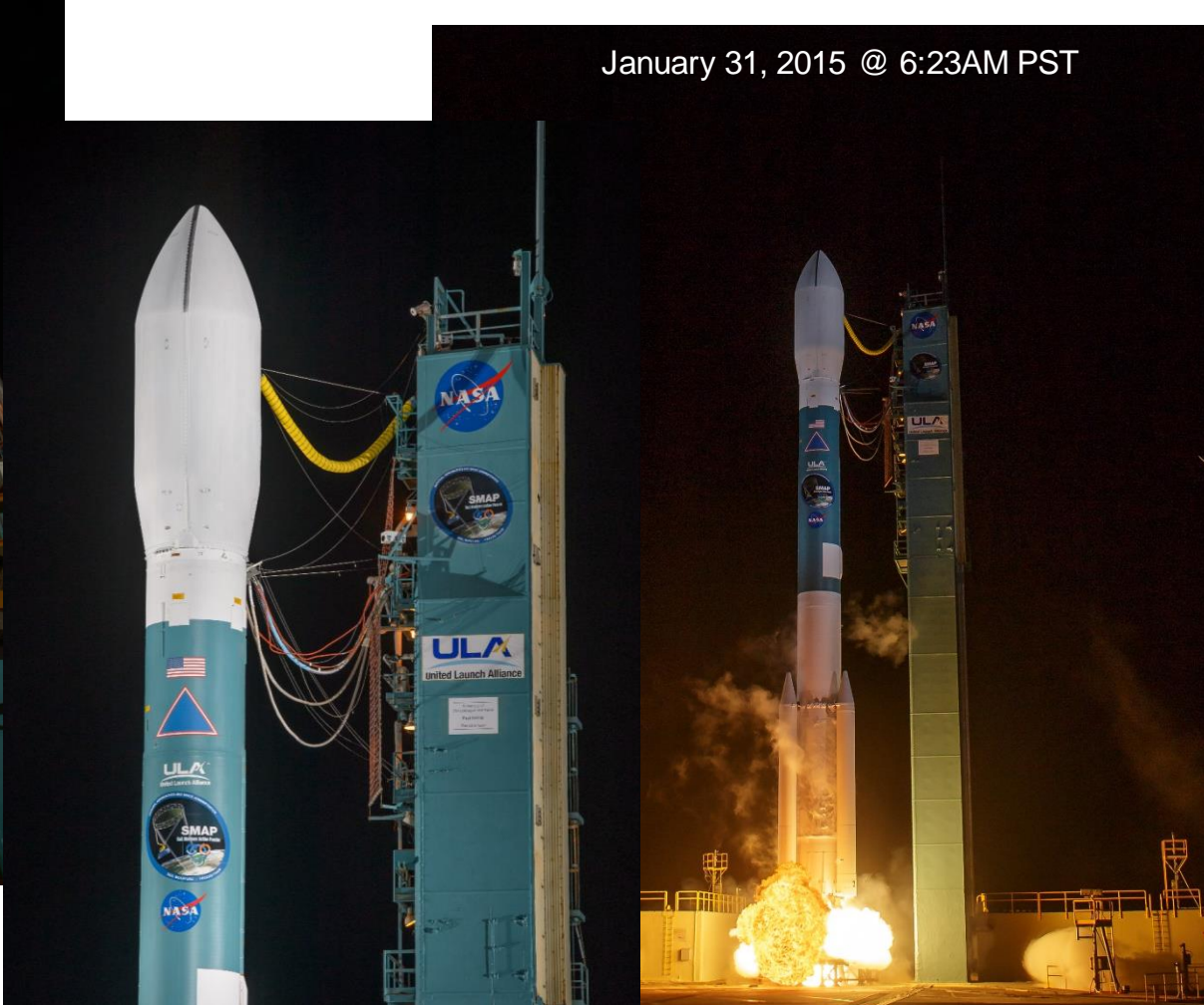


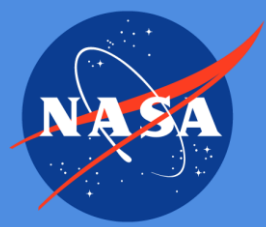


Delta II Launch Vehicle

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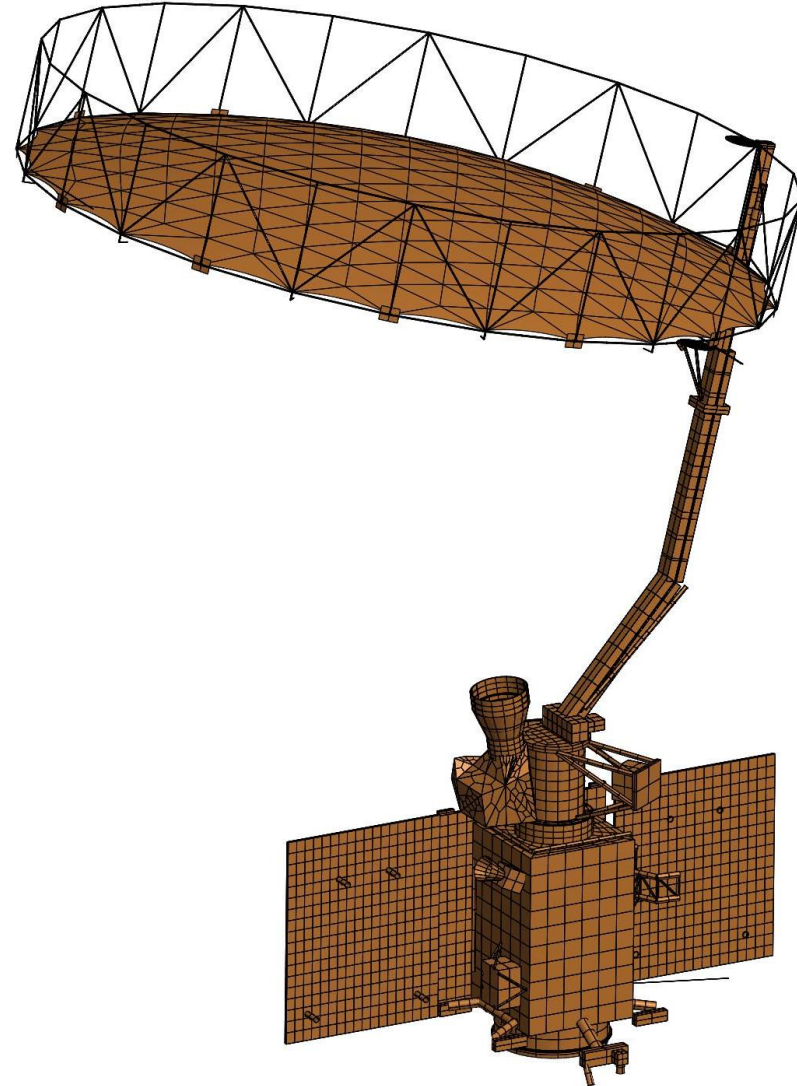
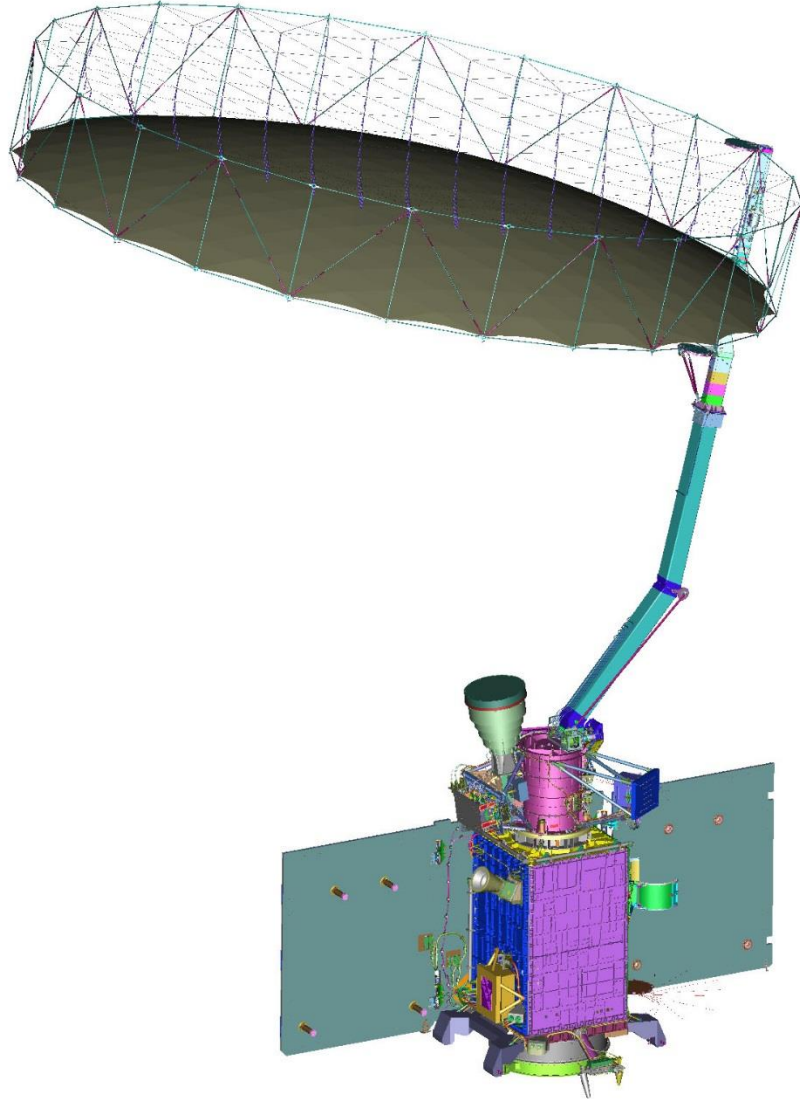


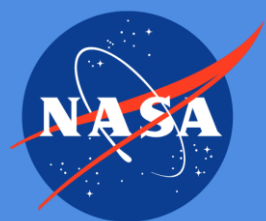


SMAP CAD & RF Models

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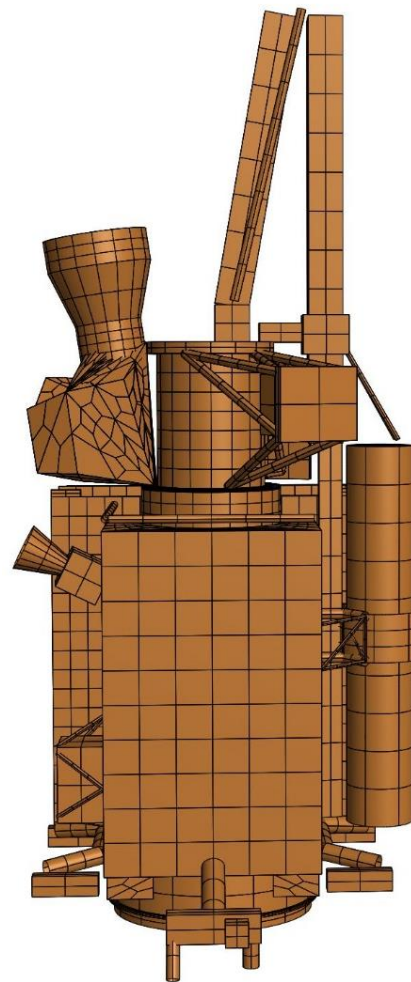
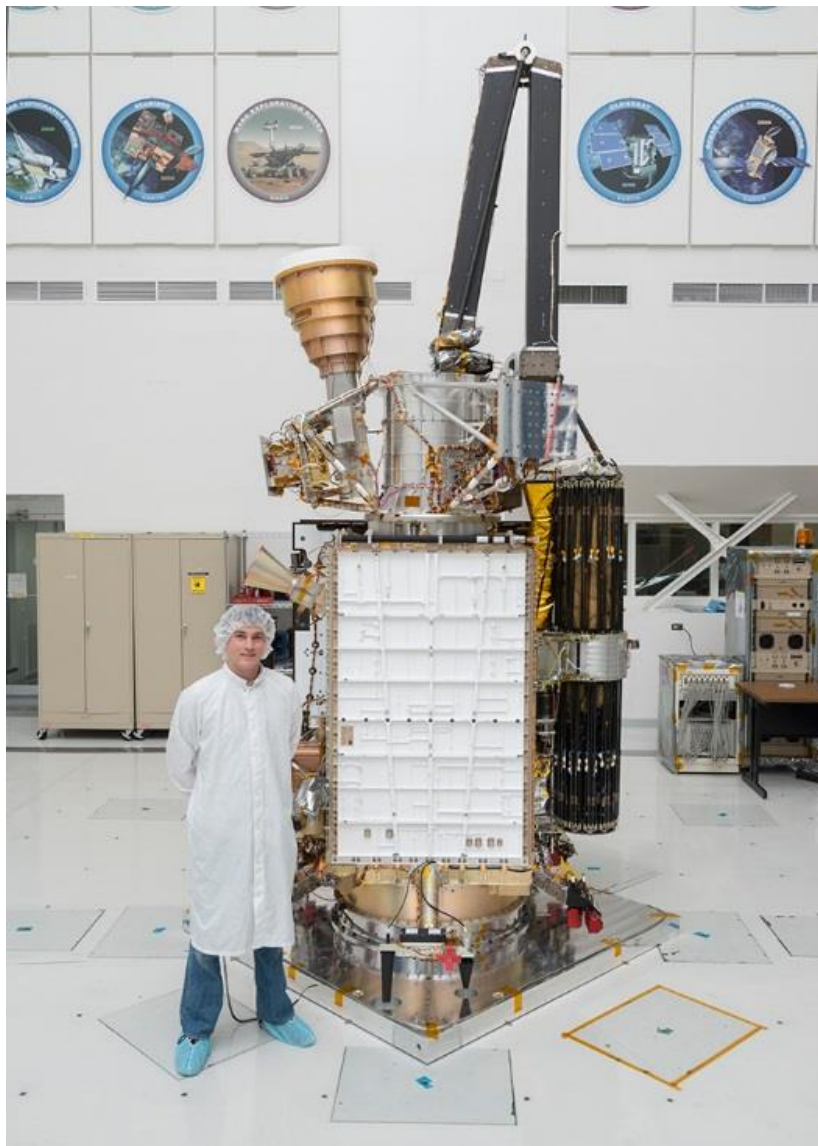


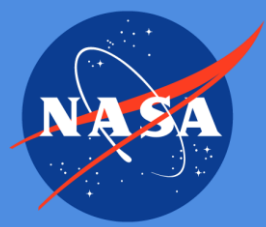


Spacecraft With Stowed AstroMesh™ Reflector

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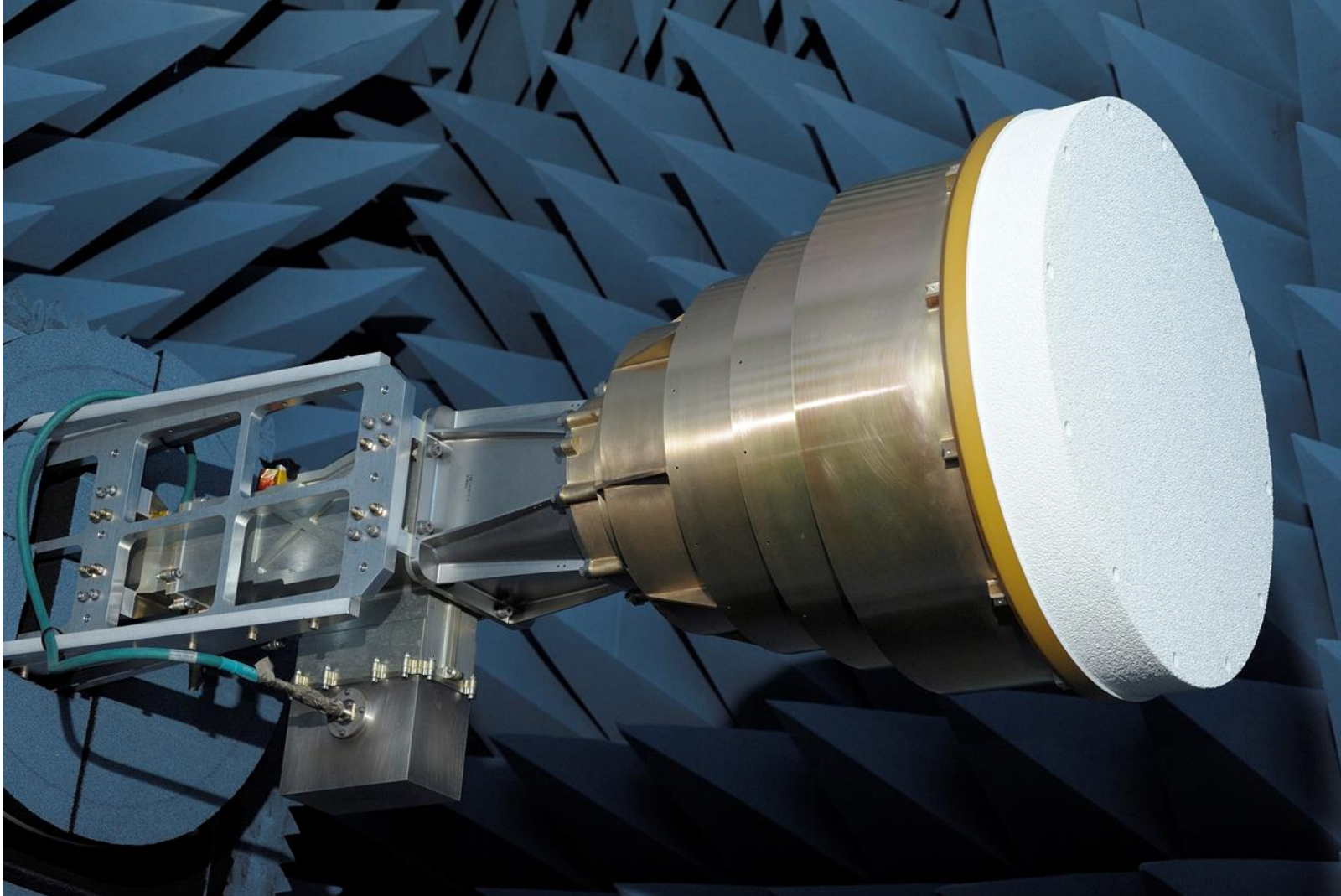


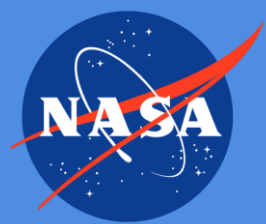


Feed Horn

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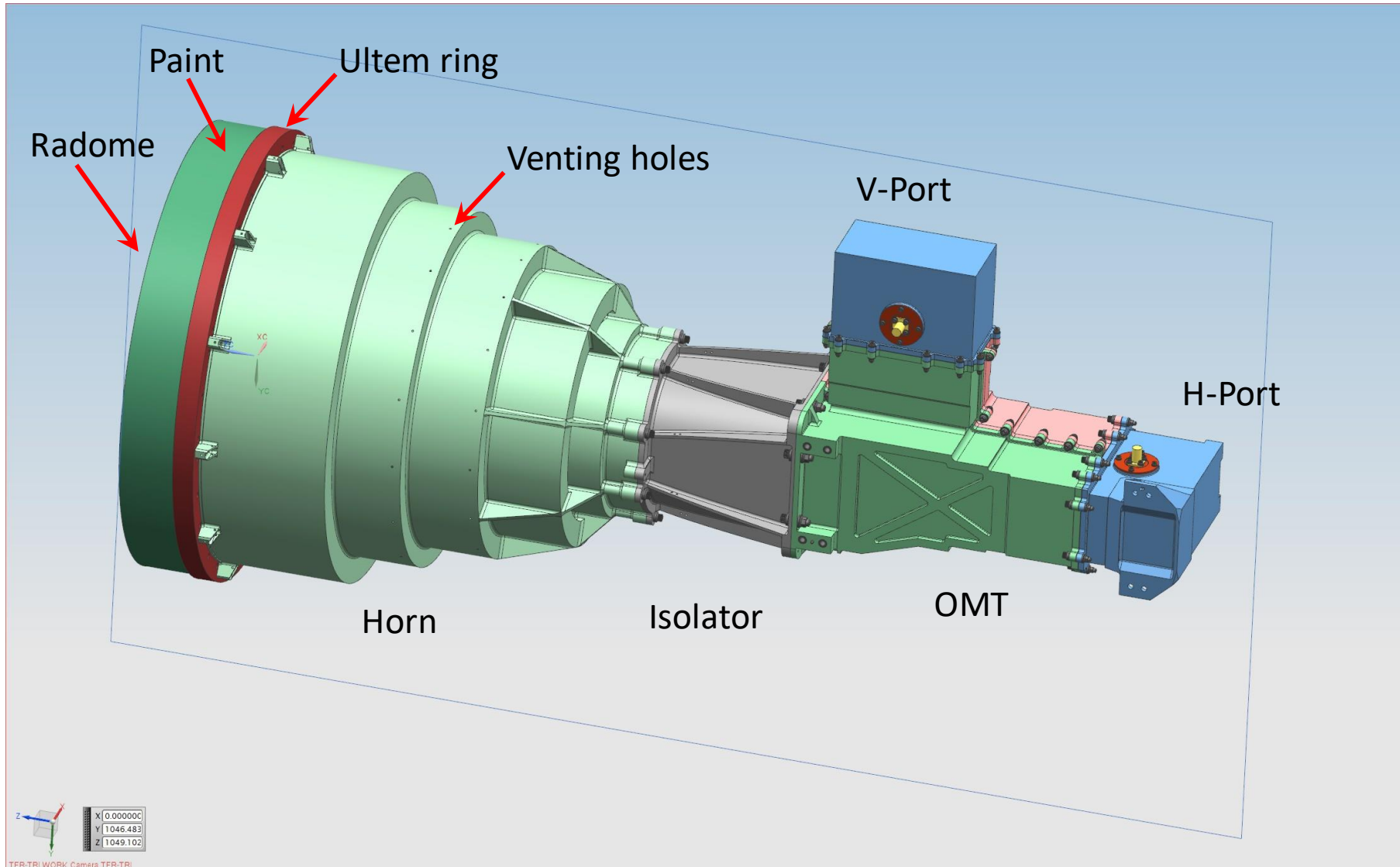


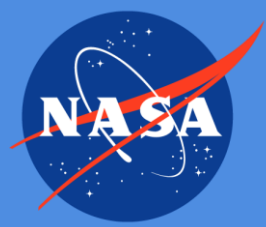


Isometric View

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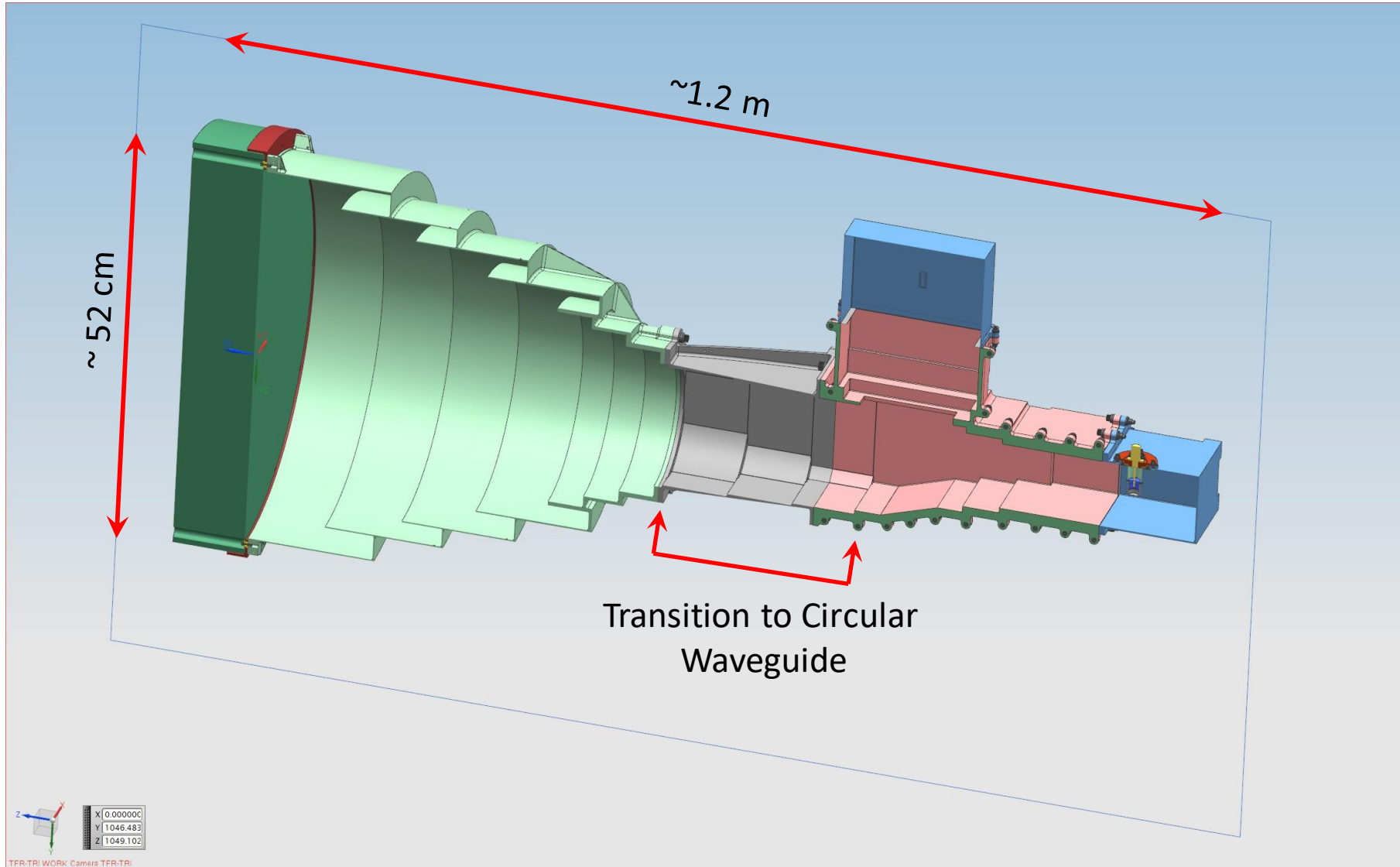


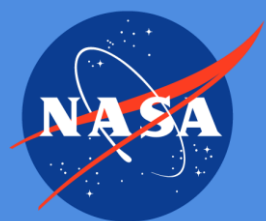


Cutaway Isometric View

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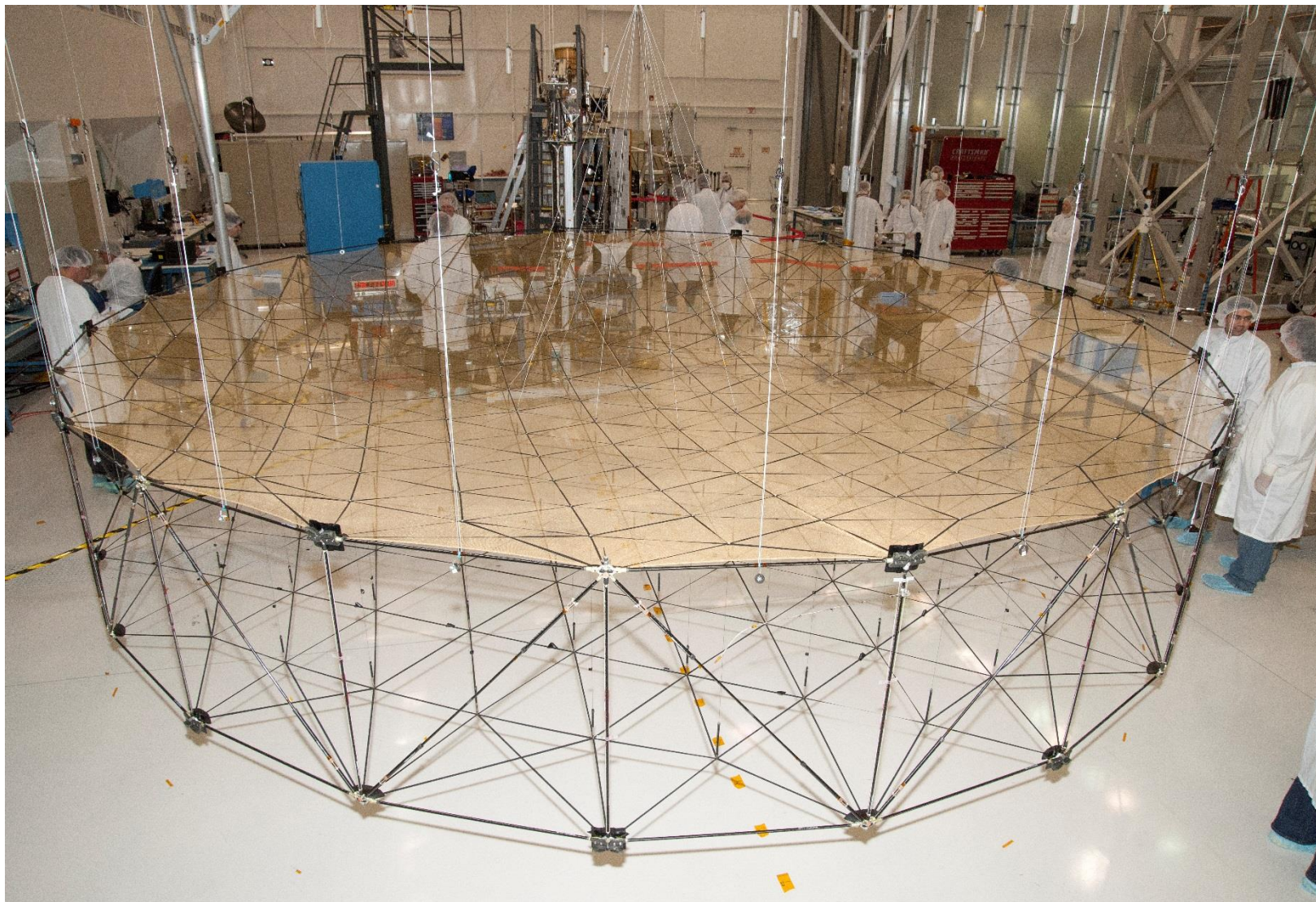




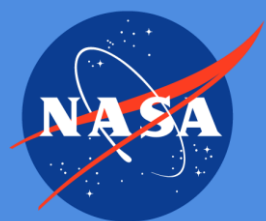
Fully Deployed AstroMesh™ Reflector

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Reflector surface RF model included a variety of small surface imperfections

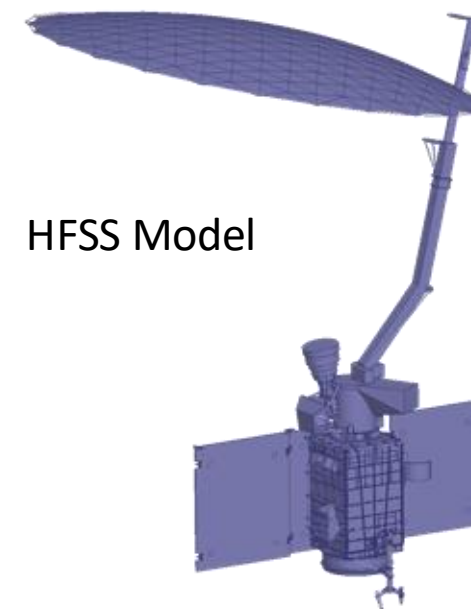
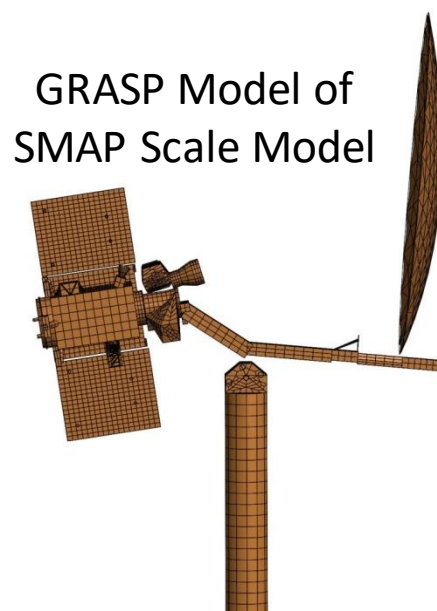
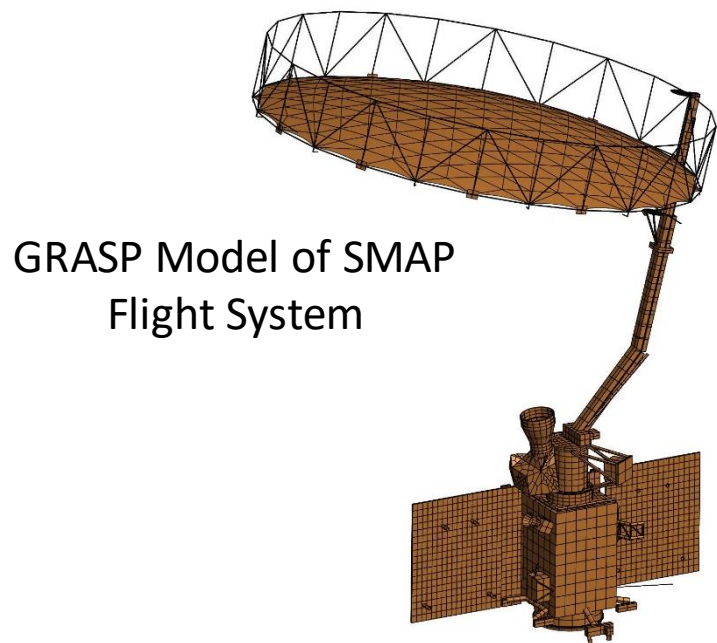


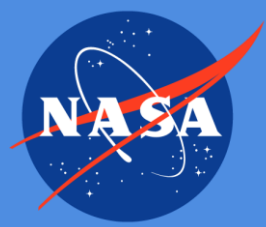
SMAP RF Modeling

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- SMAP modeling was mostly performed using the GRASP PO/MoM modeling approach.
 - PO+PTD for the reflector, MoM for everything else
- A GRASP model of the as-built scale model was also made for comparison with scale model radiation pattern measurements.
- An HFSS model was also done as an additional method of validating the GRASP model.
 - FEM for the feed assembly, IE for everything else
- The feed assembly radiation pattern was provided either in the form of an HFSS model calculation or flight feed assembly measurements when eventually available.

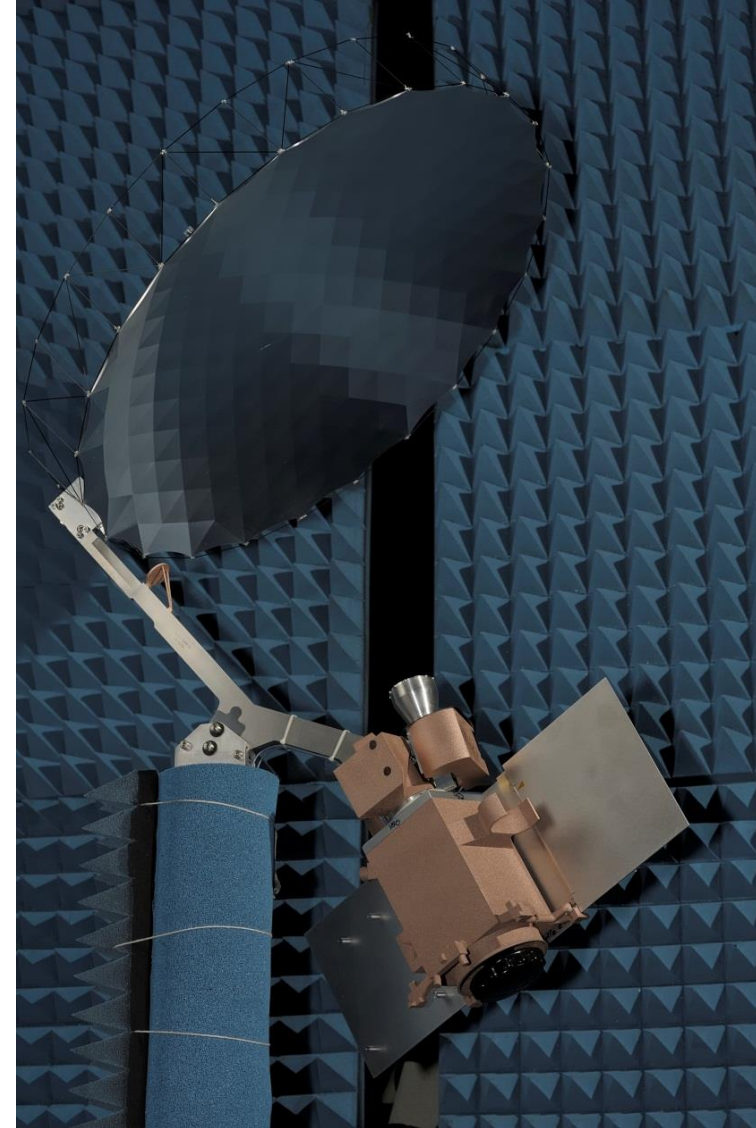
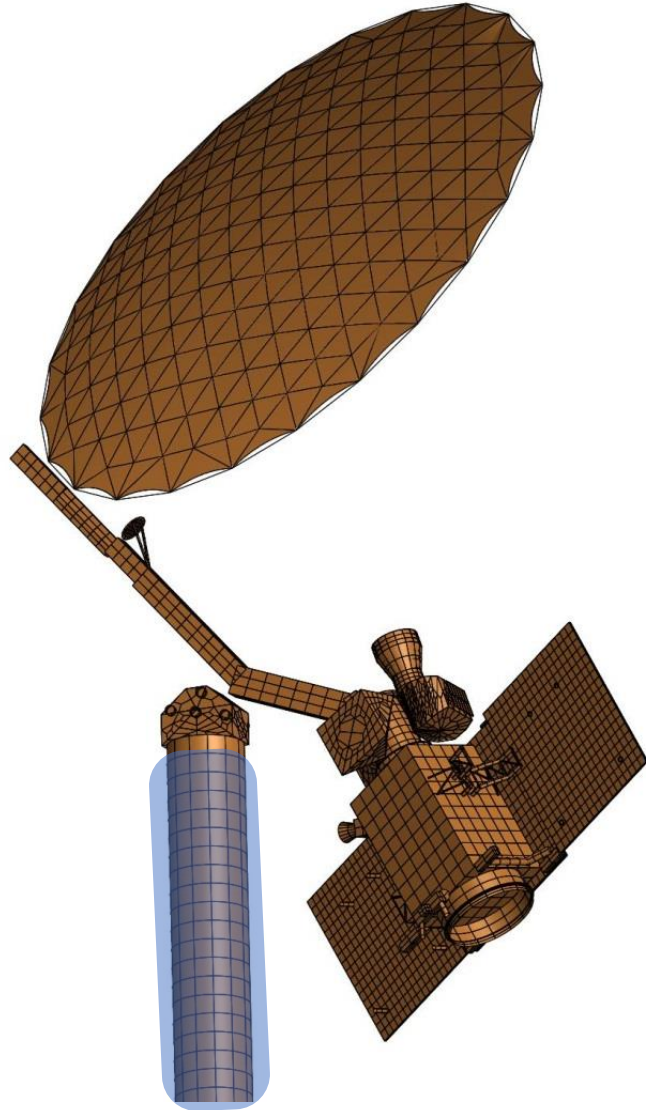


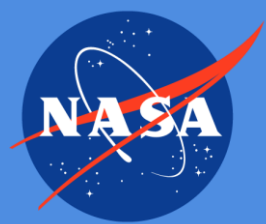


SMAP 1/10th Scale Model

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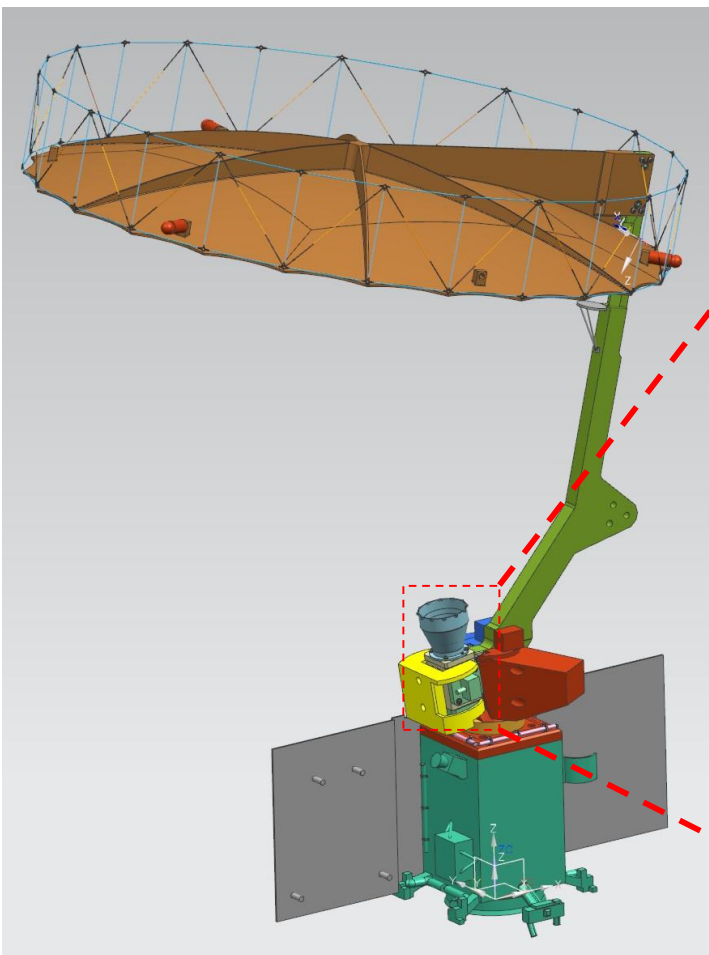


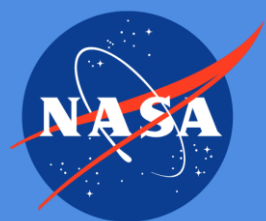


SMAP Scale Model

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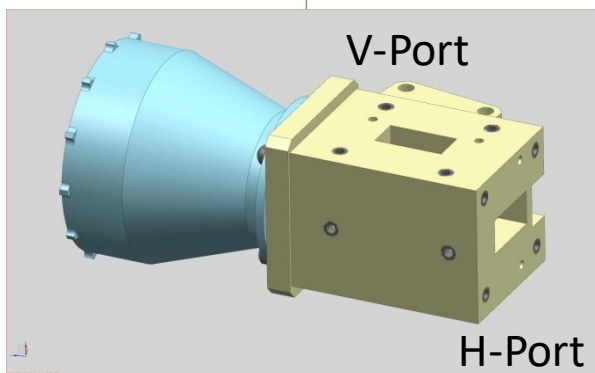
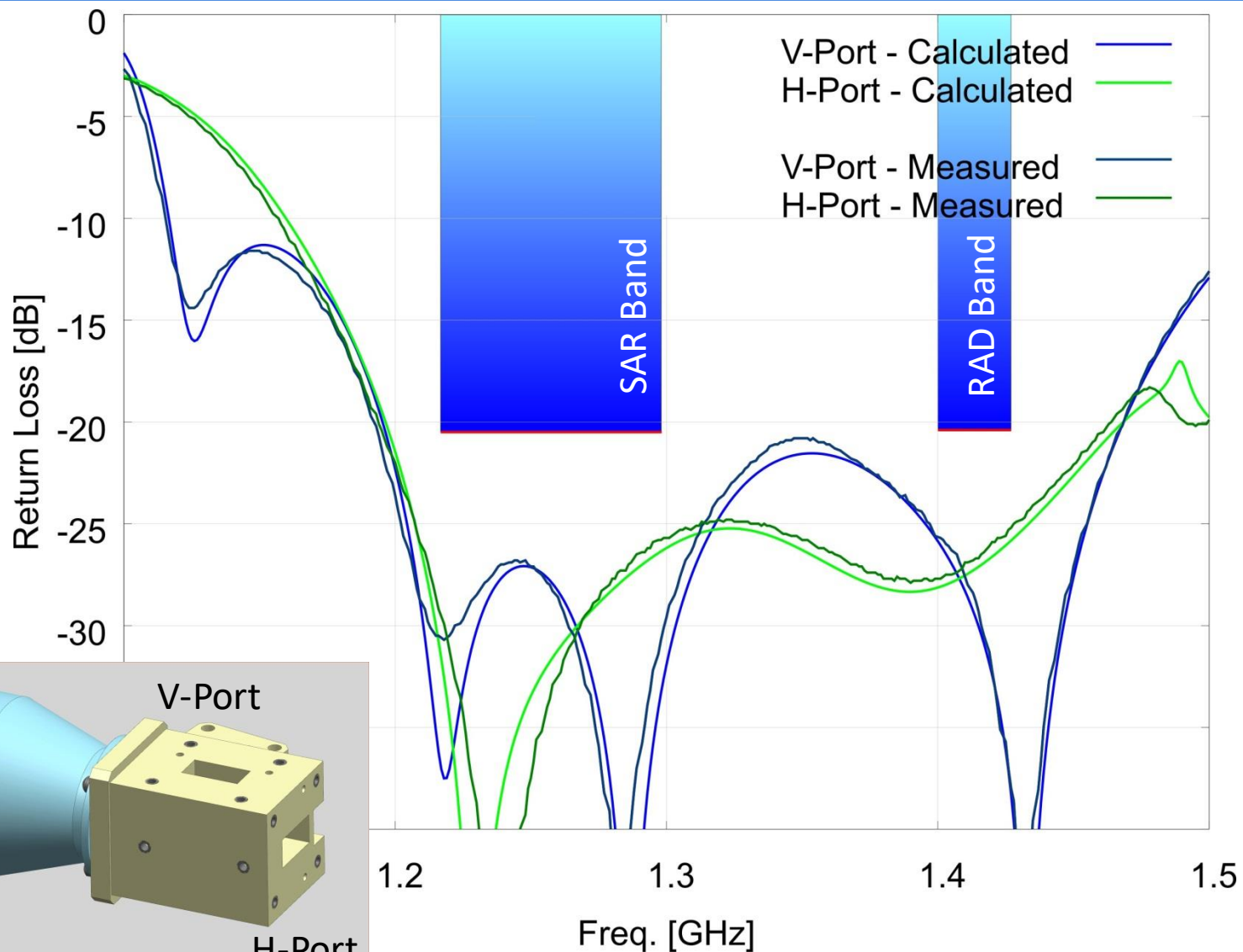


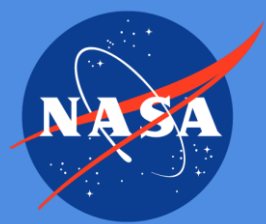


Feed Horn RL Into SM OMT

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Radiation Pattern Components

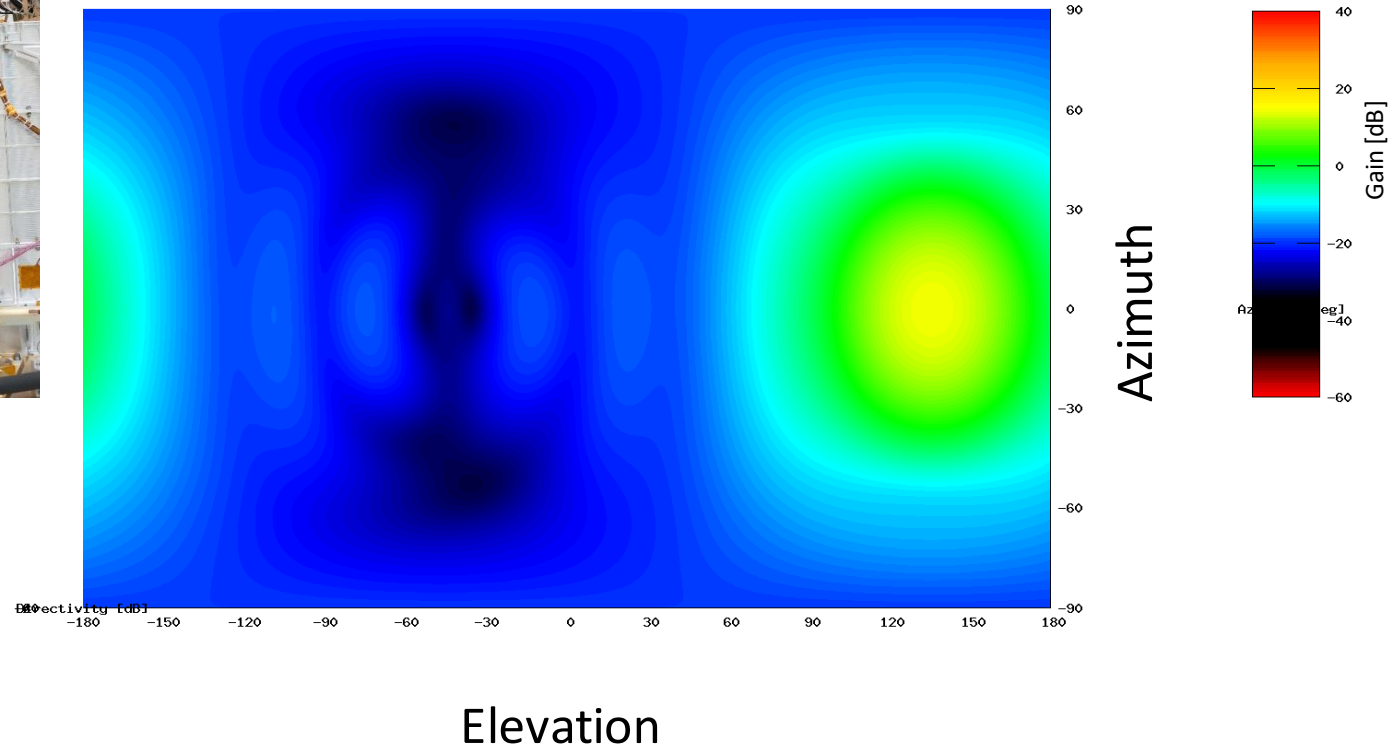
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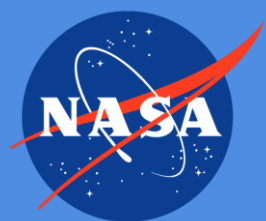
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Measured Feedhorn Pattern

SAR V-Pol



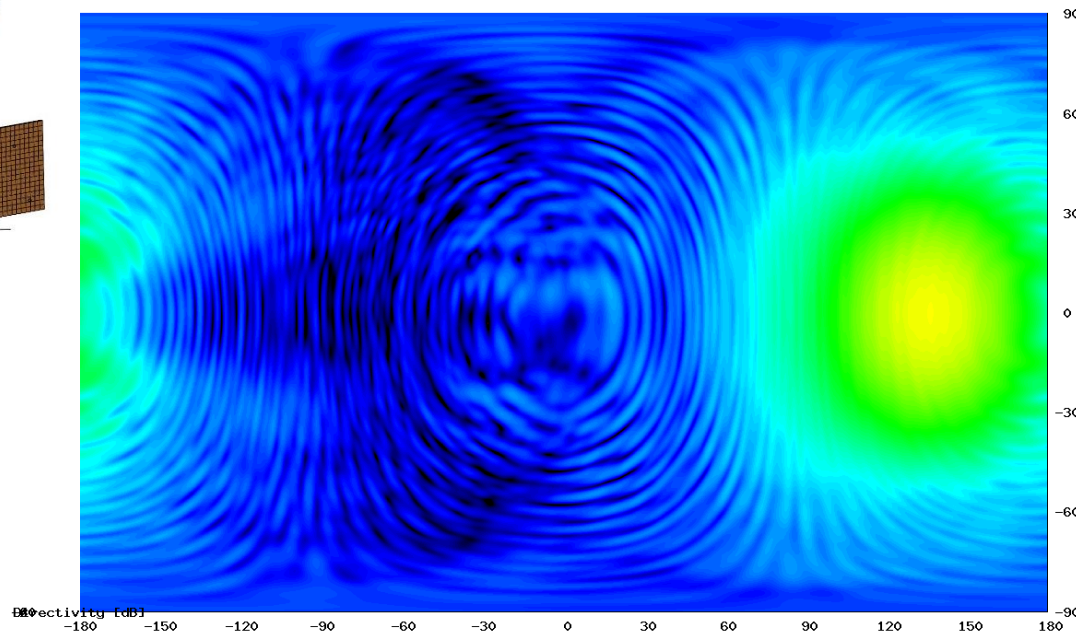
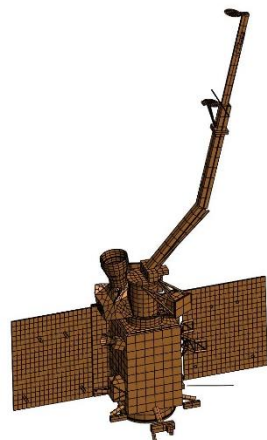


Radiation Pattern Components

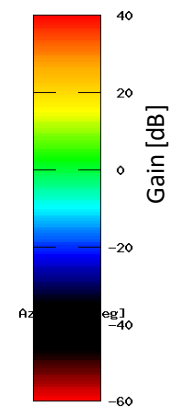
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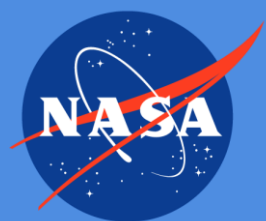
Measured Feedhorn Pattern + Scattering From S/C & Boom



SAR V-Pol



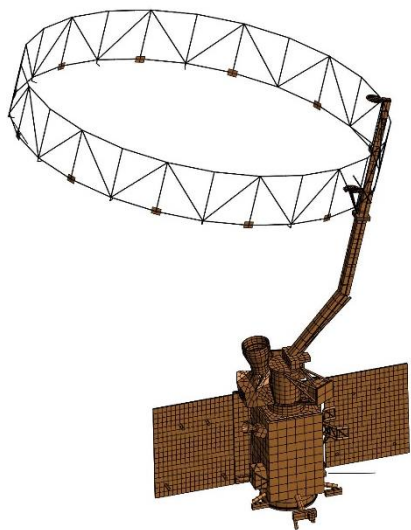
Elevation



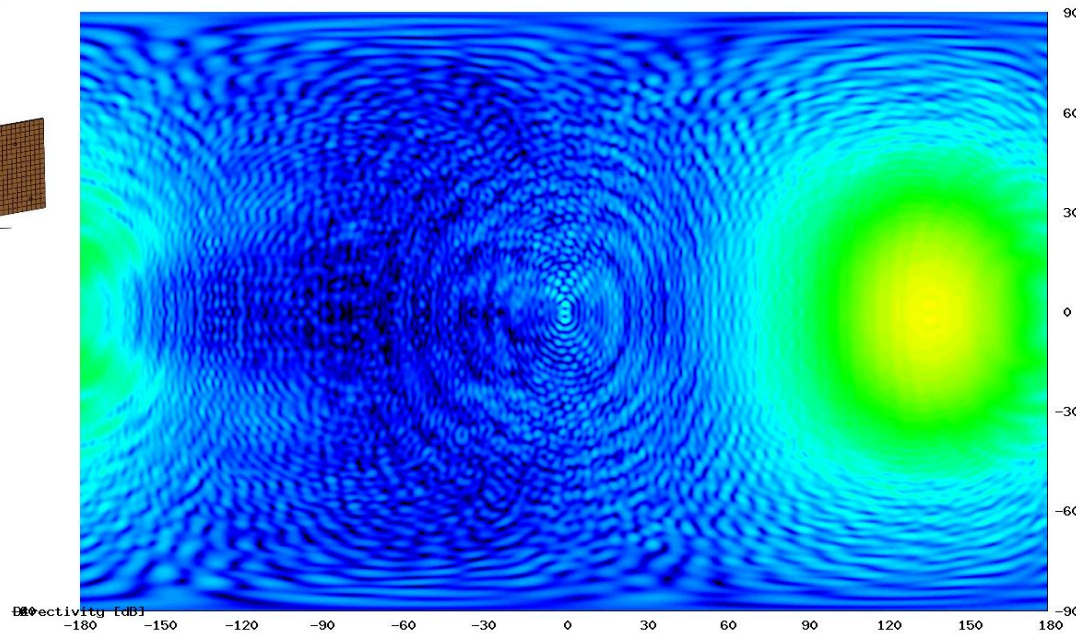
Radiation Pattern Components

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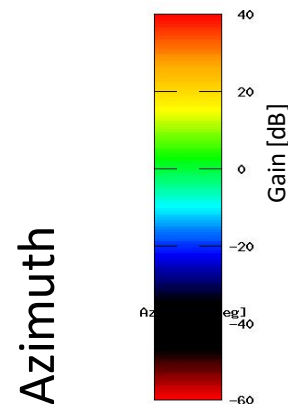
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Measured Feedhorn Pattern + Scattering From S/C, Boom & Truss



SAR V-Pol



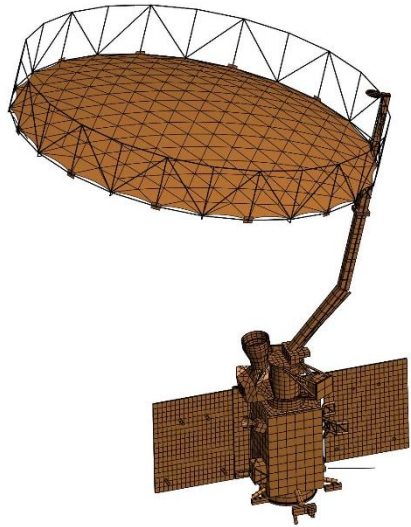
Elevation



Radiation Pattern Components

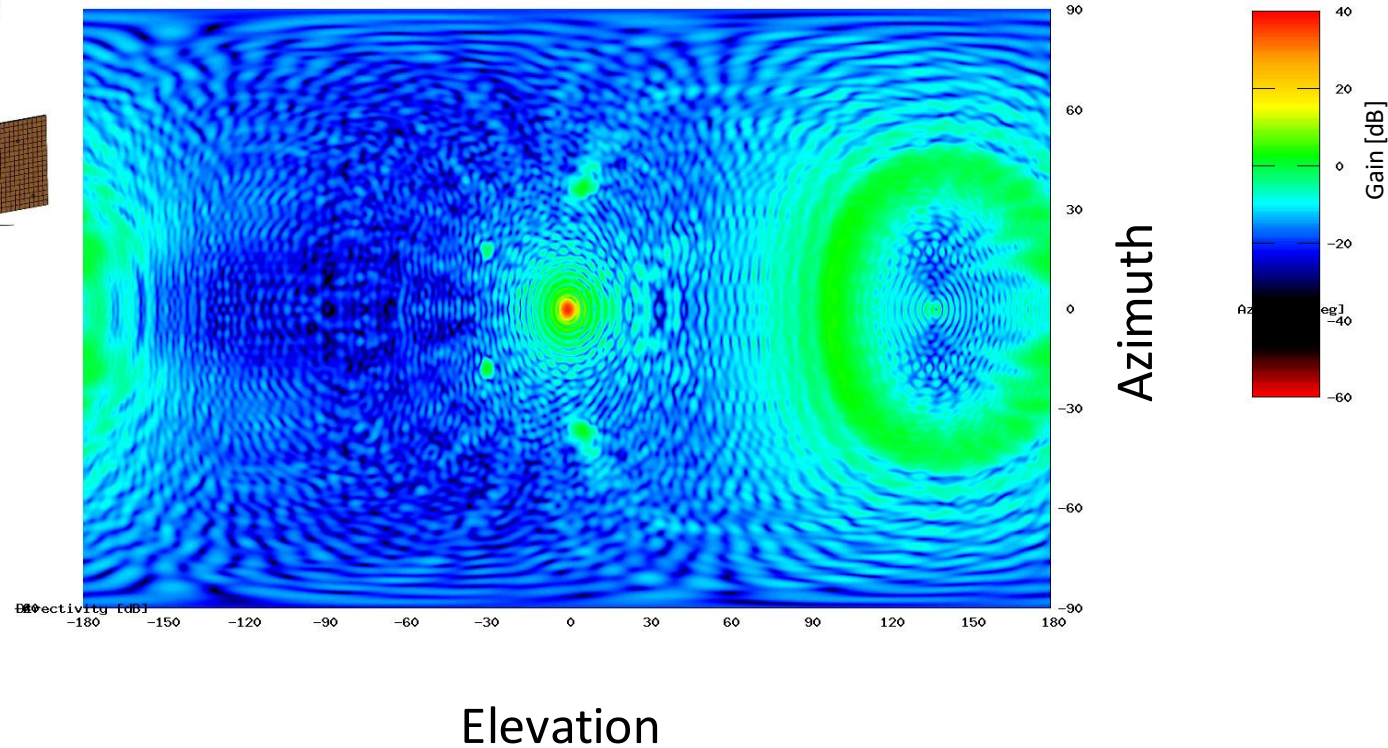
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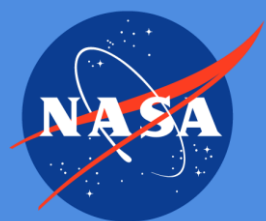
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Measured Feedhorn Pattern +
Scattering from S/C, Boom & Truss +
Reflector

SAR V-Pol





Radiation Pattern Components

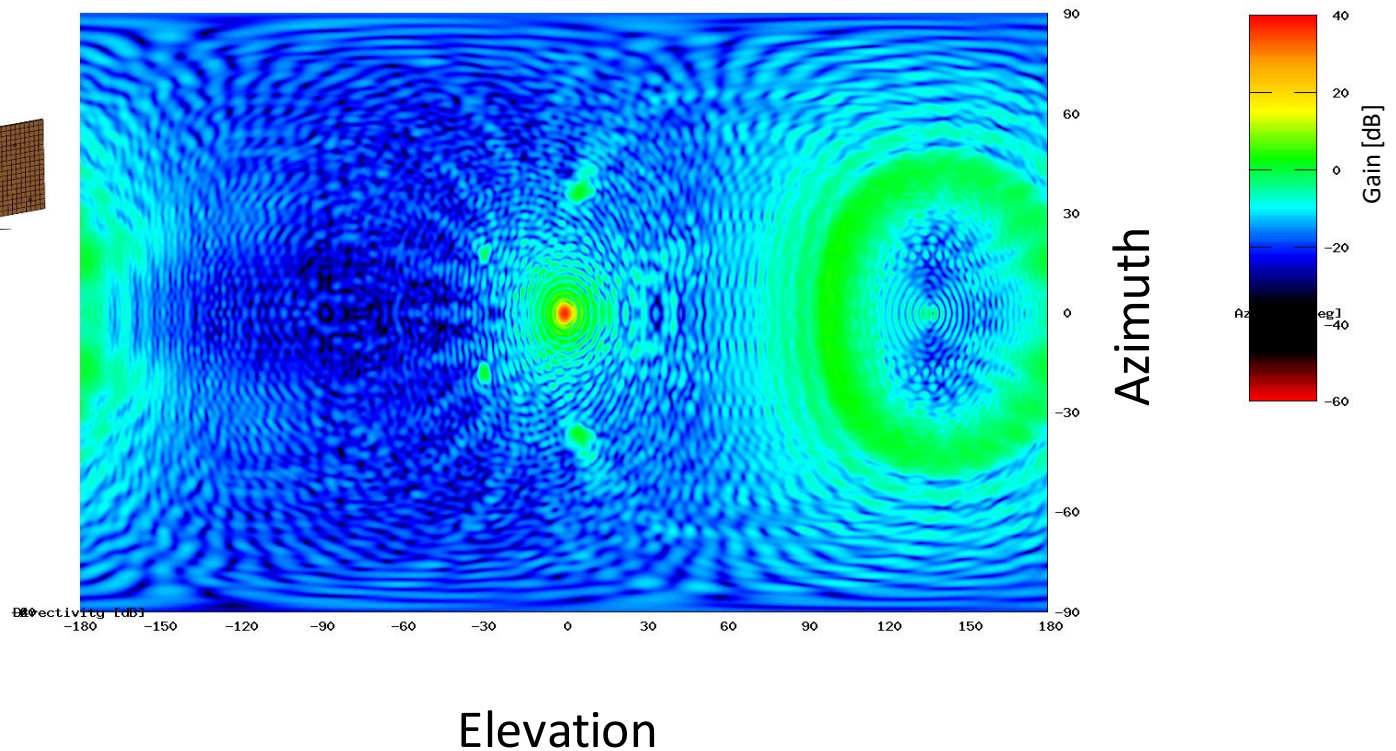
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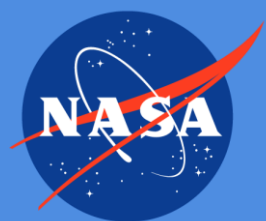
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Measured Feedhorn Pattern +
Scattering From S/C, Boom & Truss +
Reflector + 2nd Order Scattering

SAR V-Pol



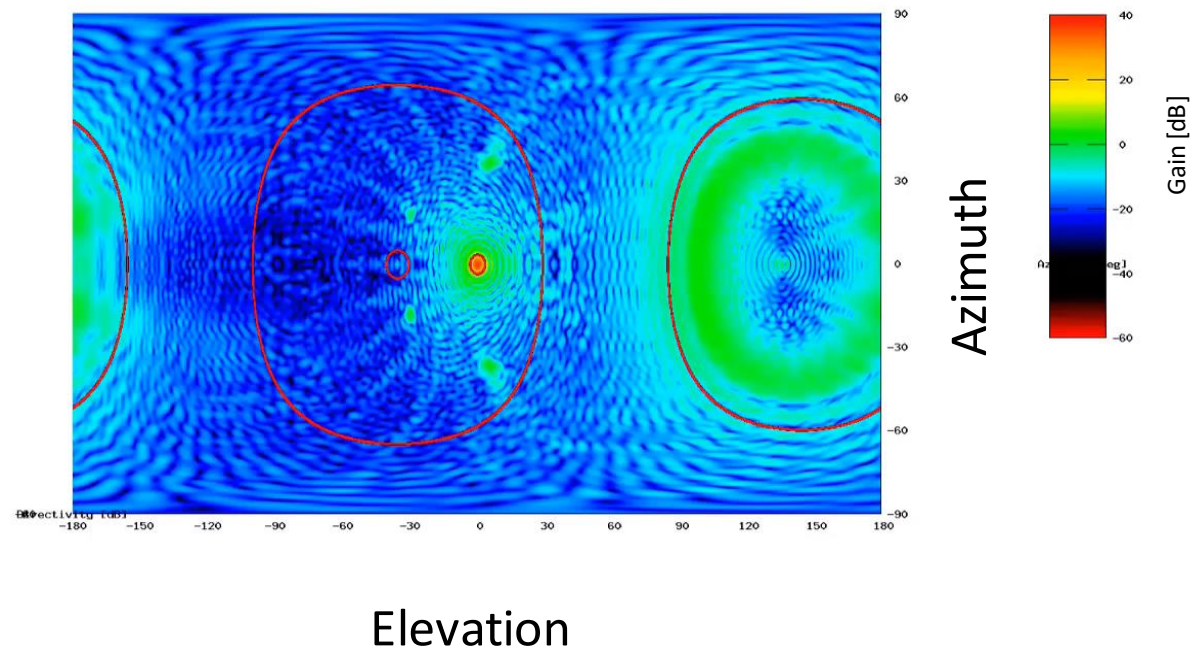


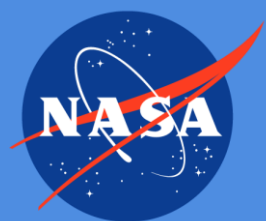
Effect of the Spinning Reflector

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SAR V-Pol

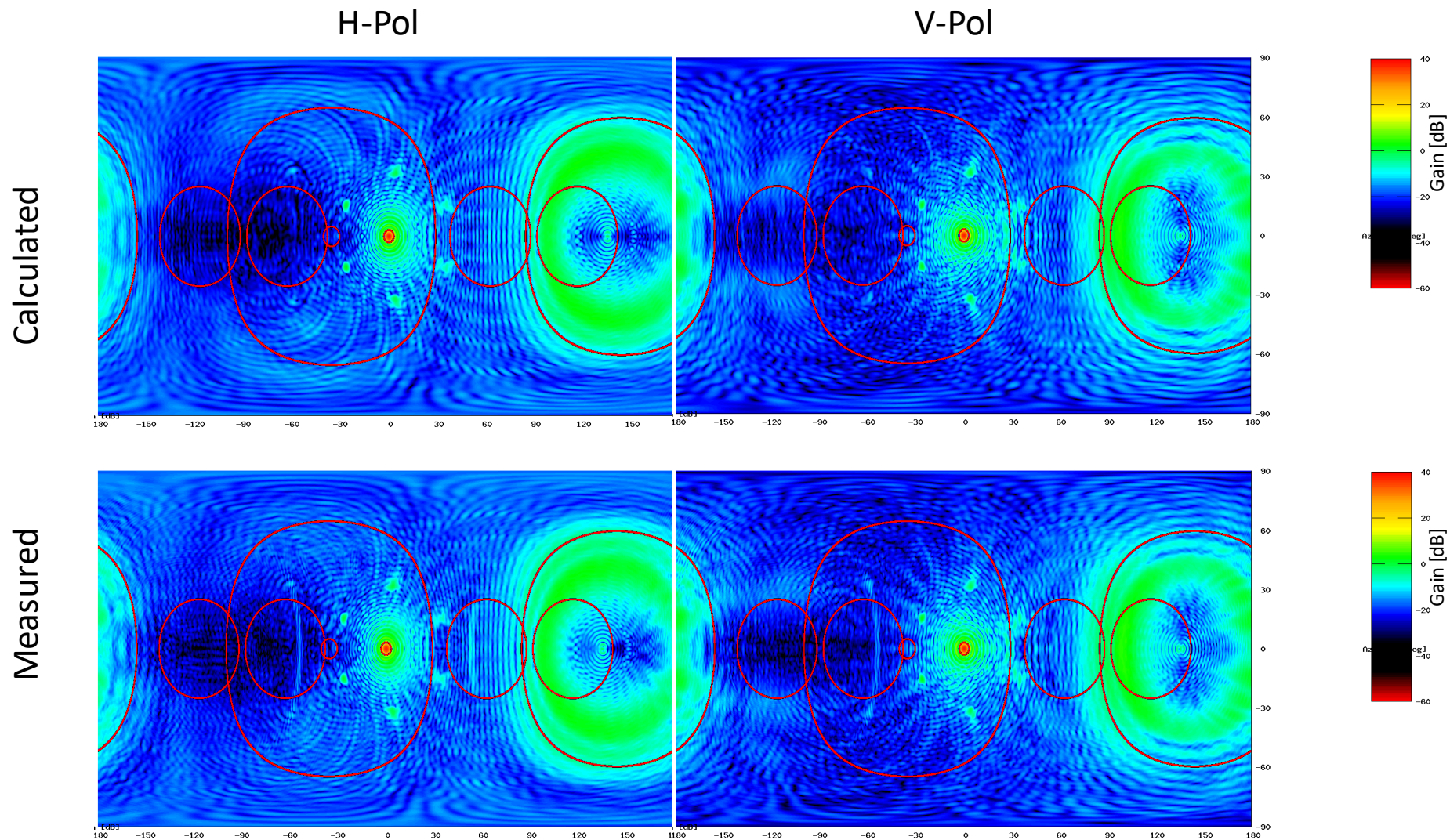


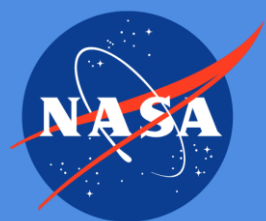


Radiation Pattern Comparison: RAD, 000

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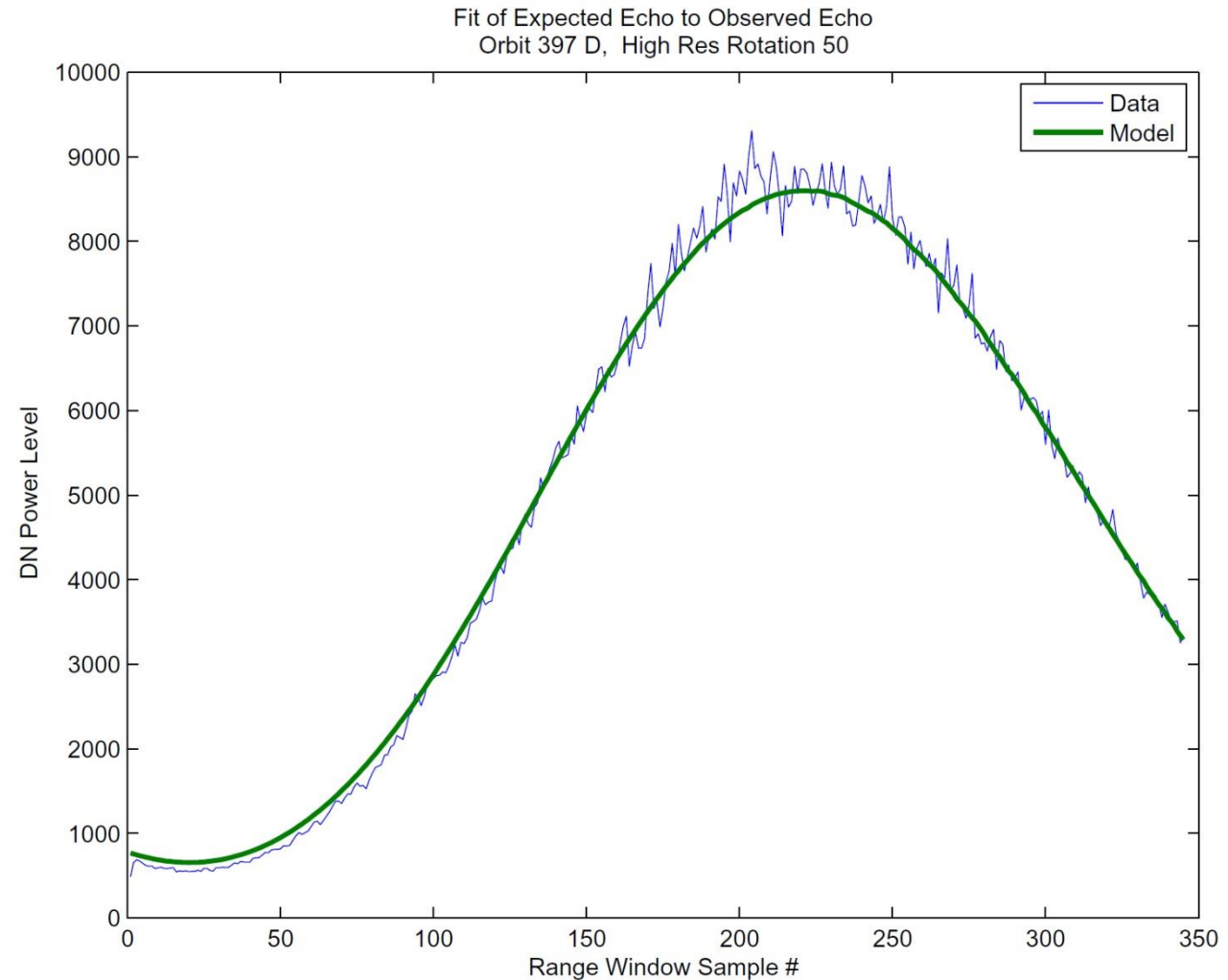


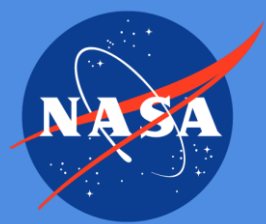
Measurement of Radar Echo From Space

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- Calculated radiation pattern could predict on-orbit performance extremely well.
- SAR pointing was predicted to be 0.270° from nominal, according to the RF model.
- With the help of corner reflectors on the ground, it was measured to be 0.291° , only 0.021° off from our calculations.



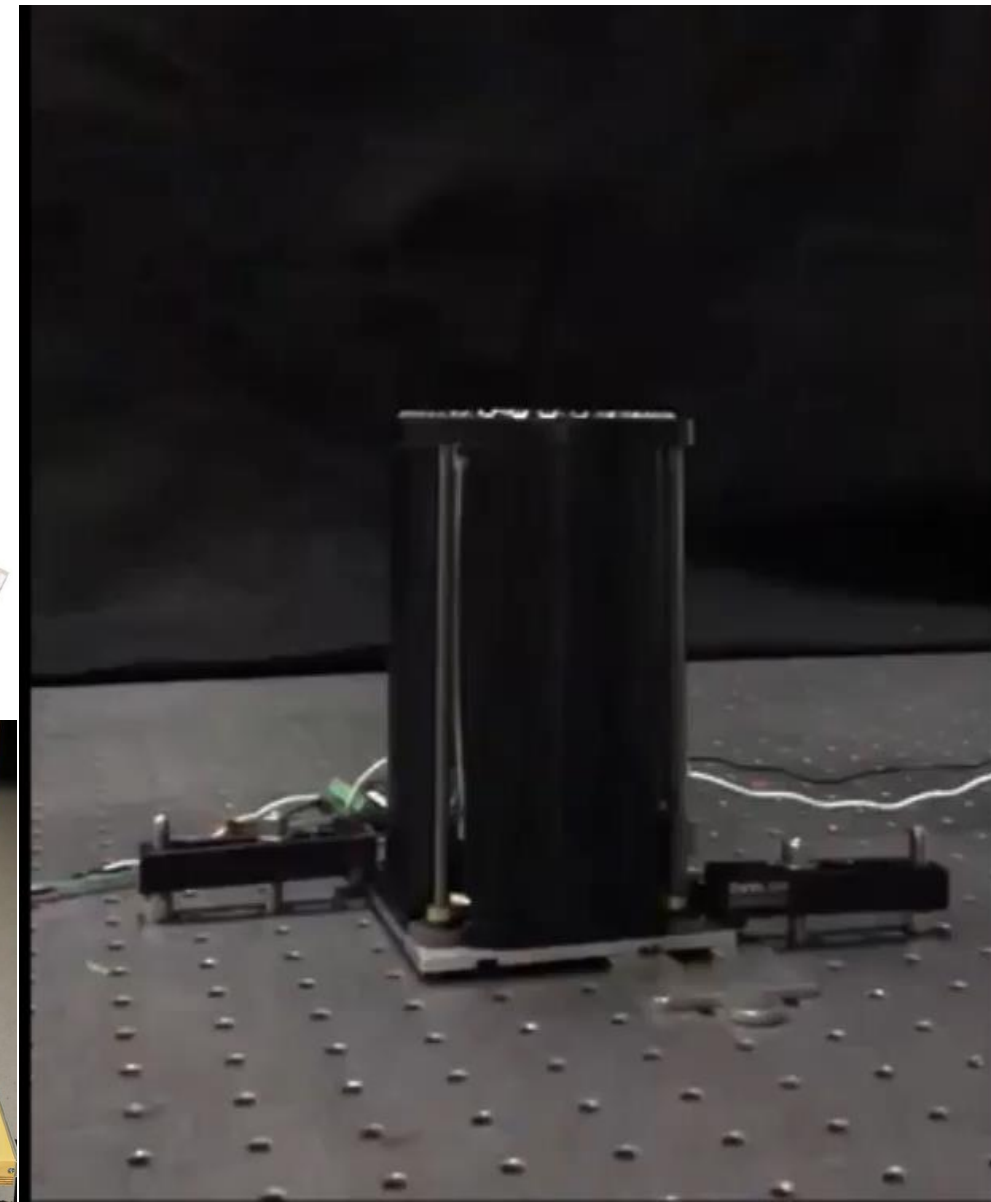
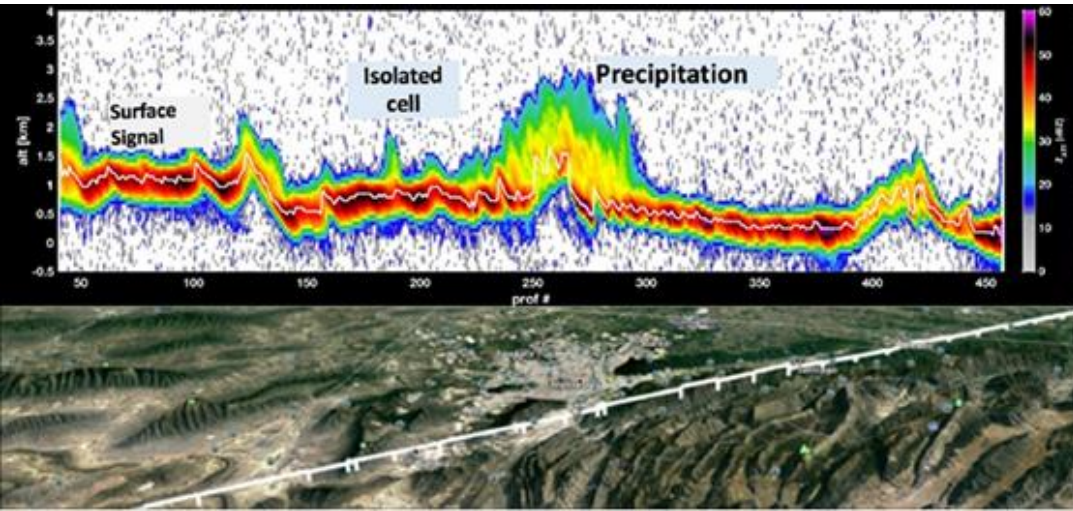


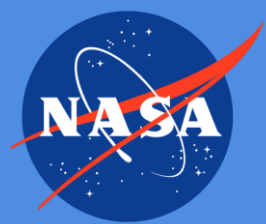
Raincube, 2018

- ### Raincube Antenna
- 0.5m Cassegrain reflector
 - Ka-band operations

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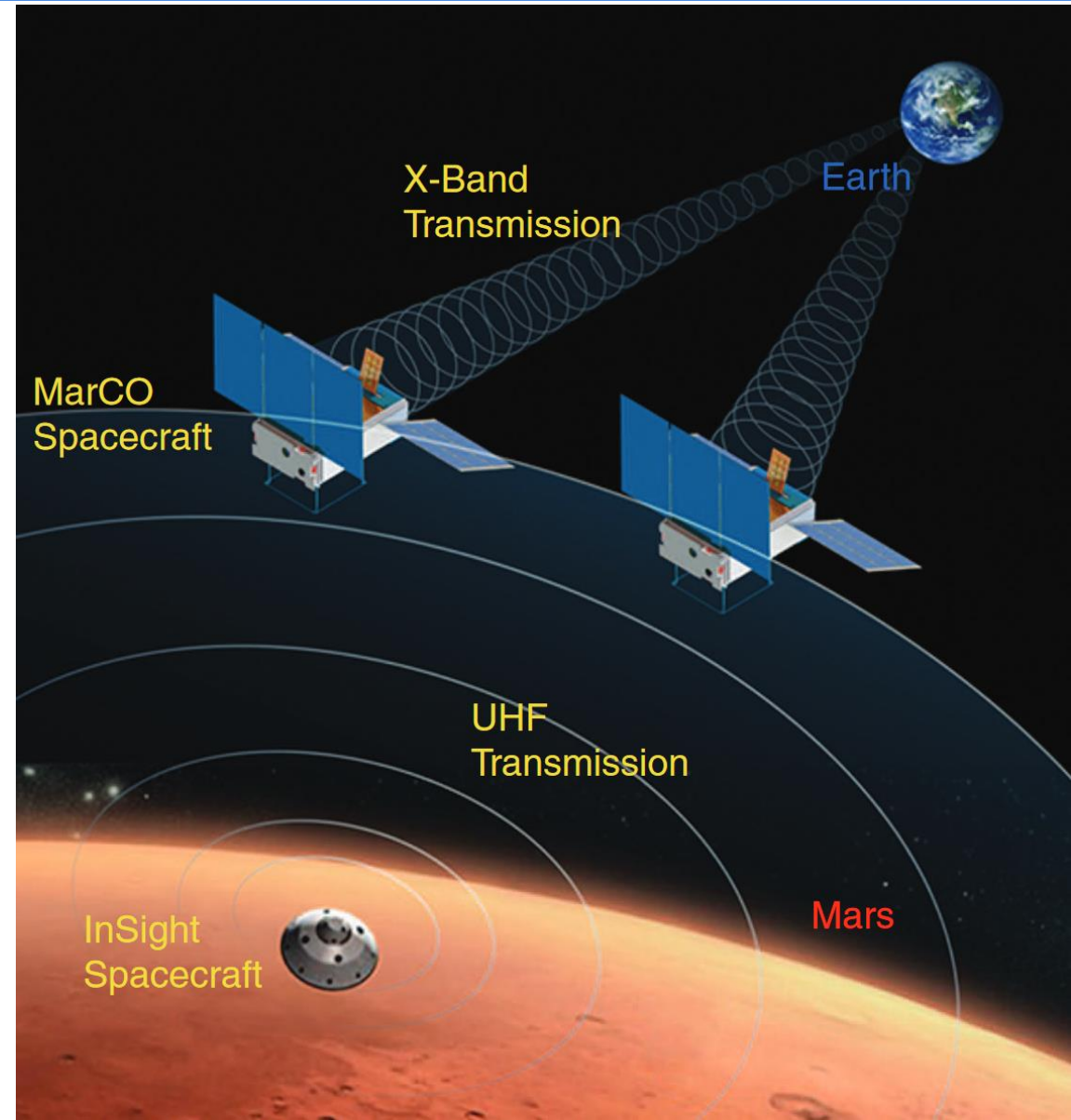
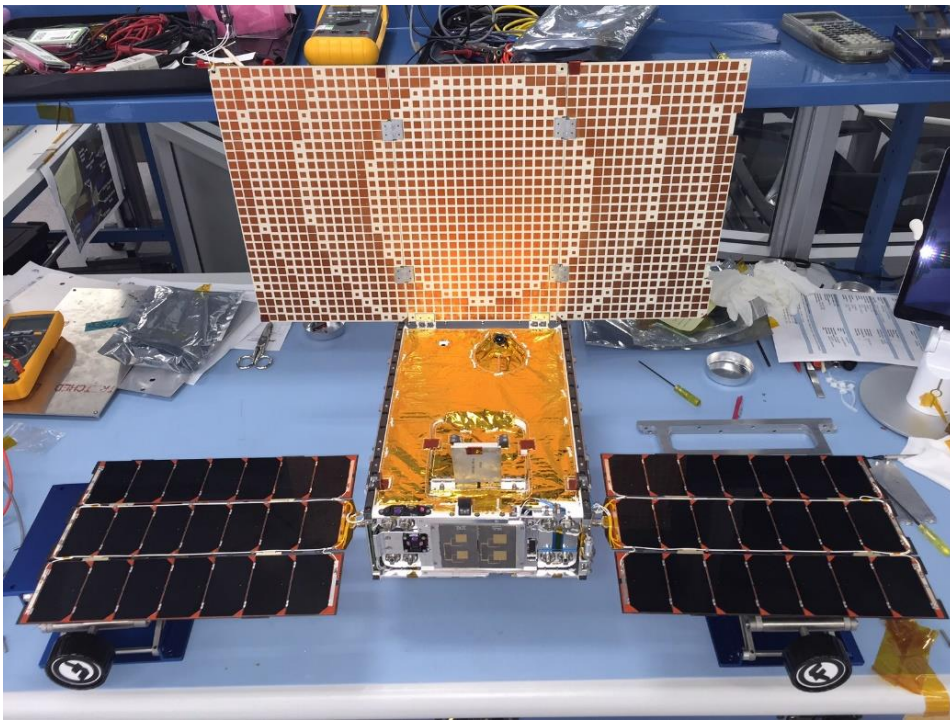
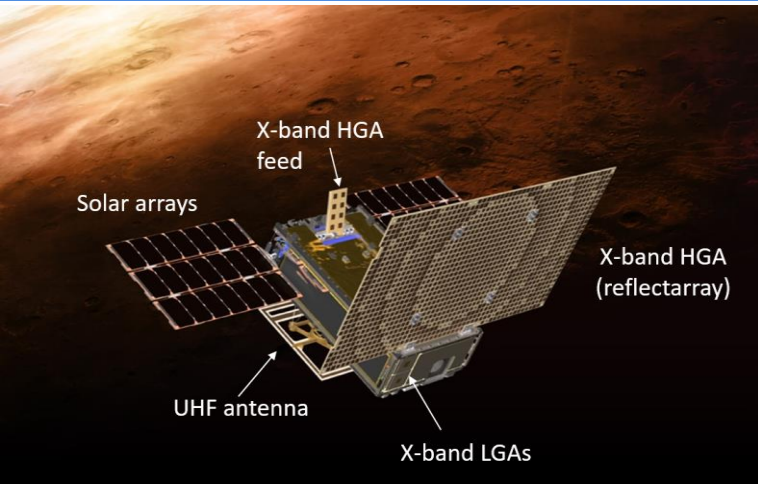




MarCO, Mars Cubesat One, 2018

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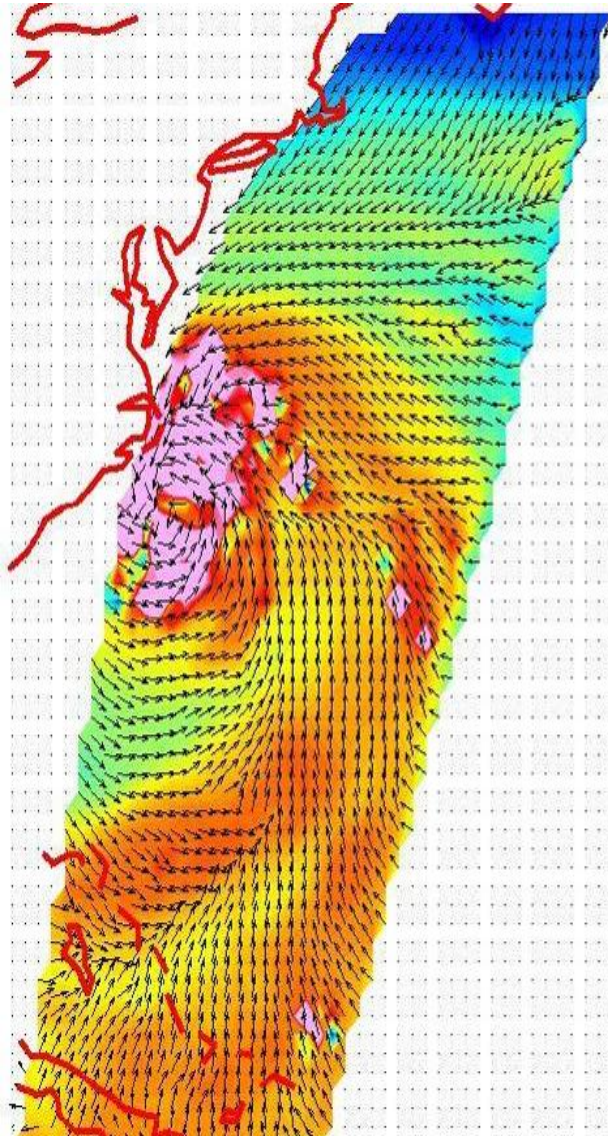




Compact Ocean Wind Vector Radiometer (COWVR), 2021

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Lance Milligan
Mechanical/Thermal
System Lead

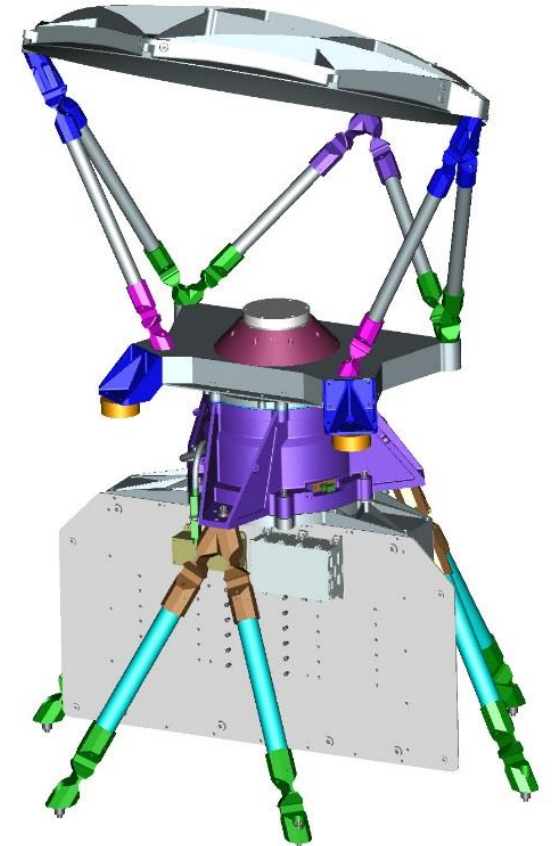
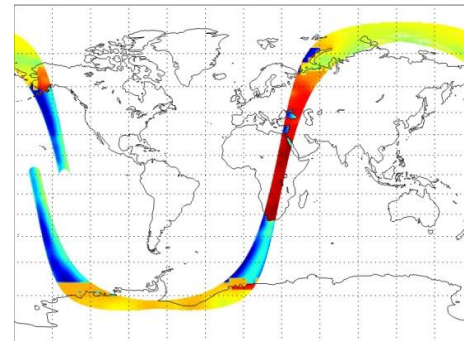
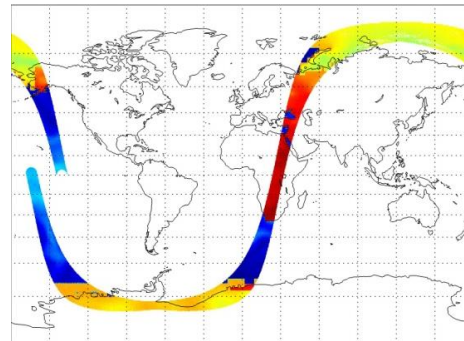
Shannon Brown
Principal Investigator

Sharmila Padmanabhan
Instrument
Performance Engineer

Richard Redick
Instrument System
Engineer

Ami Kitiyakara
Project Manager

Paolo Focardi
Antenna RF Lead



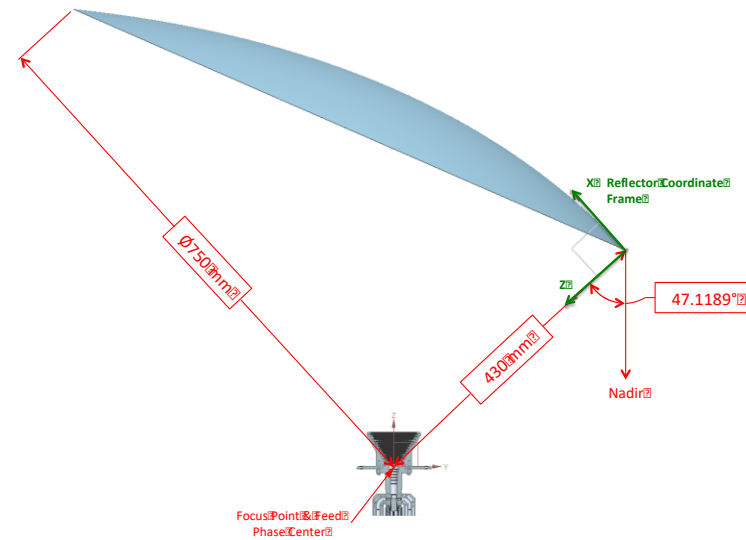
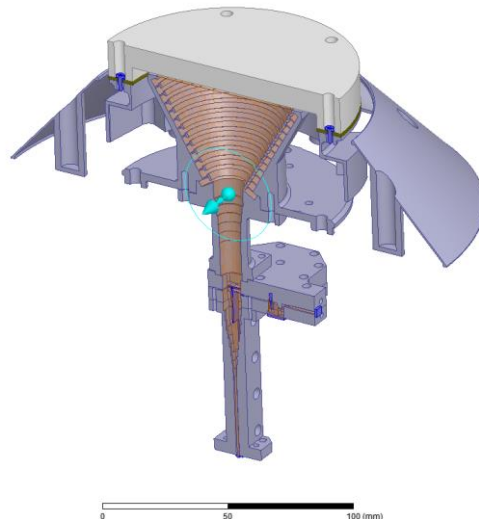


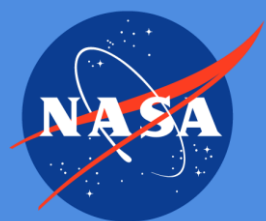
COWVR

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- COWVR's mission is to measure the wind velocity vector over the ocean.
- Launched on December 21, 2021, it is currently operating on the ISS.
- The instrument consists of:
 - 75-cm offset solid reflector.
 - A single tri-band corrugated horn (18.7 GHz, 23.8 GHz, and 33.9 GHz).
 - Feed design was inherited from Jason 3.
 - Feed horn and radiometer are fixed; reflector spins around feed axis.





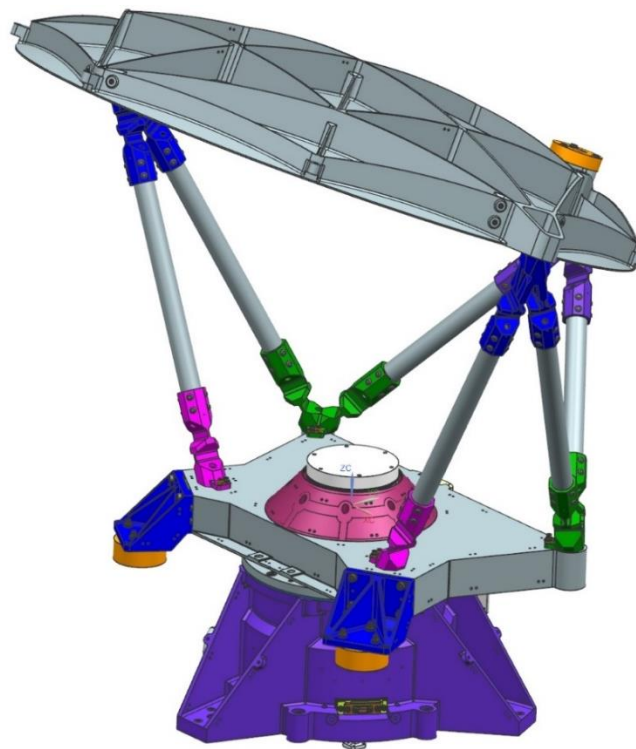
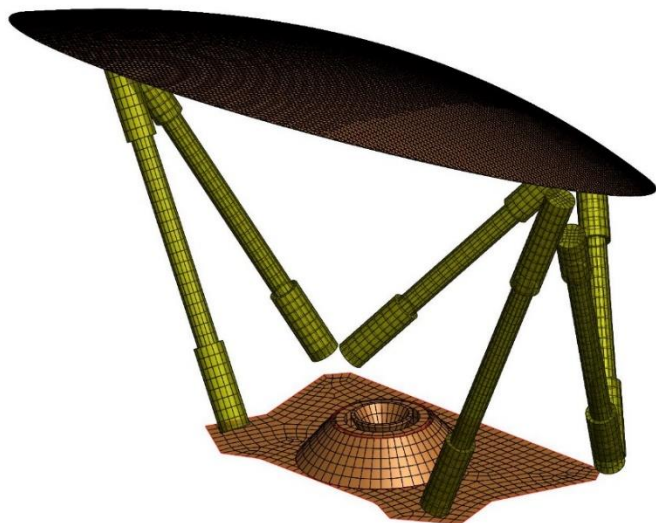
Initial RF Model

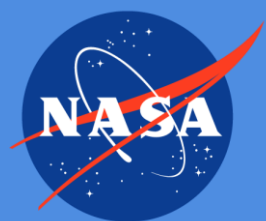
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The initial RF Model in GRASP included:

- As-built reflector surface from photogrammetry data
- As-built position and orientation of the feed
- Calculated feed pattern from HFSS model
- Top deck and struts with simplified geometry

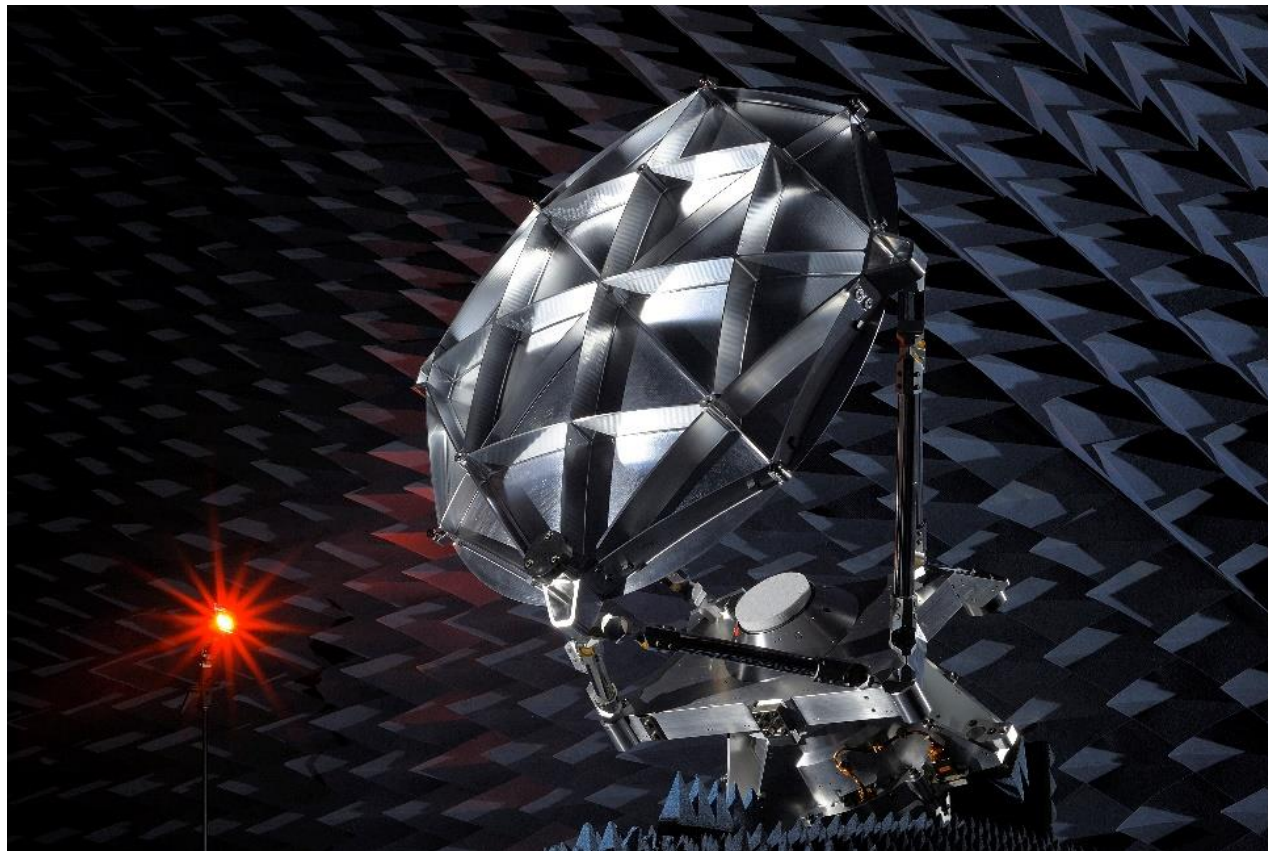


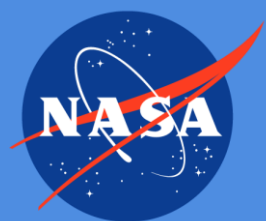


COWVR Instrument During RF Tests

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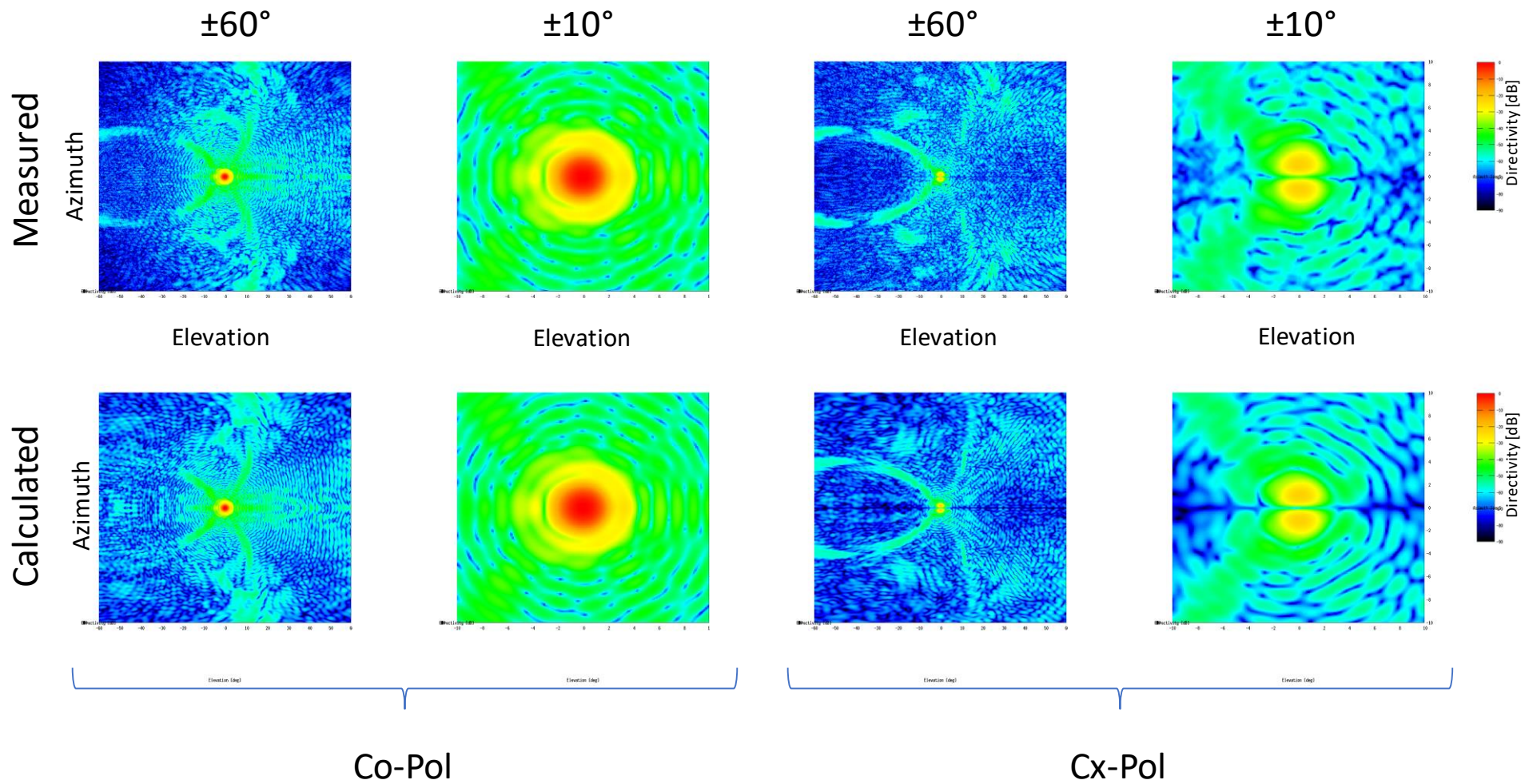


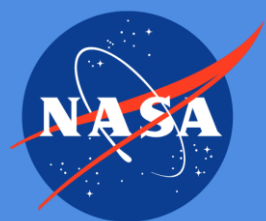


H-Pol Results, 180°, 18.7 GHz

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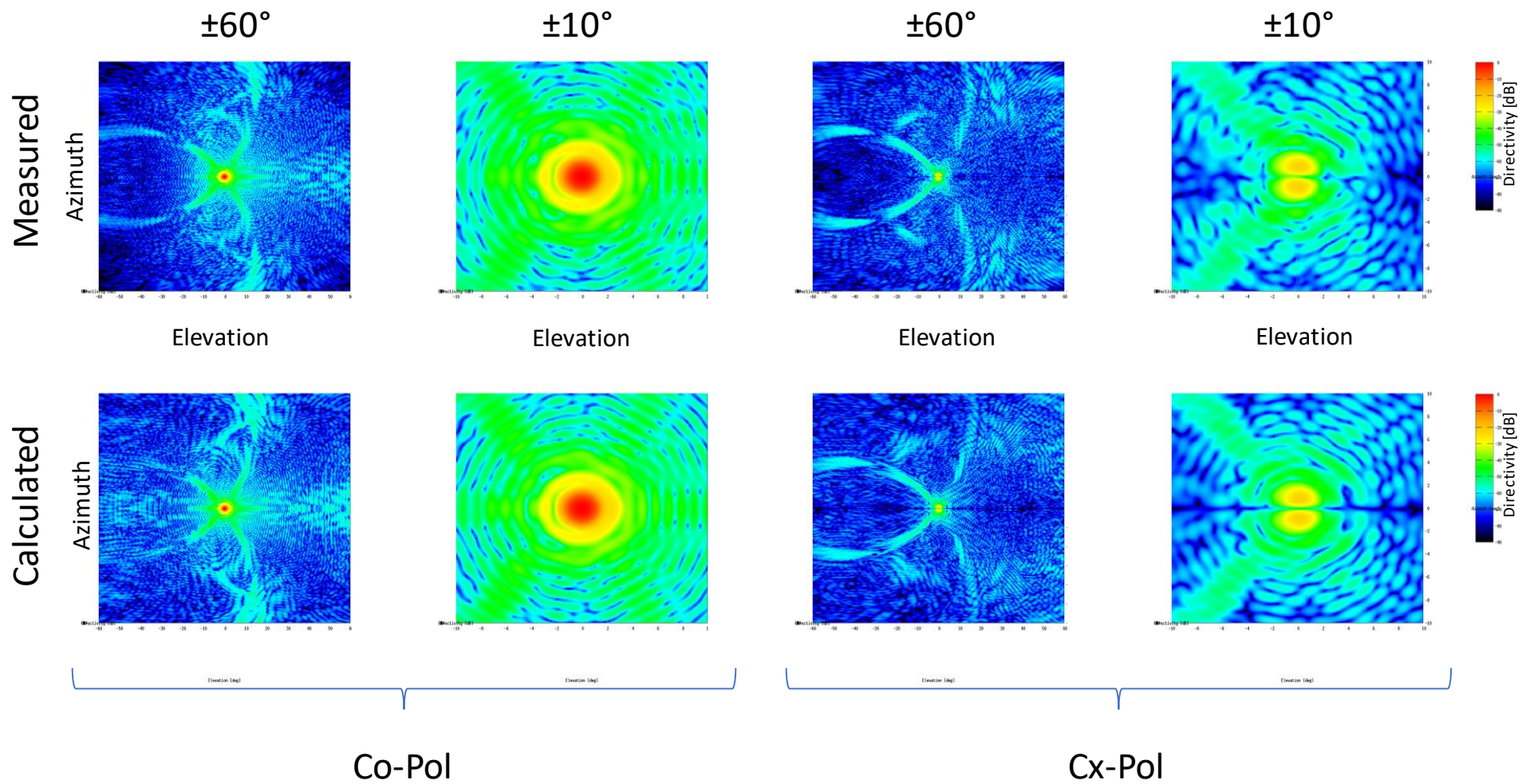




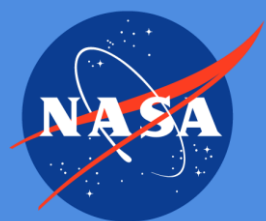
H-Pol Results, 180°, 23.8 GHz

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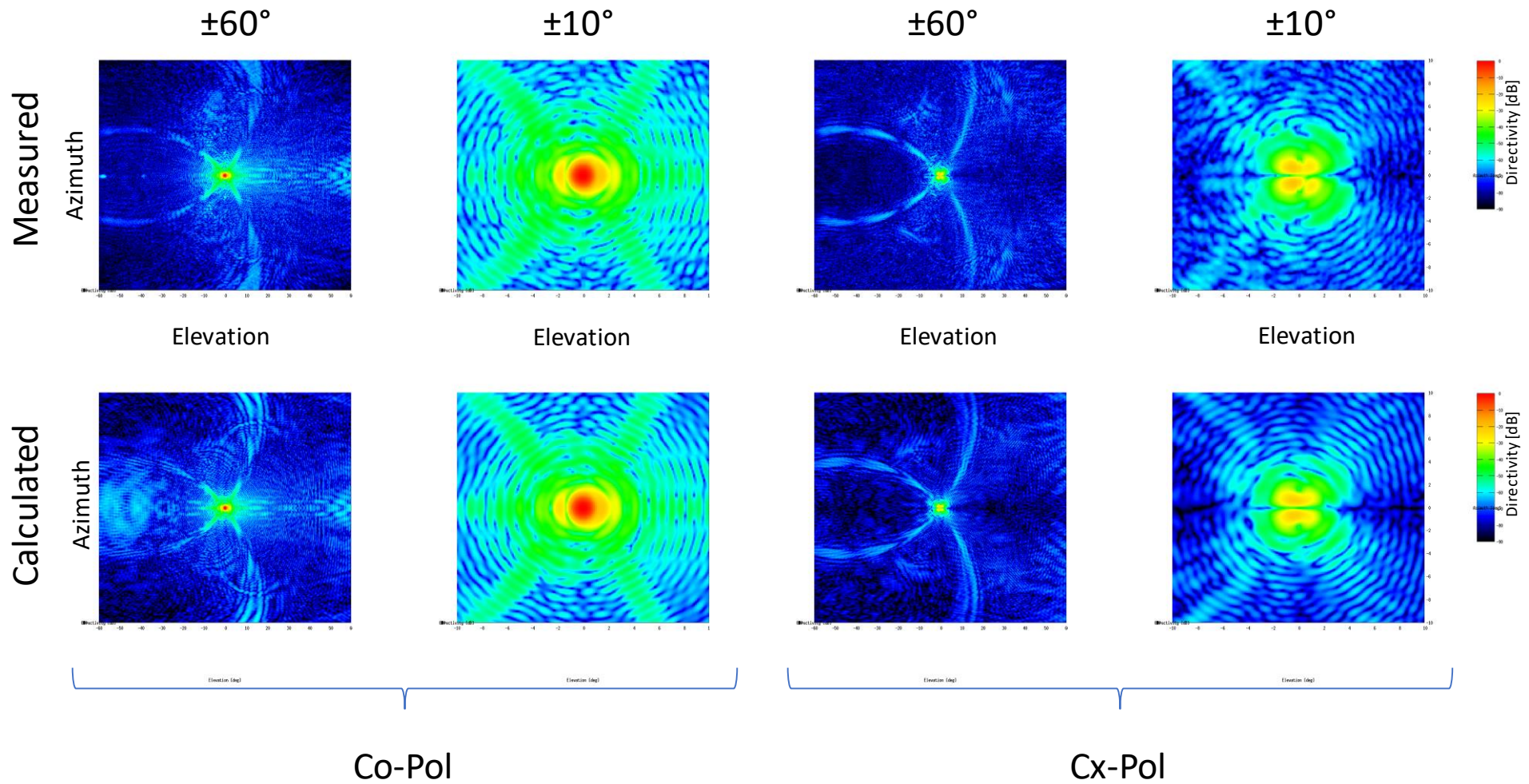
For reference: green = -45dB, light blue = -60dB



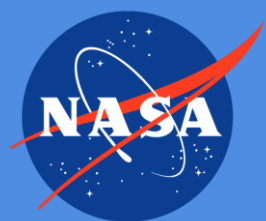
H-Pol Results, 180°, 33.9 GHz

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For reference: green = -45dB, light blue = -60dB

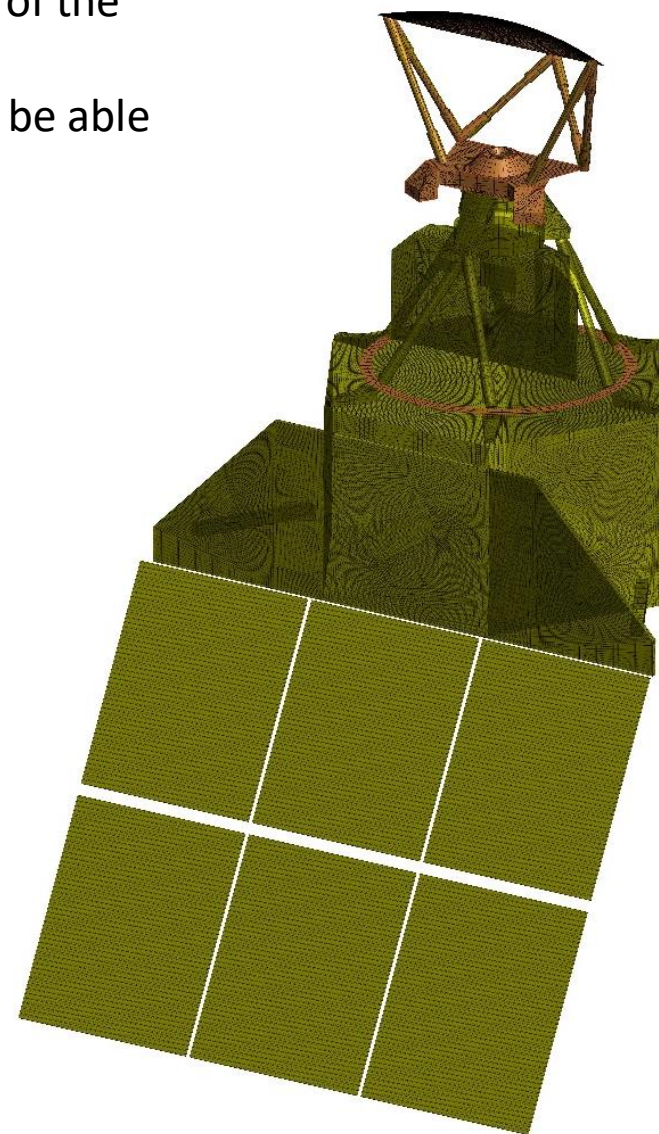
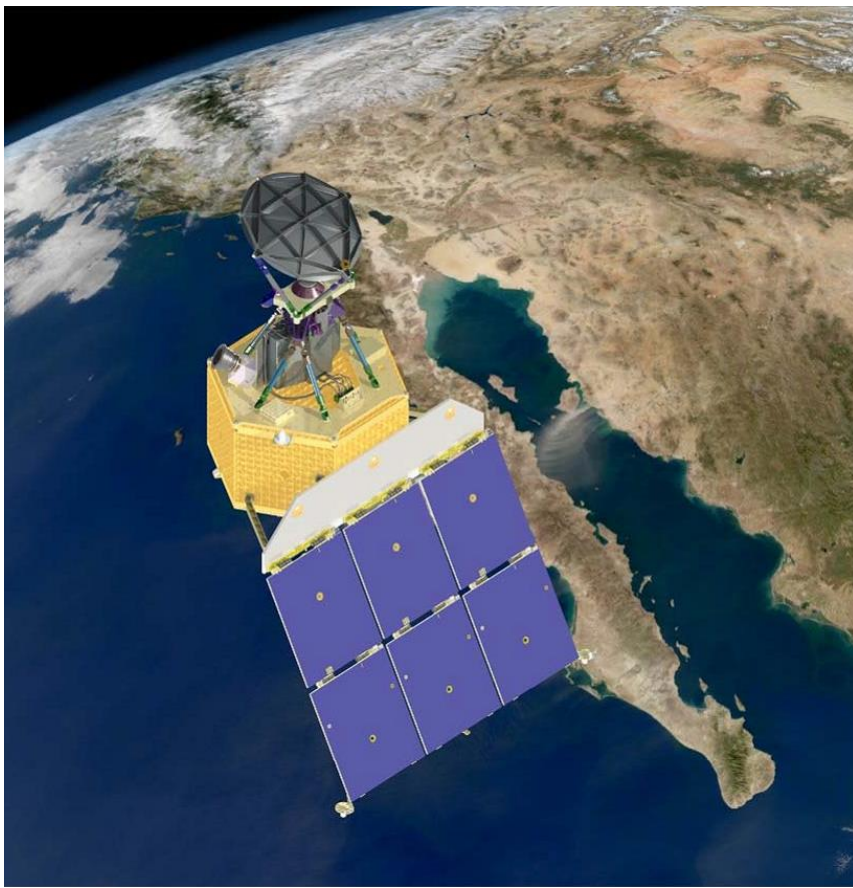


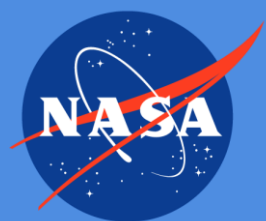
Improved RF Model

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- A larger RF model was also made in GRASP to test the effect of the solar array on the instrument's performance.
- TICRA's help was instrumental in optimizing the RF model to be able to run it on a single machine.

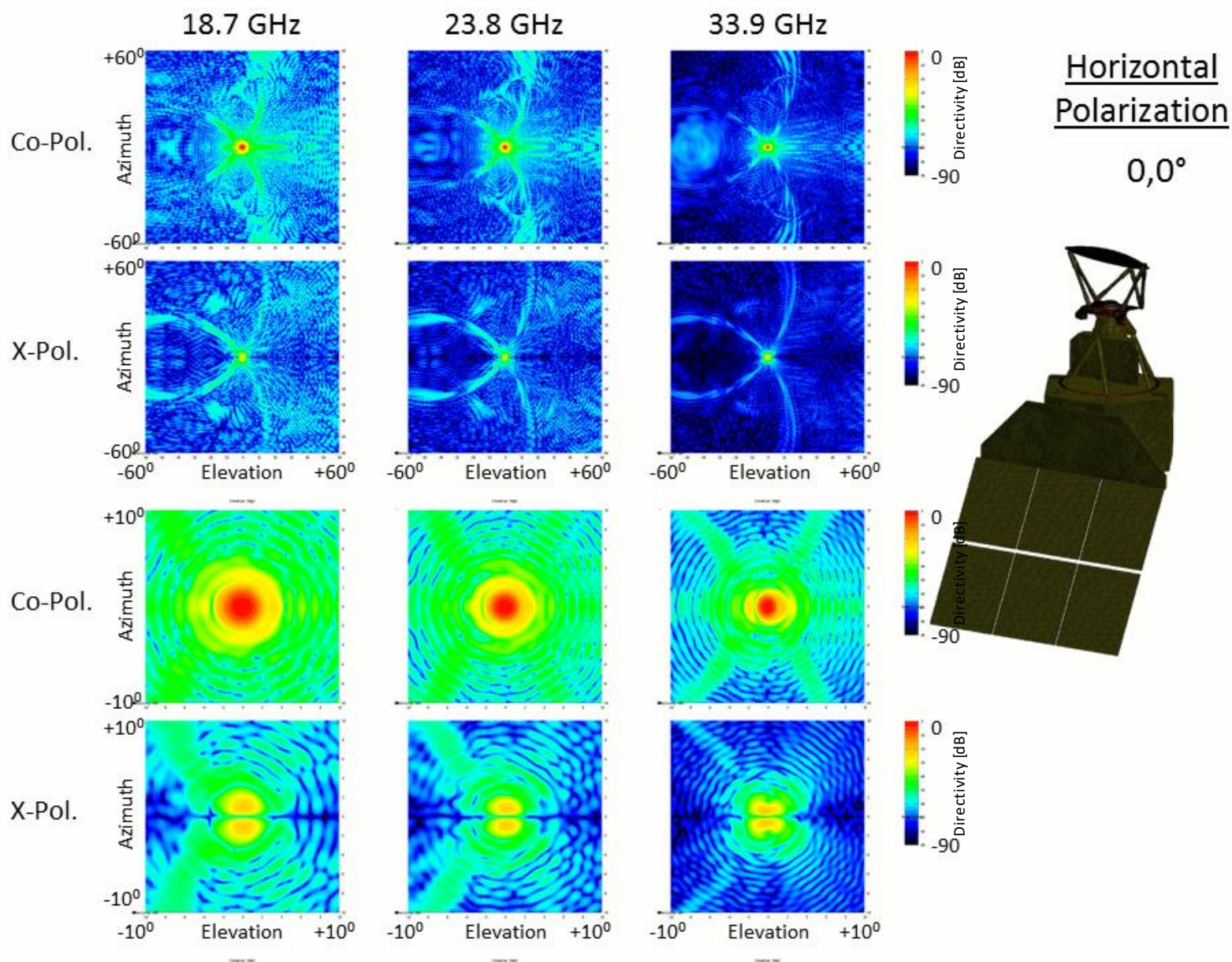


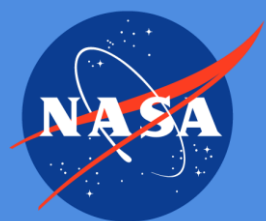


H-Pol Data

This document has

Government sponsorship acknowledged.



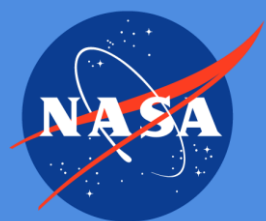


Deployment on ISS – 1

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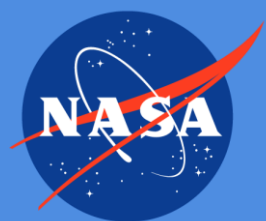


Deployment on ISS – 2

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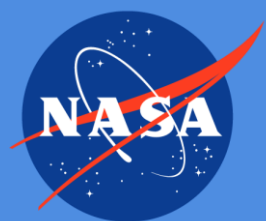


Deployment on ISS – 3

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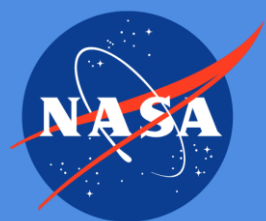


Deployment on ISS – 4

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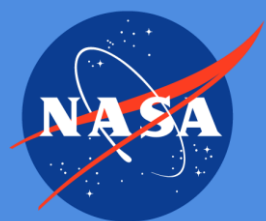


Deployment on ISS – 5

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NISAR, NASA ISRO SAR

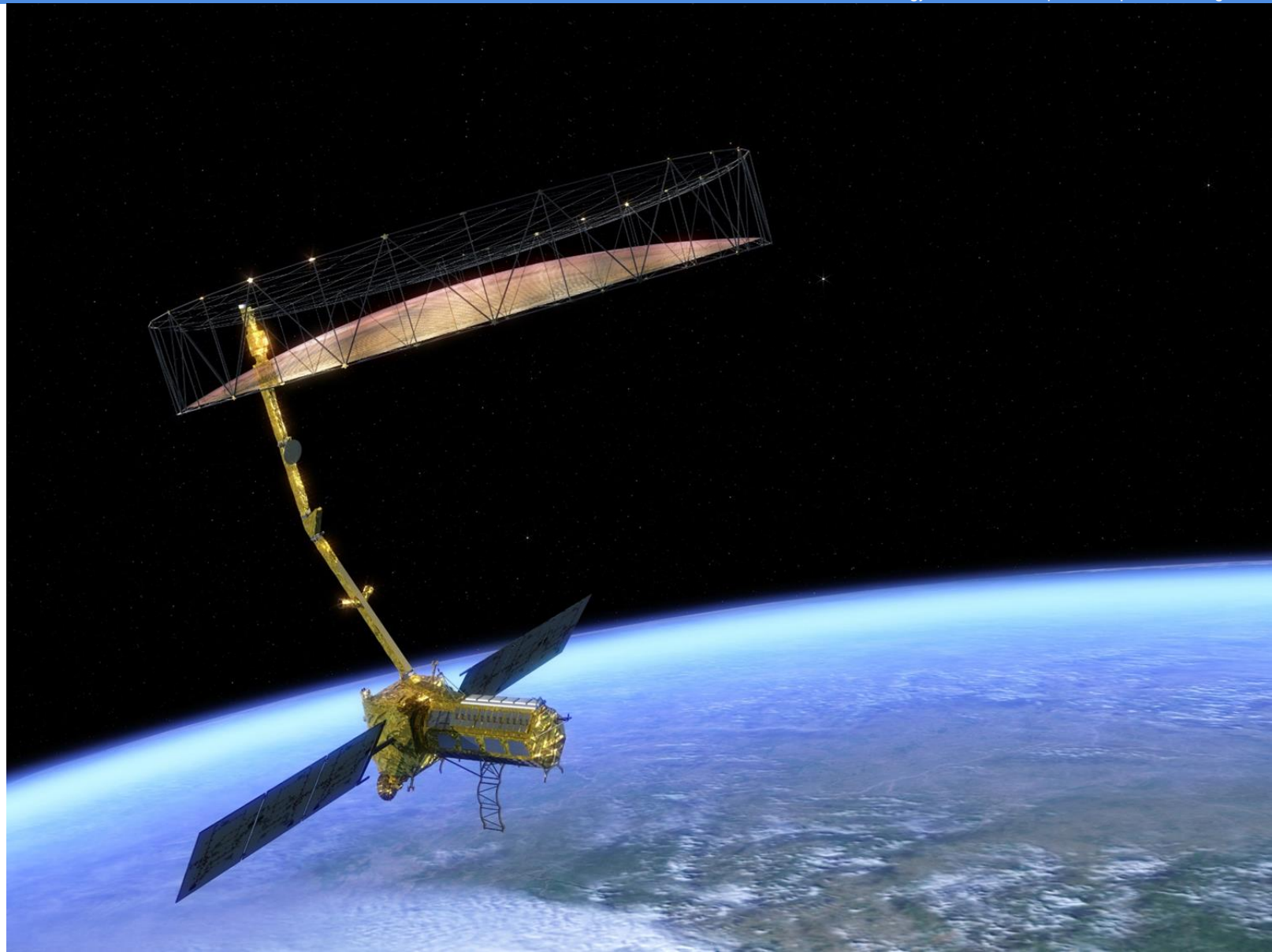
This document has been reviewed and determined not to contain export controlled CUI.

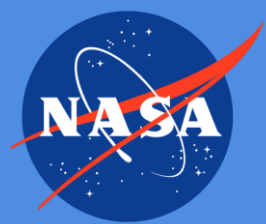
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- NISAR is a collaboration between:



- NISAR's mission is to measure how Earth's environment changes over time.
- It will be launched in early 2024 from India.
- The instrument consists of:
 - 12-m deployable mesh offset reflector
 - An L-band 2x12 element patch feed array (JPL)
 - An S-band 2x24 element patch feed array (SAC/ISRO)
 - An L-band radar (JPL)
 - An S-band radar (SAC/ISRO)
- The observatory can be operated both in left and right looking configurations.
- Sweep-SAR technique.
- 12-day repeat cycle.



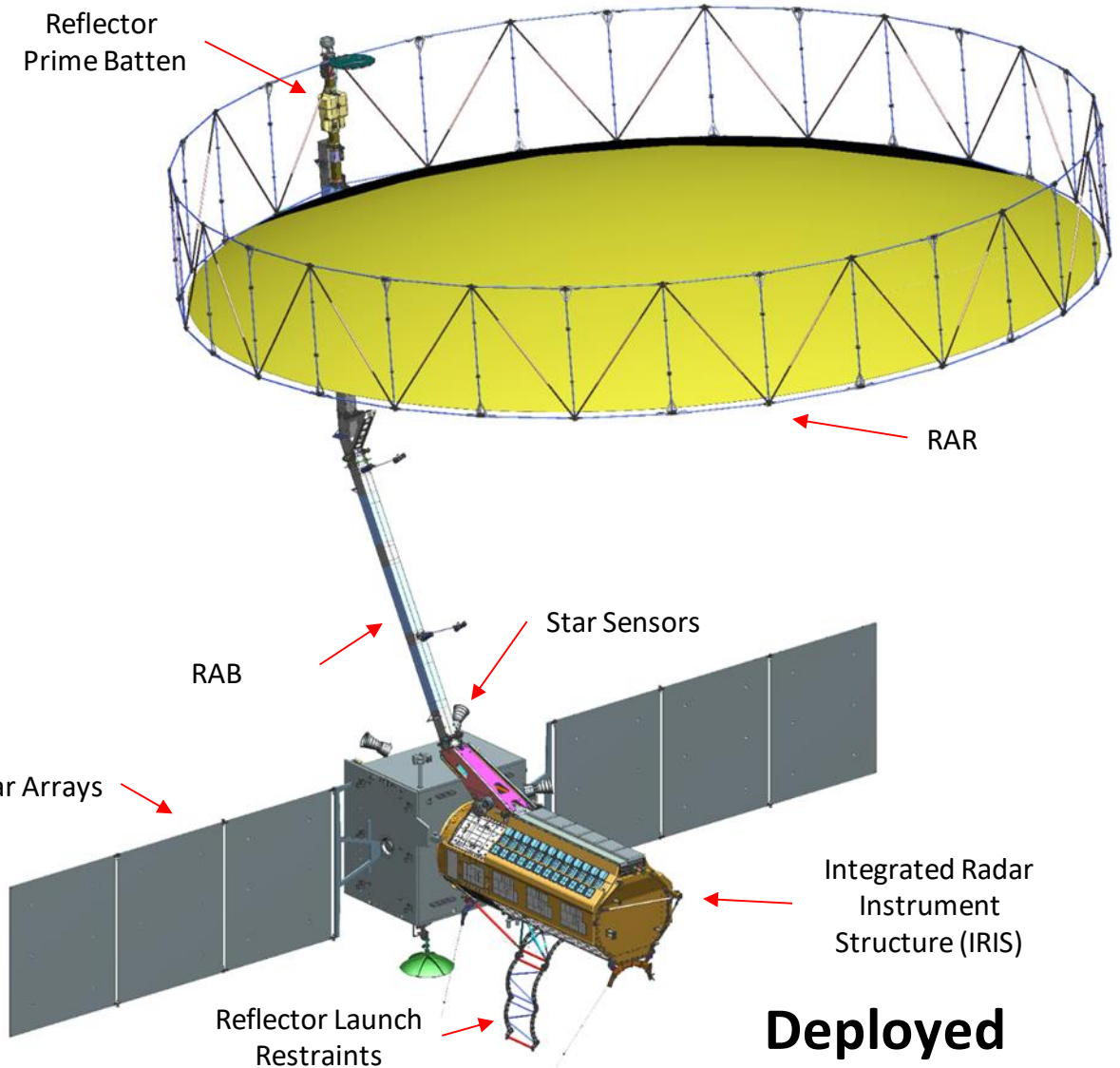
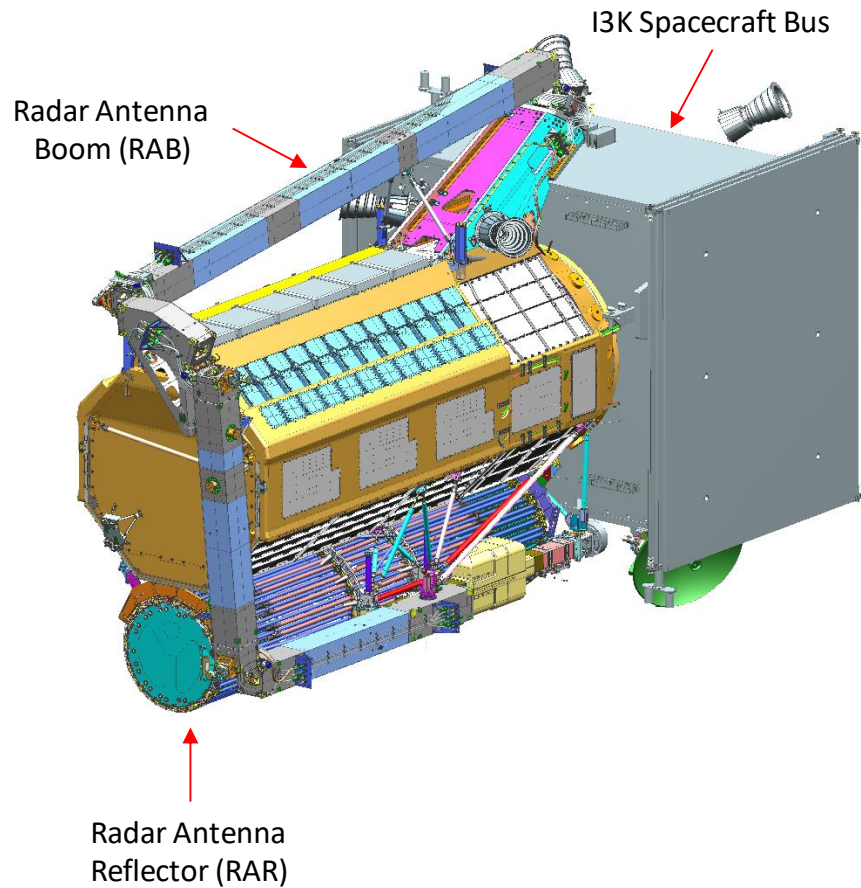


Observatory Configuration

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Stowed



Deployed

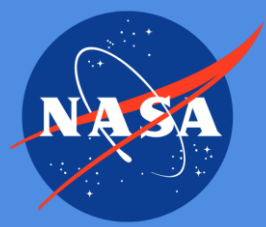


At JPL Before Being Shipped to India

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Sweep-SAR Measurement Technique

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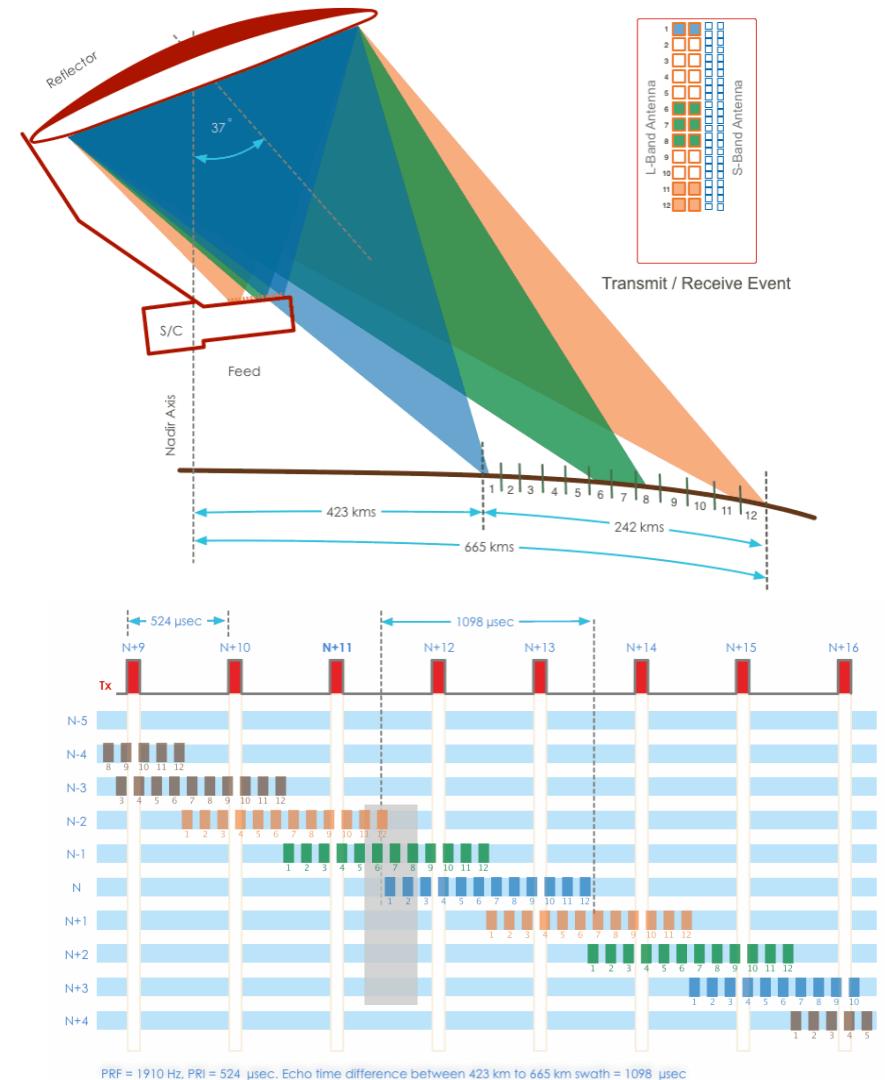
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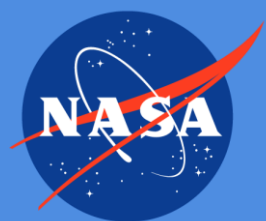
■ Sweep-SAR Basics

- On Transmit, illuminate the entire swath of interest (red beam)
- On Receive, steer the beam in fast time to follow the angle of the echo coming back to maximize the SNR of the signal and reject range ambiguities
- Allows echo to span more than 1 interpulse period

■ Consequences

- 4 echoes can be simultaneously returning to the radar from 4 different angles in 4 different groups of antenna beams
- Each echo needs to be sampled, filtered, beam-formed, further filtered, and compressed
- Onboard processing is not reversible – requires onboard calibration before data is combined to achieve optimum performance



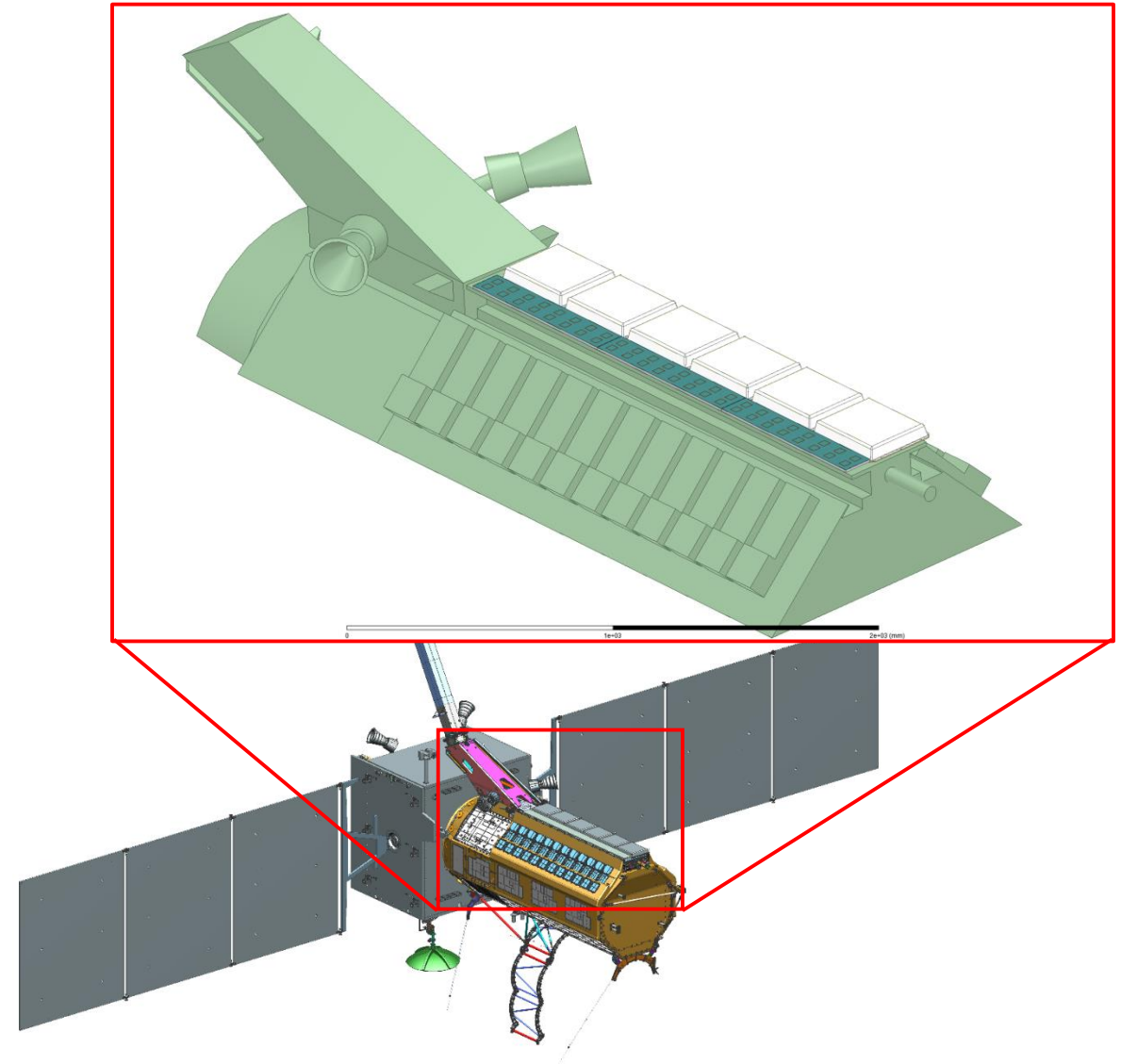


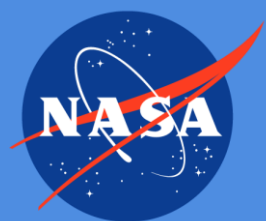
L-Band Feed RF Aperture (L-FRAP) RF Model

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- Latest Ansys HFSS RF model includes:
 - Complete L-FRAP
 - Simplified version of:
 - S-FRAP
 - Radar antenna system
 - Top 3 panels of IRIS
 - Boom base
 - Star sensors
- This RF model is used to generate radiation patterns to feed the TICRA GRASP analysis that includes the entire spacecraft.
- Each LFTA is 358 x 310 mm.
- L-FRAP is 2,158 x 310 mm.



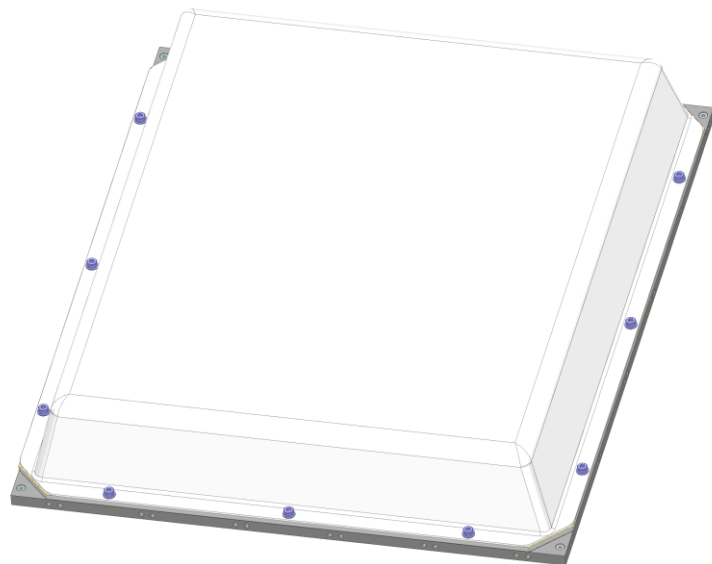


L-Band Feed Tile Assembly (LFTA) RF Model

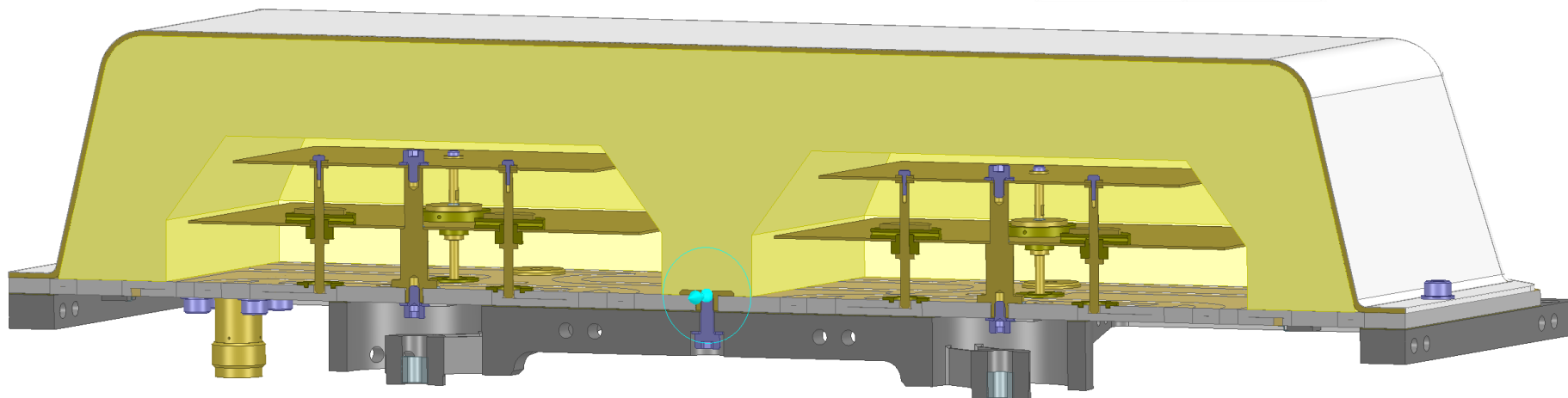
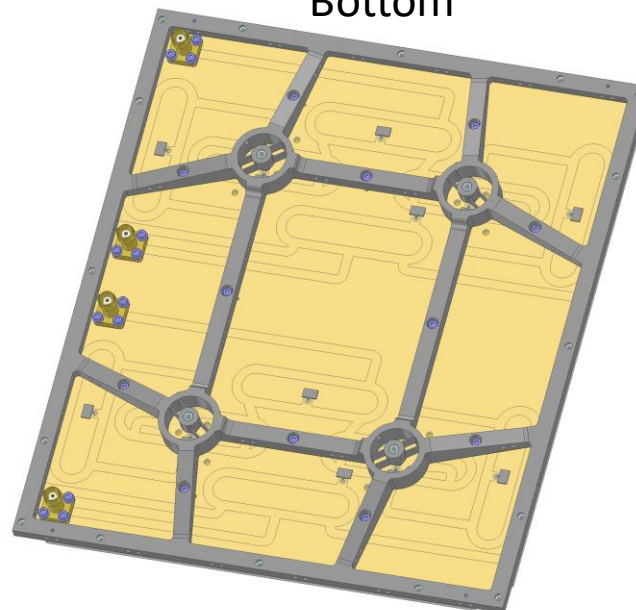
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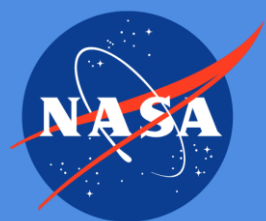
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Top



Bottom



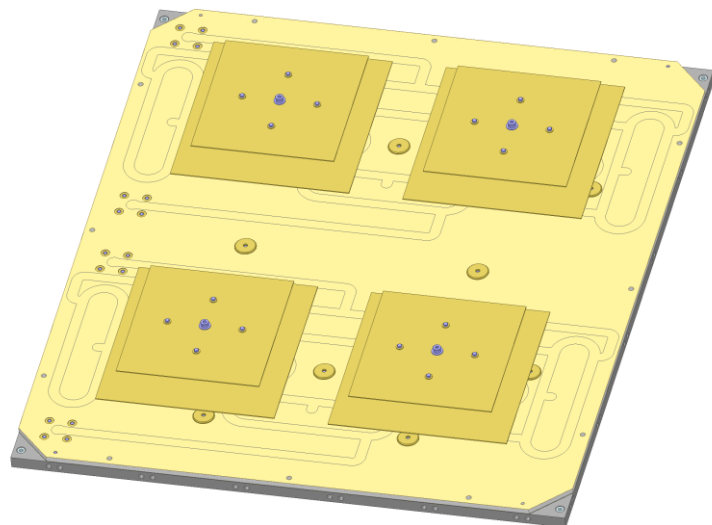


L-Band Feed Tile Assembly (LFTA) RF Model

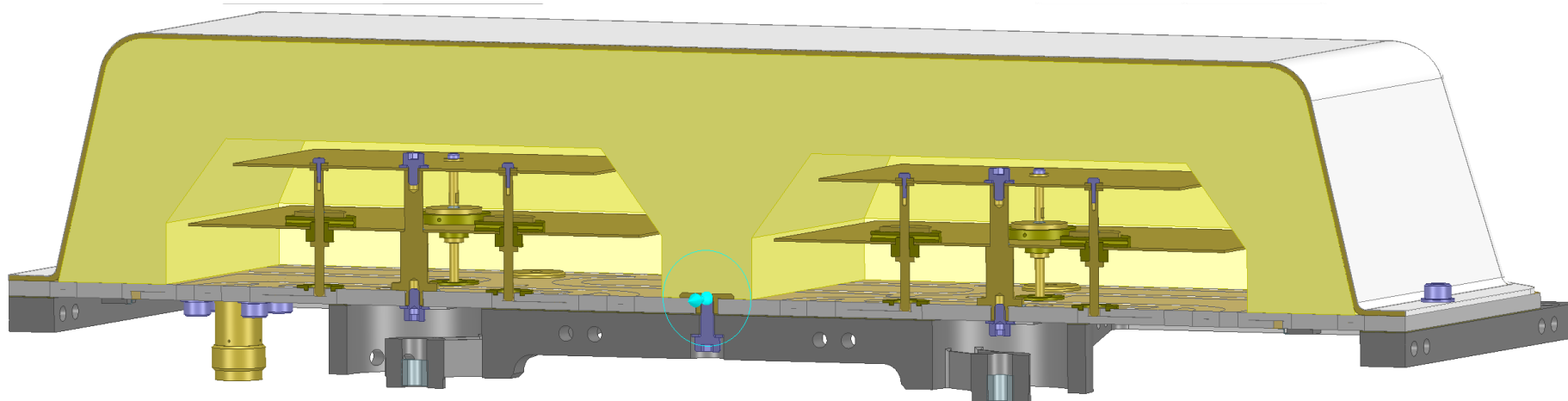
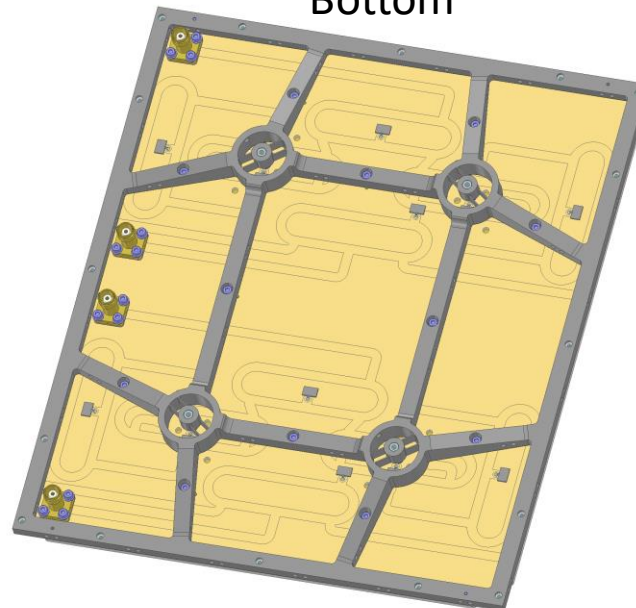
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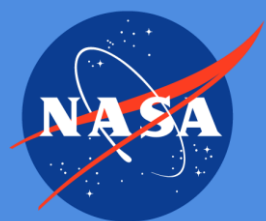
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Top



Bottom

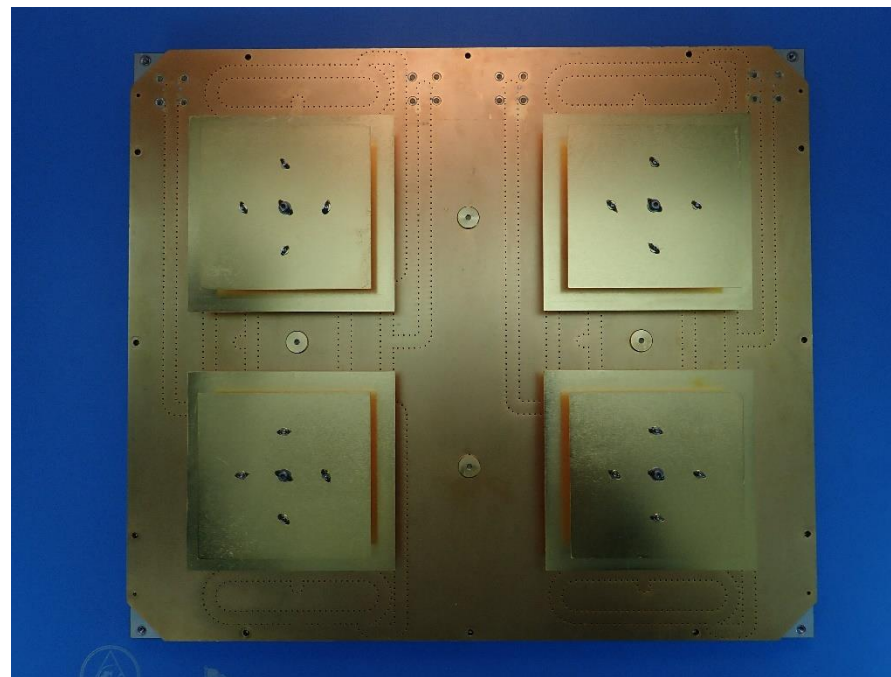
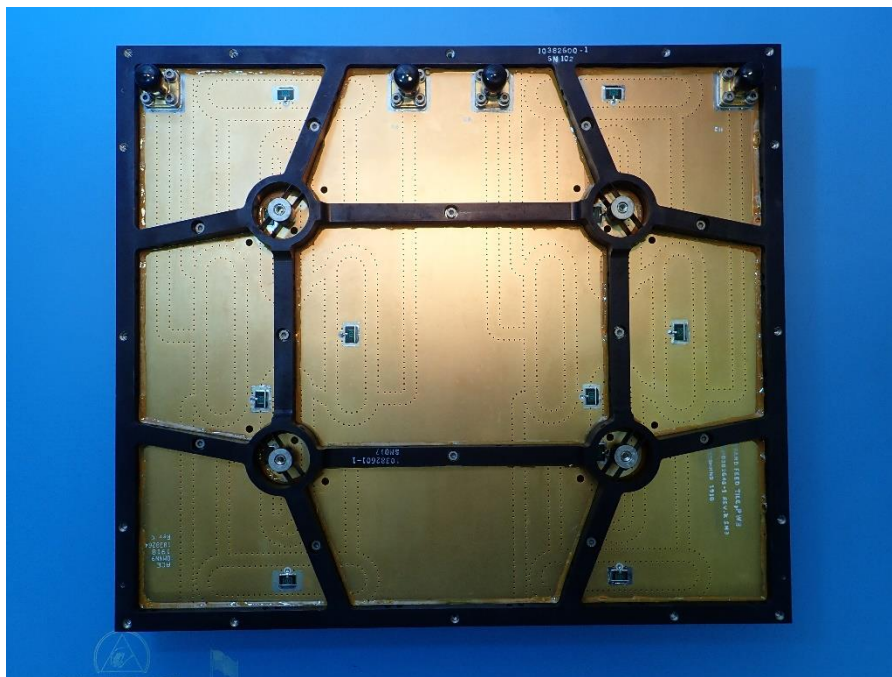


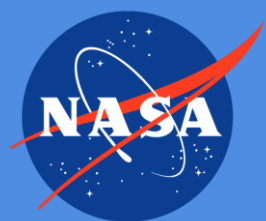


Flight Unit Without Radome, PN 10382600-1

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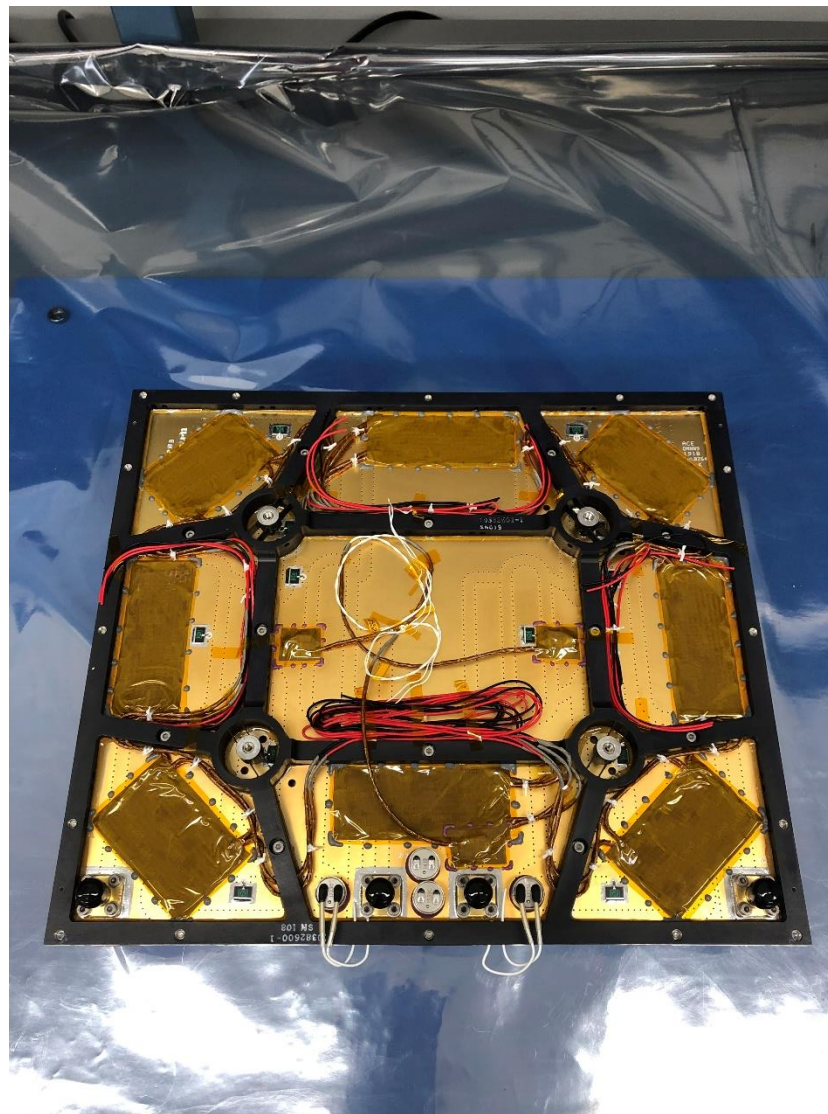


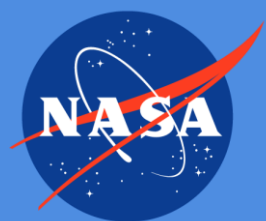


Complete FM LFTA, PN 10380071-1

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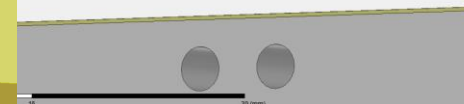
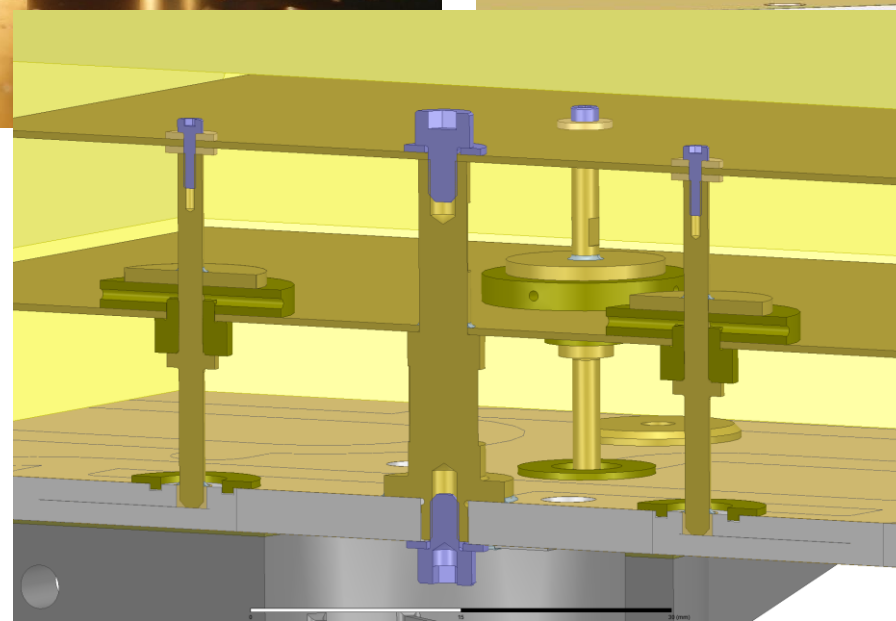
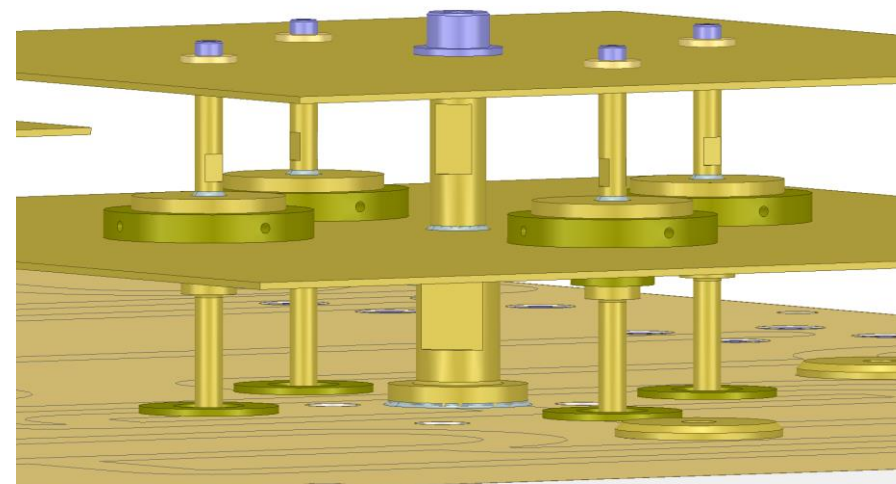
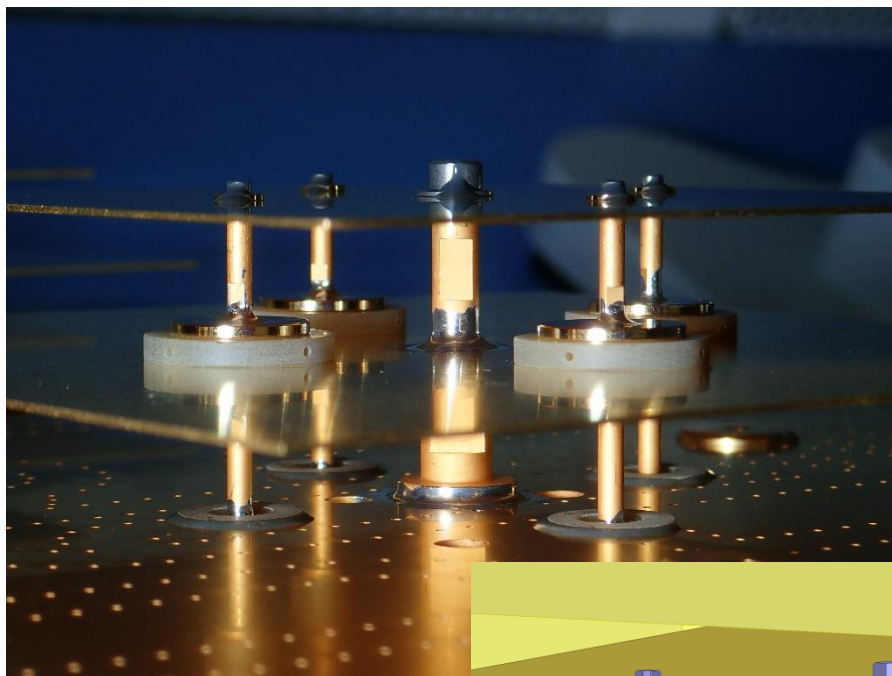


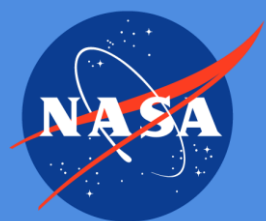


Detail of the Patch Assembly

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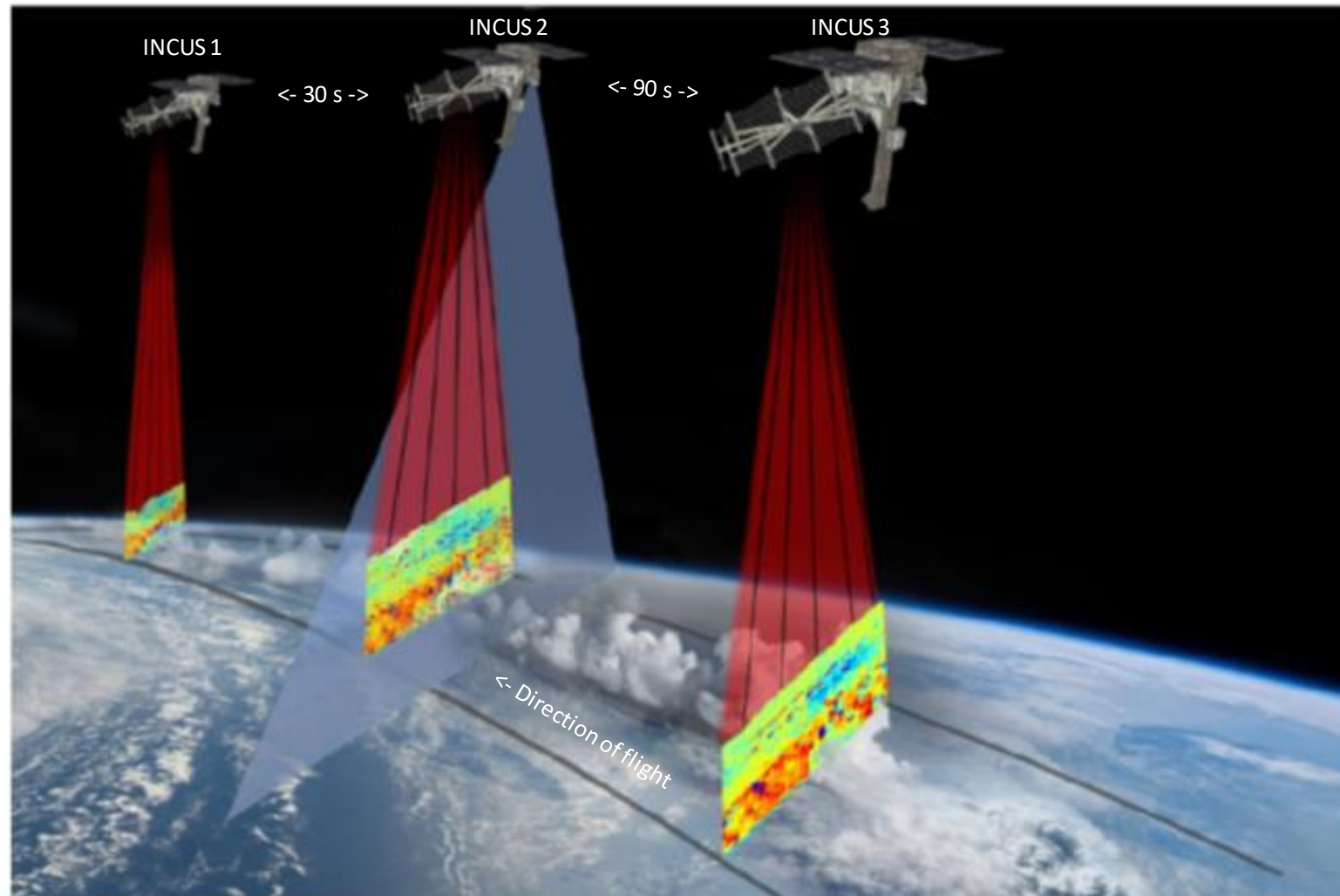


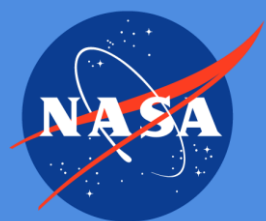
INCUS, Investigation of Convective Updrafts

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- Low Earth Orbit
 - 450-550 km, ~9 km radar swath made of 7 individual 3-km beams
- Three almost identical observatories to measure time-differenced profiles of radar reflectivity
 - Satellites are 30 and 90 seconds apart
- Each observatory has:
 - 1.6-m deployable mesh KaTENna, provided by Tenedeg
 - F/D = 0.7, 20 cm offset
 - **7-beam Ka-band feed assembly, developed at JPL**
 - Ka-band radar (dynamic atmospheric radar), heritage of Rain-Cube
- Middle observatory has a dynamic microwave radiometer with 4 channels between 150 GHz and 190 GHz, heritage of TEMPEST-D
 - 500-km swath
 - The other two observatories have mass mockups
- Commercial satellite bus provided by Blue Canyon Technologies





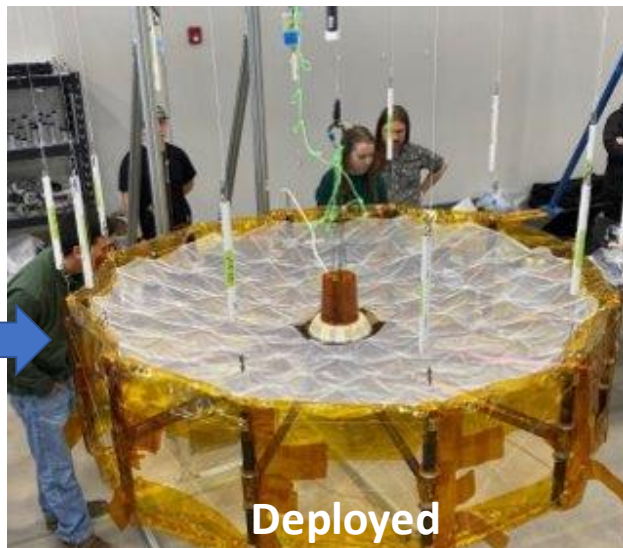
Tendeg KaTENna Design & RF Model

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1.5-m Tendeg KaTENna Design, Stowed



Deployed

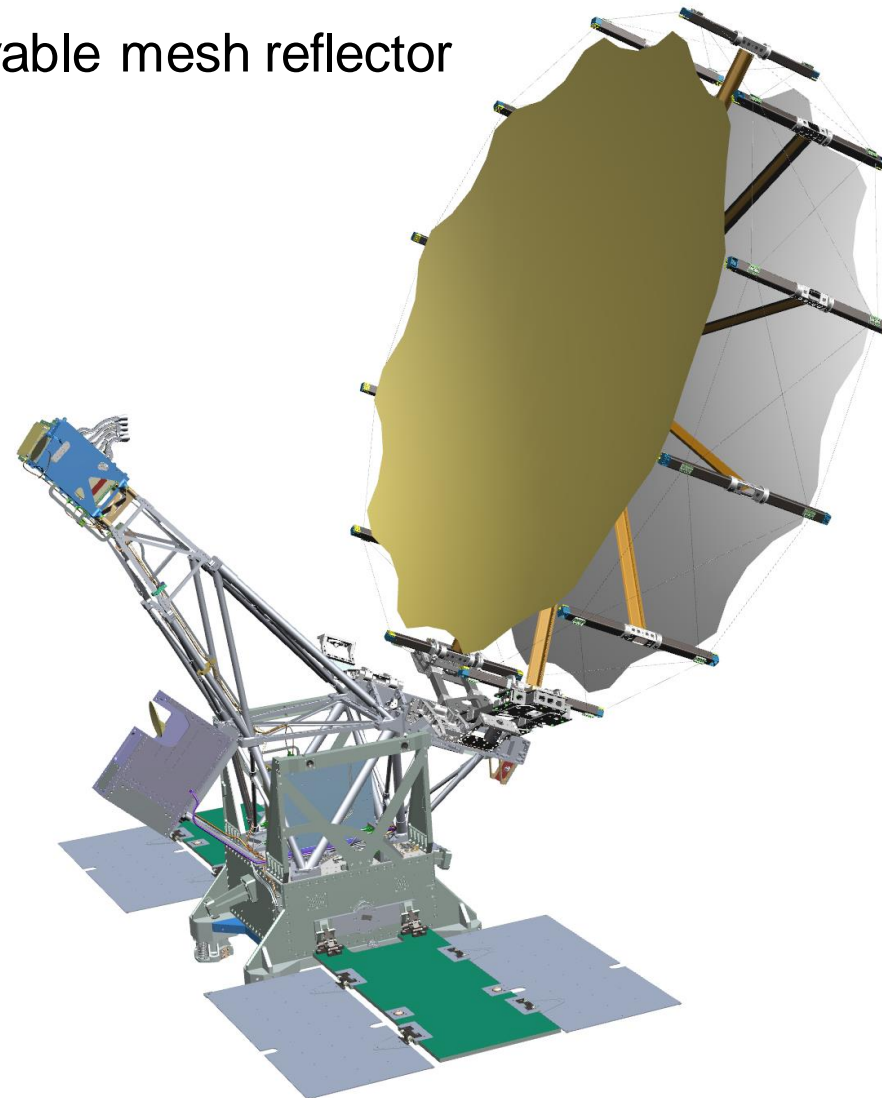
- Tendeg KaTENna, Tendeg proprietary Ka-band design
- 1.6-m deployable mesh reflector
- $F/D = 0.7$
- 20-cm offset

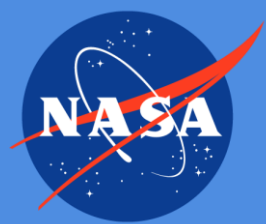
(Source: Courtesy of Tendeg)



3-m Tendeg KaTENna Design

(Source: Courtesy of Tendeg)





Golden Devices' Feed Assembly Prototypes

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SLA or SLS Polyimide models



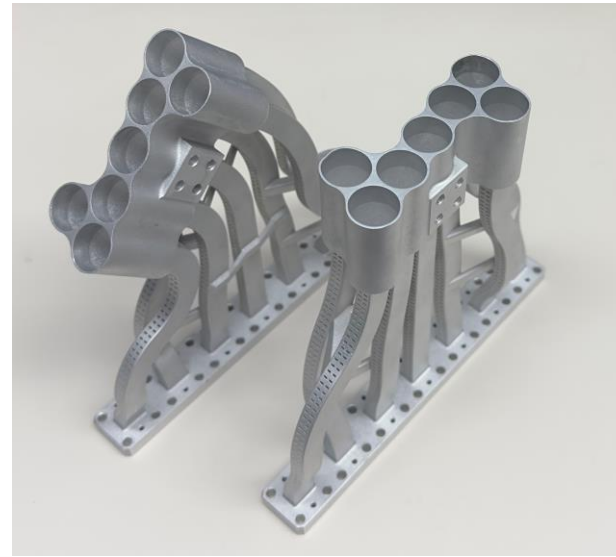
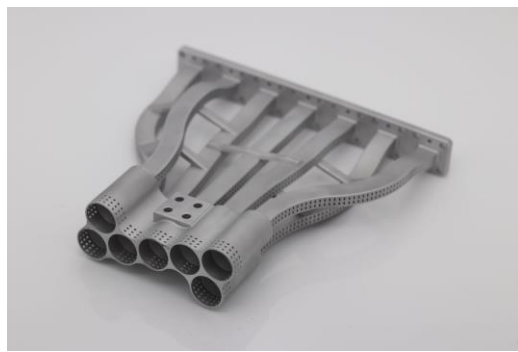
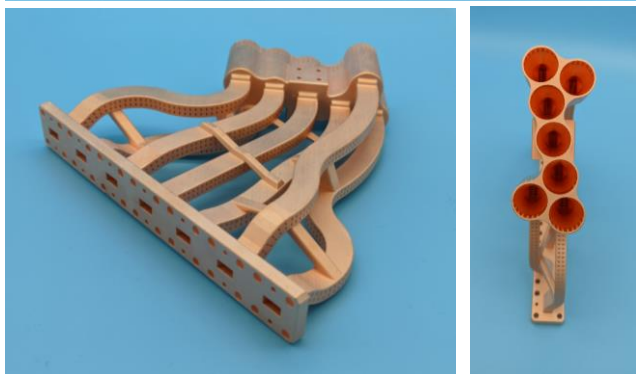
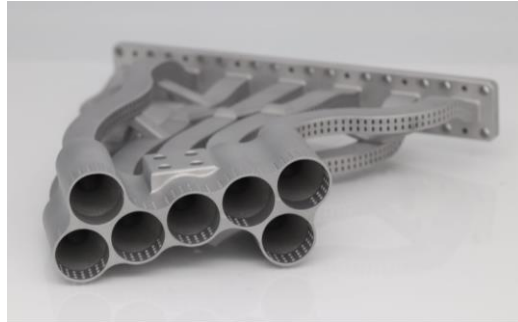
Cast Al models based on SLA mold



New Cast Al models

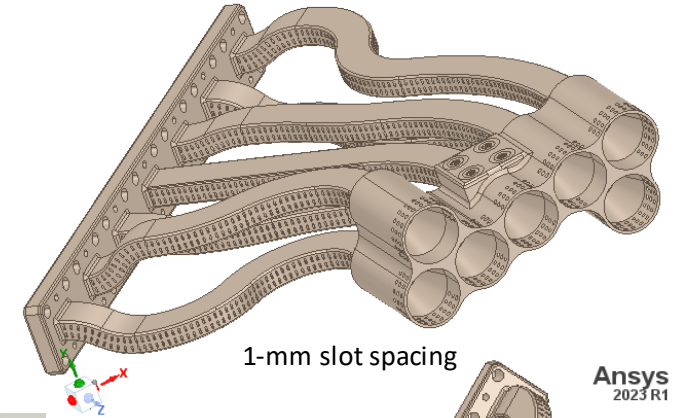


Cu or Cu+Ni plated SLS and SLA models



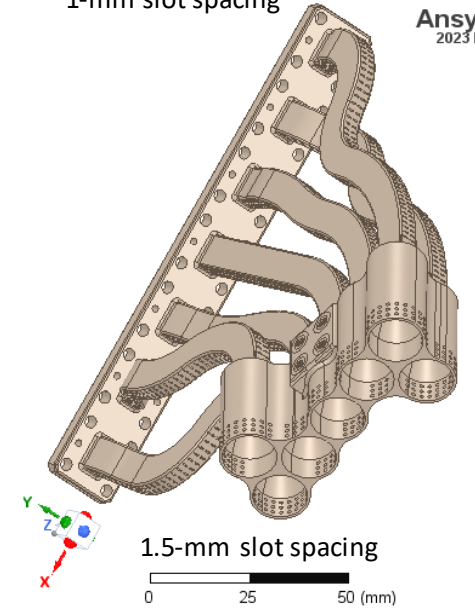
HFSS Models

Ansys
2023 R1



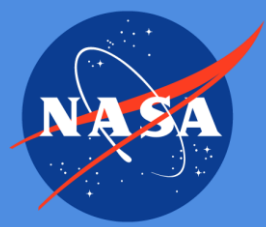
1-mm slot spacing

Ansys
2023 R1



1.5-mm slot spacing

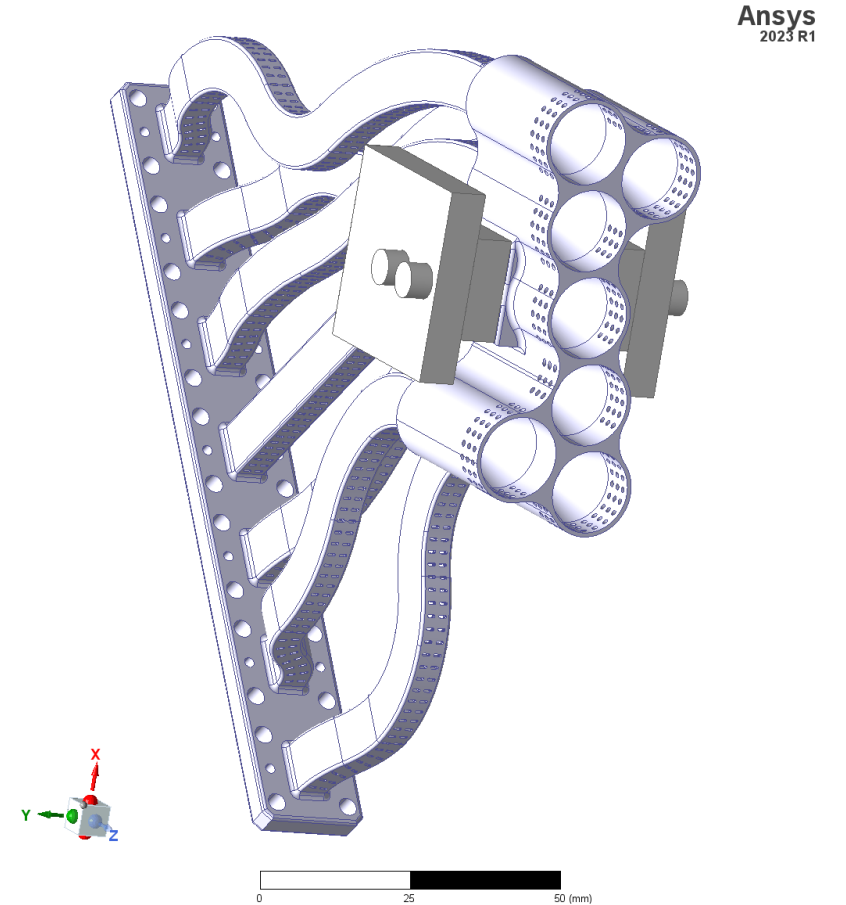
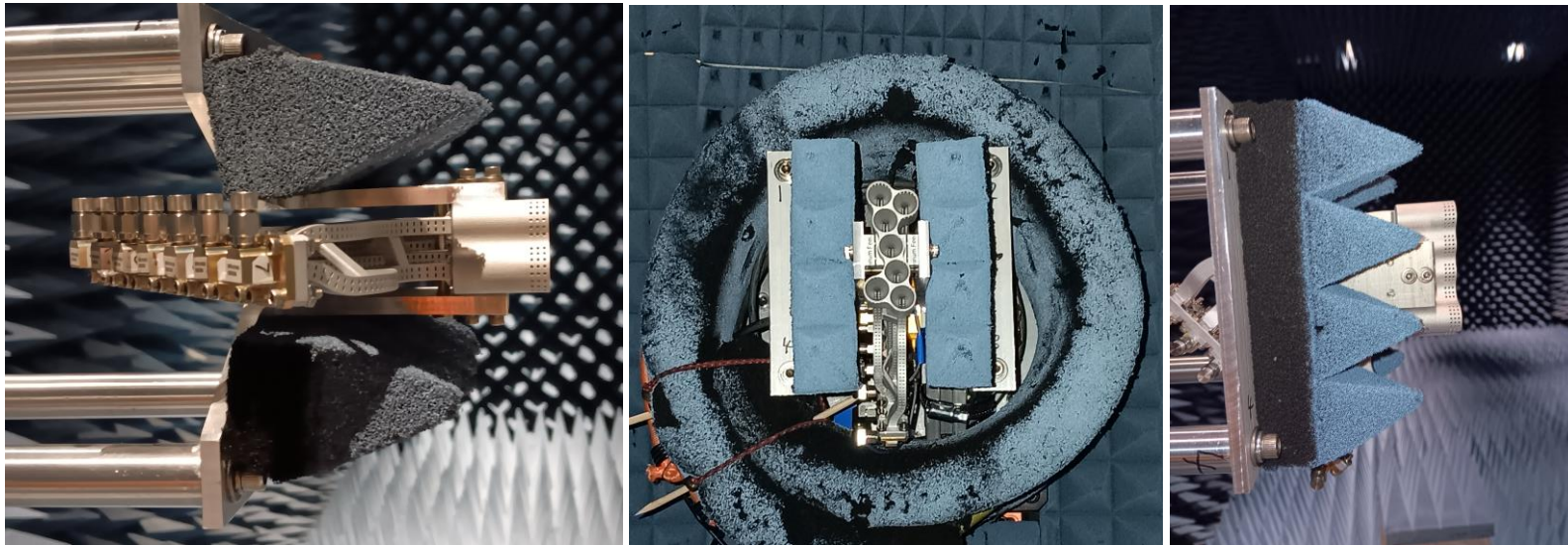
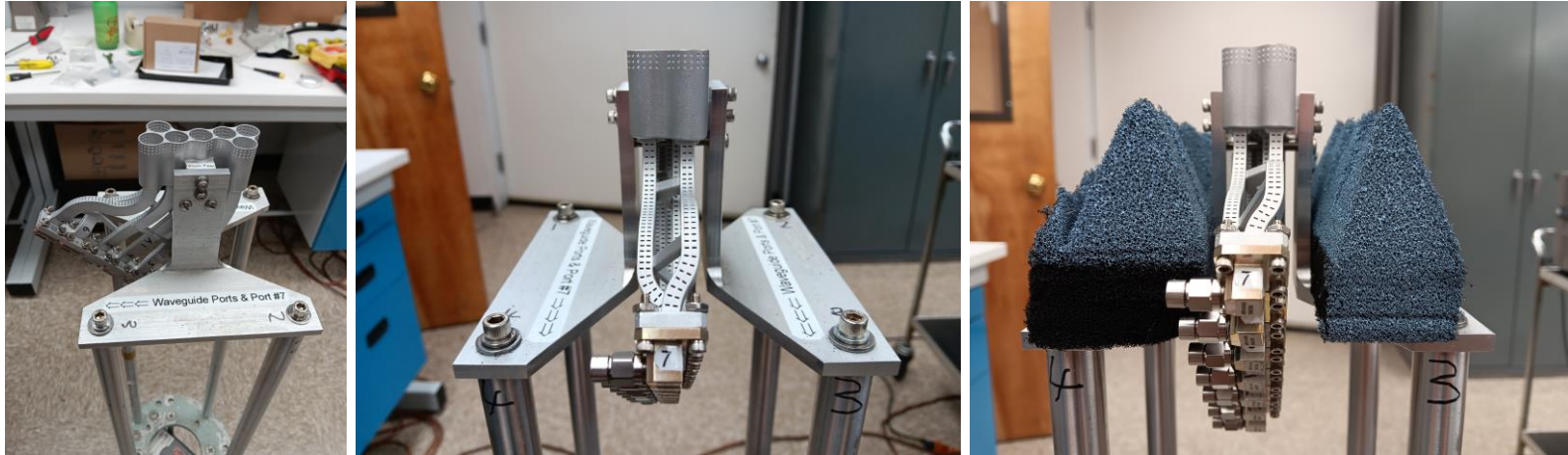
0 25 50 (mm)



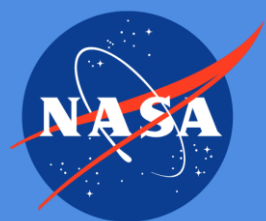
Radiation Pattern Measurements

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HFSS model used to compare with measurements

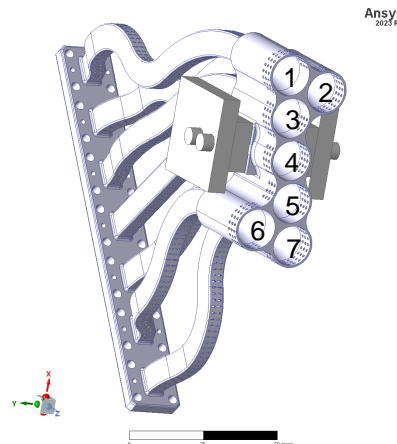


Port 4 Patterns

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Reflector subtended angle is 34° .

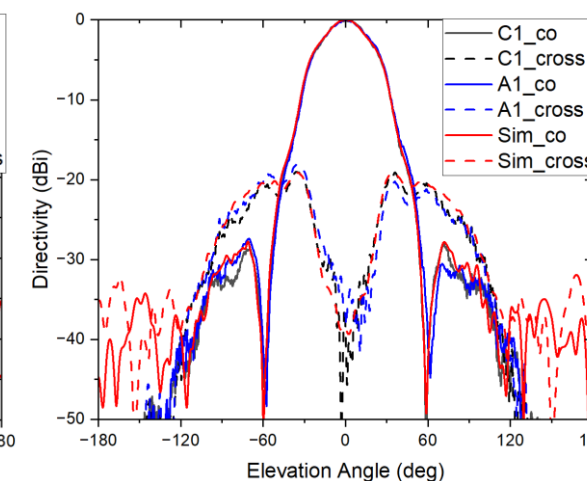
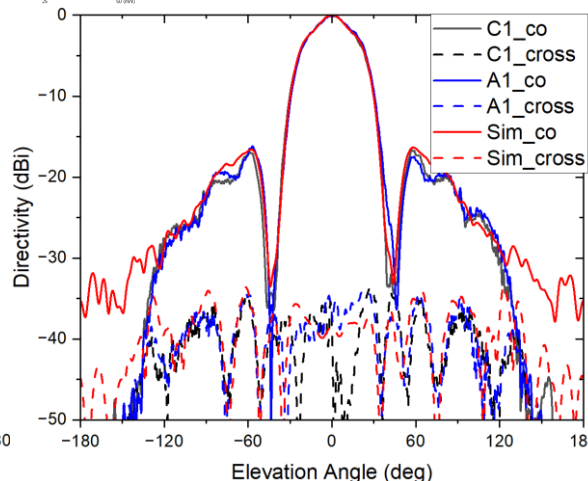
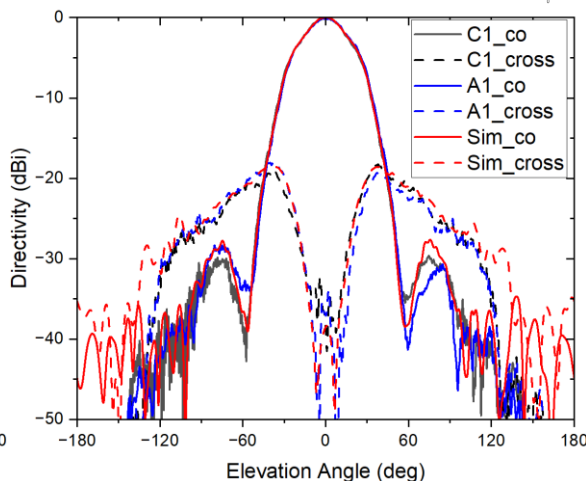
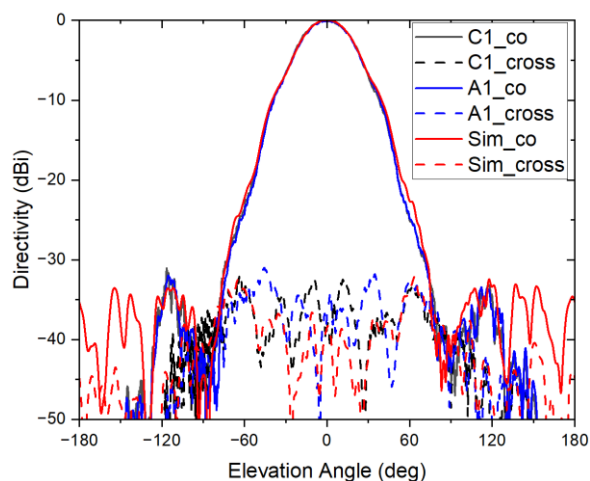


0° cut

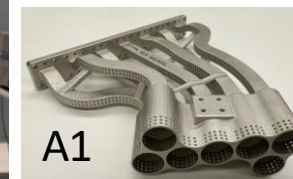
45° cut

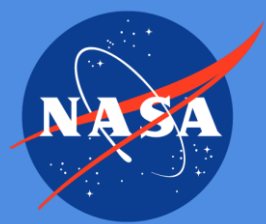
90° cut

135° cut



- All measurements were corrected for laser metrology.
- The scattering of the exposed supporting flanges is most visible on the results of this port, which is the center horn.
- The RF model predicted this effect with very good accuracy.

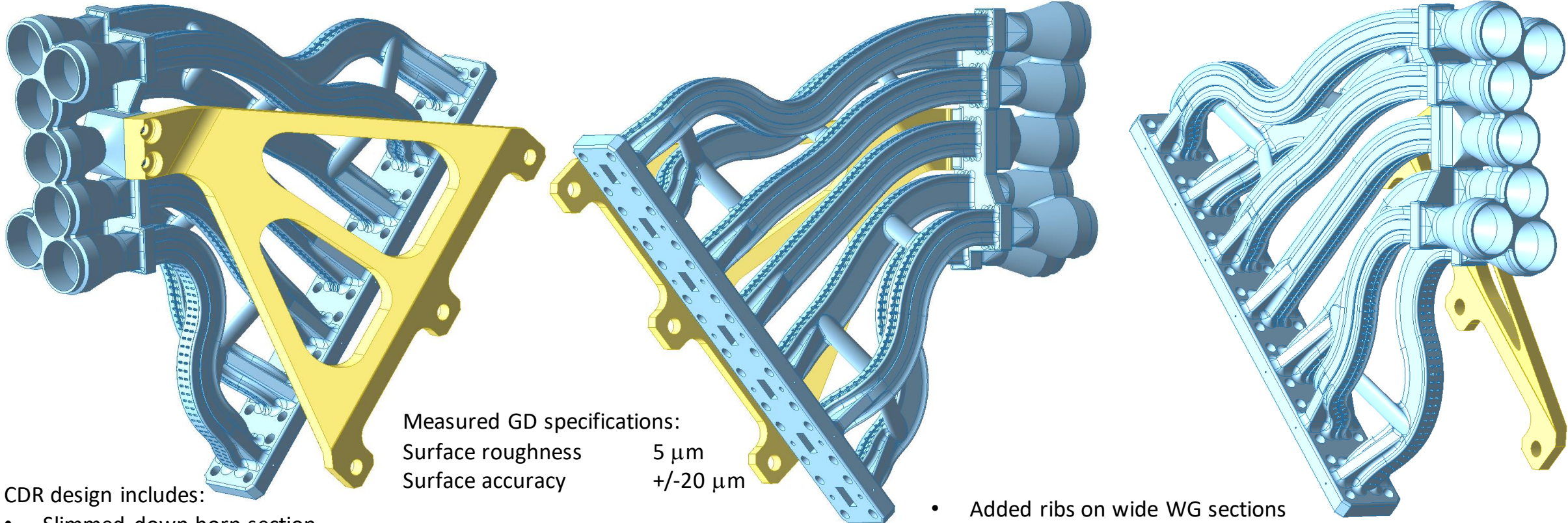




CDR Feed Design

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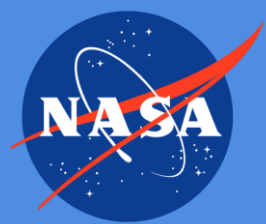
Measured GD specifications:

Surface roughness	5 μm
Surface accuracy	+/-20 μm

CDR design includes:

- Slimmed down horn section
- Single-sided bracket
- Larger fillets
- Better structural support at the FESA MB interface
- Larger bracing between WG

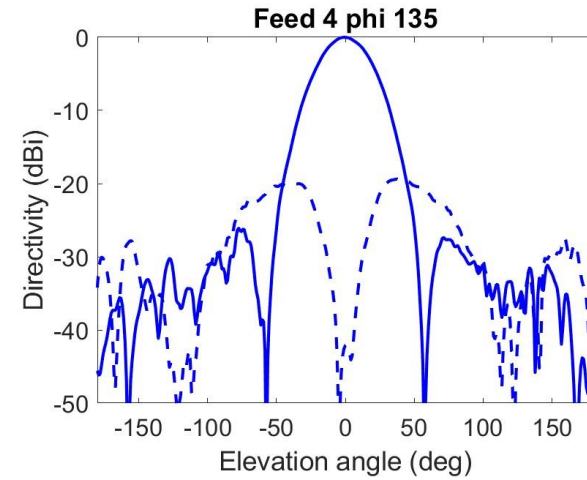
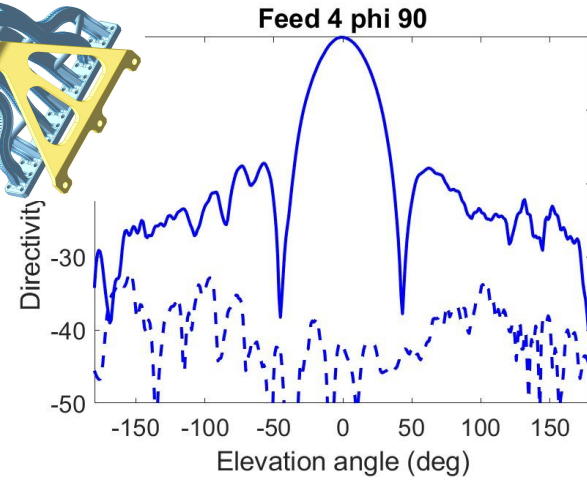
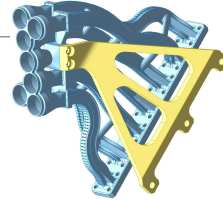
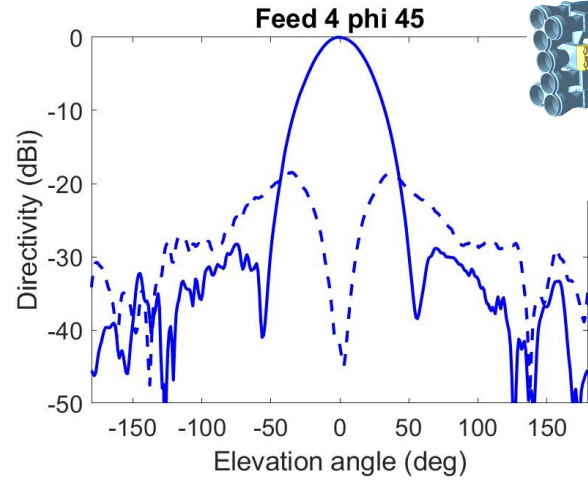
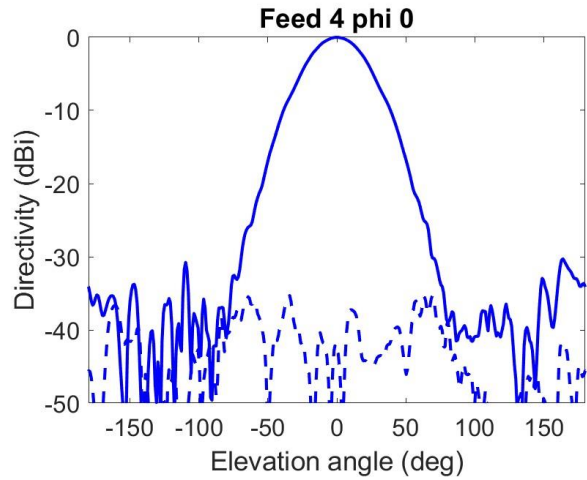
- Added ribs on wide WG sections
- Added ribs along the slotted WG sides
- Venting holes for alignment pins
- Simpler plate behind the horns
- Thicker side on WG 7



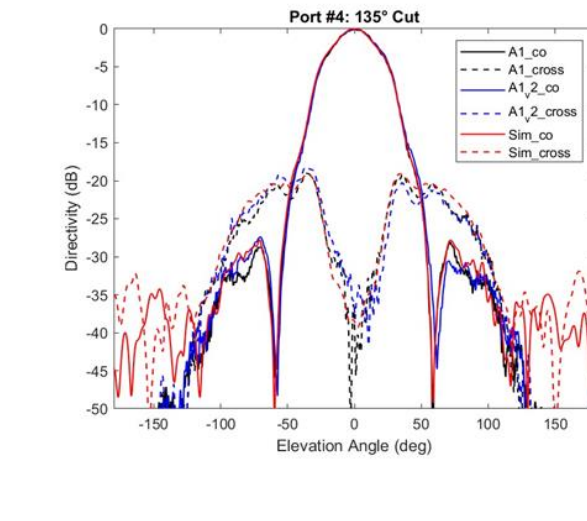
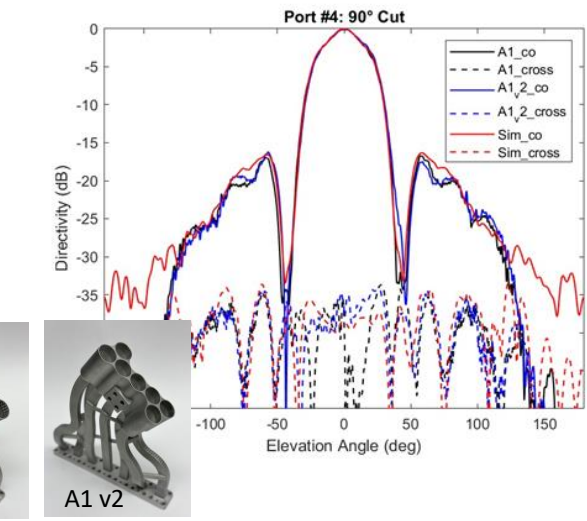
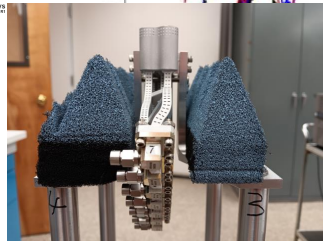
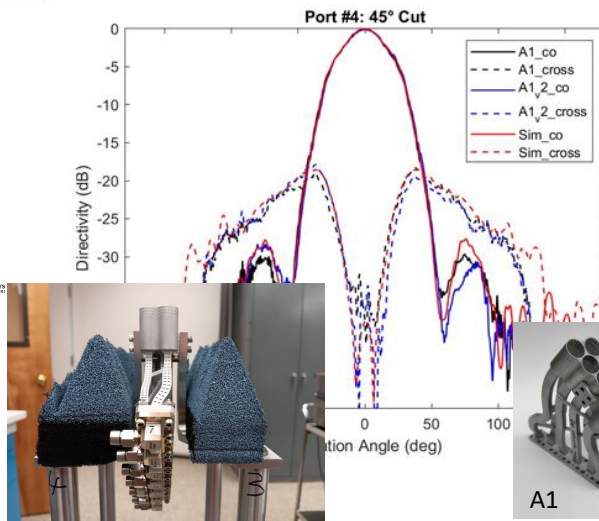
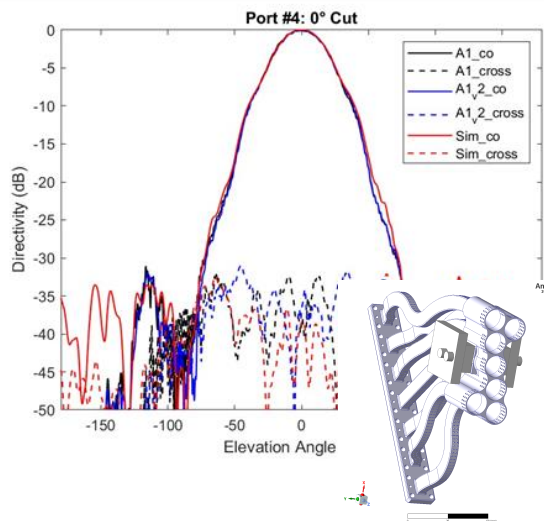
Feed 4

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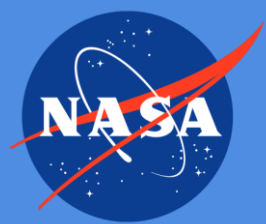
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CDR Design, Calc.



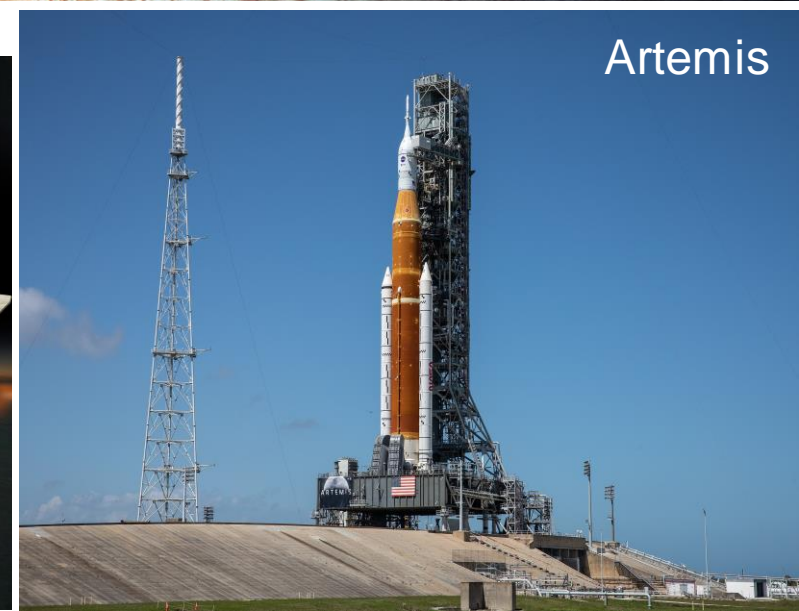
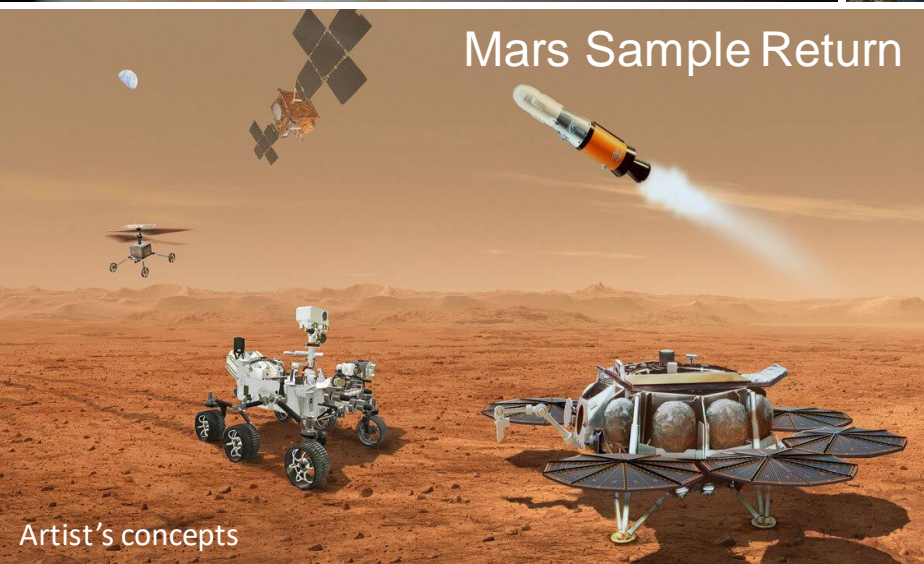
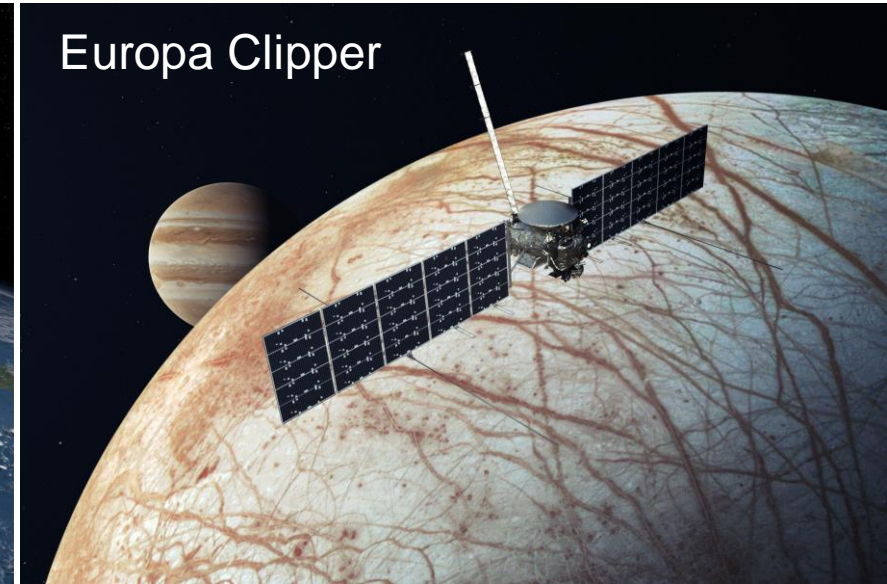
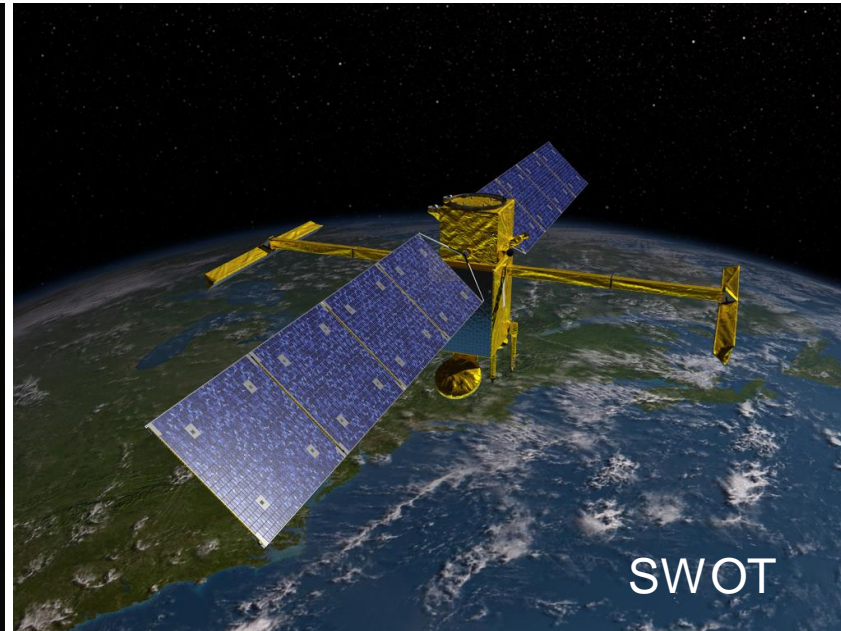
PDR Prototypes, Calc. & Meas.



Other Important Flight Projects

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A night view of the JPL campus with illuminated buildings and a sunset sky. The sky is a mix of deep blue and orange, with wispy clouds. The buildings are lit up, and the foreground is dark and somewhat desolate.

Thank you!

Paolo.Focardi@jpl.nasa.gov